# 61224,6

Micro-Gabbro 0.34 grams

## **Introduction**

61224 contained several white and off-white particles as well as 3 peppered black and white rock fragments that appeared to fit together as one (figure 1). Thin section studies showed this to be a shocked lunar gabbro (*albiet very fine-grained for a gabbro*).

# **Petrography**

Marvin and Warren (1980) describe 61224,6 and liken it to a eucritic meteorite – accept that the Fe/Mn ratio proves it is lunar in origin. Marvin (1972) originally called it "gabbroic microporphry". The gabbroic texture and mineral mode is illustrated in figure 2. Grain size is about 1- 2 mm.

Marvin and Mosie (1980) wrote 61224,6: "is a coarsegrained, pristine, plutonic gabbro with a cumulate texture in which chains of anhedral hyperthene and augite grains coexist with plagioclase  $(An_{83})$ . The plagioclase has been shocked in situ to a leafy, optically-randomized state, and the grain boundaries are occupied by a selvage of pyroxene-plagioclase glass containing minute cyrstallites".

# **Mineralogical Mode**

40 % plagioclase 60% pyroxene

#### **Mineralogy**

Olivine: none

**Pyroxene:** Marvin and Warren (1980) and Takeda et al. (1981) studied the composition and structure of pyroxene (figures 3 and 4). The Fe/Mn ratio of pyroxene seems to indicate that this is a lunar gabbro, not a meteorite (figure 5). Bersch et al. (1983) give precise analyses of high- and low-Ca pyroxene in 61224,6.

**Plagioclase:** Marvin and Warren (1980) reported plagioclase as  $An_{79.87}$ .

*Selvage glass:* Marvin and Warren (1980) reported the composition of shock-melted glass in 61224,6.







*Figure 2: Thin section photo of 61224,6 (from Marvin and Warren (1980).* 

#### Table 1. Chemical composition of 61224

reference weight SiO2 % TiO2 Al2O3 FeO MnO MgO CaO Na2O K2O P2O5 S % sum	Marvin80 Warren80 50.7 0.4 13.2 9.91 0.16 12.77 11.6 0.91 0.02	(a) (a) (a) (a) (a) (a) (a)
Sc ppm V	20.8	(a)
Cr Co Ni Cu	1990 23.6 8	(a) (a)
Zn Ga	4 2 8	(a) (a)
Ge ppb As Se Rb Sr Y	4.3	(a) (a)
Zr Nb Mo Ru Rh Pd ppb Ag ppb	170	(a)
Cd ppb In ppb Sn ppb Sb ppb Te ppb Cs ppm	4.1	(a)
Ba La Ce Pr	32 1.47 4.3	(a) (a) (a)
Nd	9	(a)
Eu	1.43	(a) (a)
Tb	0.22	(a)
Dy Ho Er Tm		
Yb	1.06	(a)
Hf	0.55	(a) (a)
Ta W ppb	0.16	(a)
Re ppb	12.6	(a)
Ir ppb	0.15	(a)
Au ppb	0.08	(a)
Th ppm U ppm	0.19 0.6	(a) (a)
technique:	(a) INAA	()



compiled by C Meyer





Figure 4: Takeda et al. (1981) determined the composition of fine pyroxene lamallae in 61224 (small dots).



Figure 5: Composition of pyroxene in 61224,6 compared with pyroxene from meteorites and other lunar samples (Marvin and Warren 1980).

# **Chemistry**

Marvin and Warren (1980) and Warren and Wasson (1980) both reported the composition of 61224,6 (table). Warren (1992) termed it "pristine".

# **Radiogenic age dating**

None

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## **Processing**

Marvin and Mosie (1980) searched for additional pieces of 61224,6 in the finer fractions of this soil (61223 and 61222).



#### **References for 61224,6**

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