

10092

Ilmenite Basalt (low K)

46 grams



Figure 1: Photo of 10092,0. NASA S76-25871. Sample is about 3 cm.

Introduction

10092 is a low-K, ilmenite basalt (figure 1). It was originally labeled 10002,22, but because of its size, was renumbered 10092. It was returned in rock box #1003.

Petrography

According to Beatty and Albee (1978), 10092 is similar in texture to 10045 and 10020. These rocks have an overall texture characterized by an open network of randomly-oriented plagioclase laths and ilmenite platelets with dominant pyroxene in between. Minor minerals include a silica phase, ulvöspinel, Cr-spinel, troilite and a K-rich glass.

Olivine in 10092 is too Mg-rich to be in equilibrium with the bulk composition, indicating that it may have cumulate origin.

Mineralogical Mode for 10092

Beatty and
Albee 1978

Olivine	4.9
Pyroxene	45
Plagioclase	31
Ilmenite	15.6
Glass	0.14
silica	2.6
troilite	0.23
phosphate	0.06

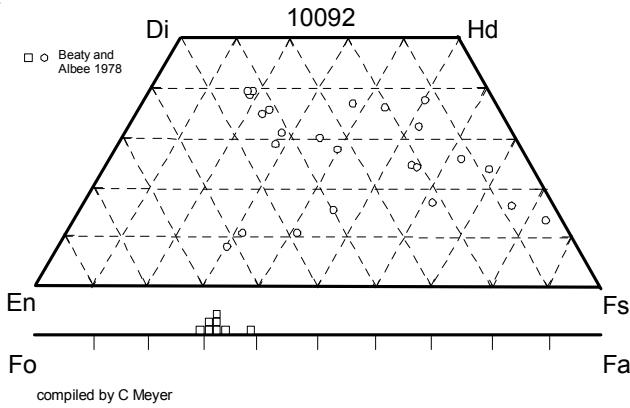
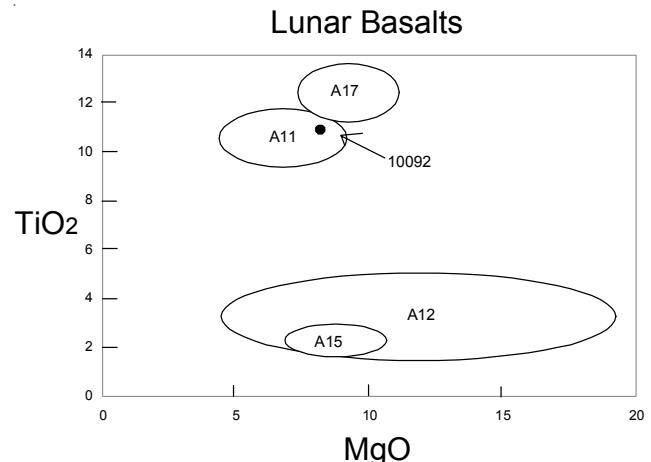


Figure 2: Pyroxene and olivine composition of 10092 (from Beatty and Albee 1978).



Mineralogy

Olivine: Olivine is Fo₇₂₋₆₀.

Pyroxene: Pyroxene has compositional zoning (figure 2).

Plagioclase: Plagioclase is An₉₂₋₈₂.

Ilmenite: Ilmenite has 2.5% MgO (Beatty and Albee 1978).

Chemistry

Rhodes and Blanchard (1980) obtained an analysis of 10092 (table 1, figures 3 and 4).

Radiogenic age dating

10092 has not been dated.

Processing

Apollo 11 samples were originally described and cataloged in 1969 and “re-cataloged” by Kramer et al. (1977). There are 5 thin sections.

