10031 Vitrophyre Basalt

2.7 grams

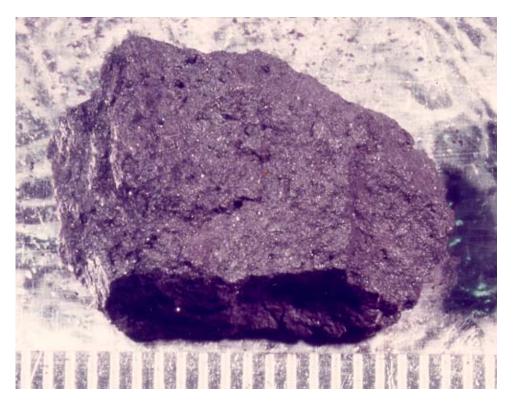


Figure 1: Photo of 10031. Scale 1.8 cm. NASA S76-21144.

Introduction

10031 was returned as a "contingency sample". It is apparently a quickly cooled (quenched) example of the high-K ilmenite basalts from Apollo 11.

It is 3.6 b.y. old and has had an exposure to cosmic rays for ~ 300 m.y.

Petrography

Beaty et al. (1979) describe 10031 as a "vitrophyre". Phenocrysts of olivine and armalcolite (with ilmenite overgrowth) are contained in a fine-grained to glassy groundmass. "10031 contains about 60% crystals which consist of equant armalcolite mantled by rutile-bearing ilmenite, platey ilmenite (50 x 5 microns) with feathery edges, pyroxene (to 100 microns) and minute troilite spheres. These crystals are enclosed in a brown glass which contains numerous incipient crystallites of ilmenite and pyroxene." (from Beaty et al.)

Grove and Beaty (1980) were able to reproduce the texture of 10031 experimentally and determine the cooling rate.

Mineralogy

Olivine: Olivine is Fo₇₅.

Pyroxene: Beaty et al. give the composition of some pyroxene grains.

Armalcolite: The composition of armalcolite in 10031 is discussed in Beaty et al.

Ilmenite: Ilmenite in 10031 forms elongate platy grains (figure 2) and also forms overgrowths on armalcolite grains and has exsolution of rutile.

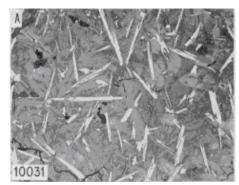


Figure 2: Texture of 10031 from Beaty et al. 1979.

Chemistry

Lunar sample 10031 has the composition typical of the high-K suite of Apollo 11 basalts (table 1).

Radiogenic age dating

Guggisberg et al. (1979) determined the age of 10031 by the Ar/Ar plateau technique (figure 3).

Cosmogenic isotopes and exposure ages

Guggisberg et al. (1979) determined an ${}^{37}Ar/{}^{38}Ar$ exposure age of 300 m.y. There are 2 thin sections.

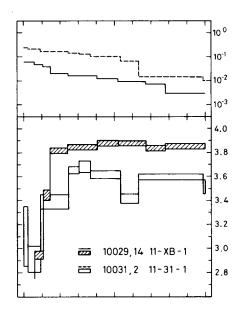


Figure 3: Argon plateau for 10031 compared with 10029 (from Guggisberg et al. 1979).

Summary of Age Data for 10031

Ar/Ar

Guggisberg et al. 1979

 3.6 ± 0.08 b.y. (poorly defined)

Table 1. Chemical composition of 10031.

reference weight	Beaty 1979	
SiO2 % TiO2 Al2O3 FeO MnO MgO CaO Na2O K2O P2O5 S % sum	11.4 8 20.7 0.225 8 10.6 0.503 0.3	(a) (a) (a) (a) (a) (a) (a)
Sc ppm V Cr Co Ni Cu Zn Ga Ge ppb As Se Rb Sr	87 65 2370 28	(a) (a) (a) (a)
Y Zr Nb Mo Ru Rh Pd ppb Ag ppb Cd ppb In ppb Sn ppb Sb ppb Te ppb	370	(a)
Cs ppm Ba La Ce	330 27.3 78	(a) (a) (a)
Pr Nd Sm Eu	64 20.7 2.23	(a) (a) (a)
Gd Tb Dy Ho Er	4.3 30	(a) (a)
Tm Yb Lu Hf Ta W ppb Re ppb Os ppb Ir ppb Pt ppb	17.2 2.43 15.4 2.5	(a) (a) (a) (a)
Au ppb Th ppm U ppm technique:	2.8 (a) INAA	(a)

References for 10031

Beaty D.W., Hill S.M.R., Albee A.L., Ma M.-S., and Schmitt R.A. (1979a) The petrology and chemistry of basaltic fragments from the Apollo 11 soil, part 1. *Proc.* 10th Lunar Sci. Conf. 41-75.

Grove T.L. and Beaty D.W. (1980) Classification, experimental petrology and possible volcanic histories of the Apollo 11 high-K basalts. *Proc. 11th Lunar Planet. Sci. Conf.* 149-177.

Guggisberg S., Eberhardt P., Geiss J., Grogler N., Stettler A., Brown G.M. and Pecket A. (1979) Classification of the Apollo-11 basalts according to Ar³⁹-Ar⁴⁰ ages and petrological properties. *Proc.* 10th Lunar Planet. Sci. Conf. 1-39.

James O.B. and Jackson E.D. (1970) Petrology of the Apollo 11 ilmenite basalts. *J. Geophys. Res.* **75**, 5793-5824.

Kramer F.E., Twedell D.B. and Walton W.J.A. (1977) **Apollo 11 Lunar Sample Information Catalogue** (revised). Curator's Office, JSC 12522

LSPET (1969) Preliminary examination of lunar samples from Apollo 11. *Science* **165**, 1211-1227.

Schmitt H.H., Lofgren G., Swann G.A. and Simmons G. (1970) The Apollo 11 samples: Introduction. *Proc. Apollo 11 Lunar Science Conf.* 1-54.