List of Photo #s for 10092

S76-25871 – 76 color mug

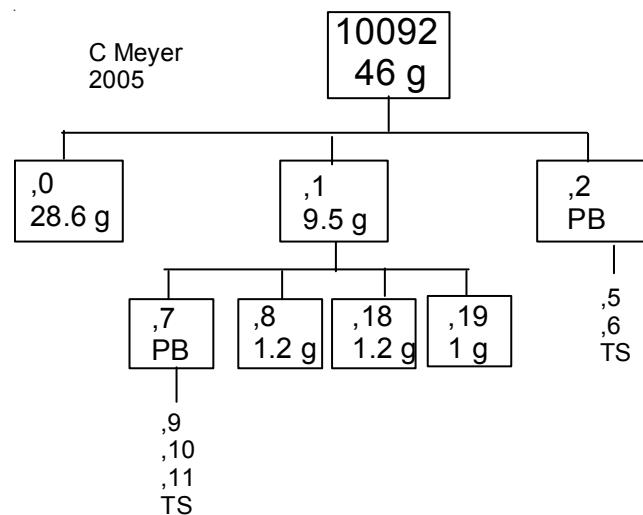
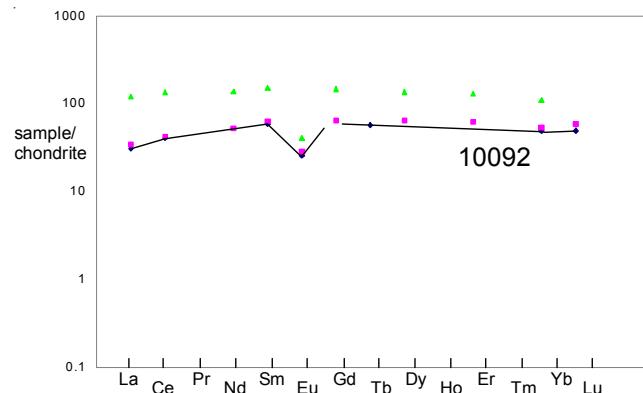


Table 1. Chemical composition of 10092.

reference	Rhodes80	Beaty 78	Neal2001
weight			
SiO ₂ %	38.85	(a)	38.05 (c)
TiO ₂	10.87	(a)	12.74 (c)
Al ₂ O ₃	9.44	(a)	9.66 (c)
FeO	19.35	(a)	18.38 (c)
MnO	0.3	(a)	0.23 (c)
MgO	8.52	(a)	8.9 (c)
CaO	10.99	(a)	11.38 (c)
Na ₂ O	0.34	(b)	0.33 (c)
K ₂ O	0.06	(a)	0.02 (c)
P ₂ O ₅	0.07	(a)	0.03 (c)
S %			0.11 (c)
sum			
Sc ppm	82		85.6 (d)
V			96 (d)
Cr	2930	(b)	3010 (c) 2446 (d)
Co	19.3		21.4 (d)
Ni			3.6 (d)
Cu			43.7 (d)
Zn			72.7 (d)
Ga			3.64 (d)
Ge ppb			
As			
Se			
Rb			0.81 (d)
Sr			144 (d)
Y			92 (d)
Zr			210 (d)
Nb			18.2 (d)
Mo			0.06 (d)
Ru			
Rh			
Pd ppb			
Ag ppb			
Cd ppb			
In ppb			
Sn ppb			
Sb ppb			20 (d)
Te ppb			
Cs ppm			0.03 (d)
Ba			73.2 (d)
La	7.3	(b)	8.03 (d)
Ce	25	(b)	24.9 (d)
Pr			4.41 (d)
Nd			23.2 (d)
Sm	8.9	(b)	9.15 (d)
Eu	1.48	(b)	1.57 (d)
Gd			13.4 (d)
Tb	2.1	(b)	2.37 (d)
Dy			16 (d)
Ho			3.34 (d)
Er			9.73 (d)
Tm			1.39 (d)
Yb	8.1	(b)	9.03 (d)
Lu	1.22	(b)	1.23 (d)
Hf	6.8	(b)	7.56 (d)
Ta	1.4	(b)	1.18 (d)
W ppb			70 (d)
Re ppb			
Os ppb			
Ir ppb			
Pt ppb			
Au ppb			
Th ppm			0.68 (d)
U ppm			0.22 (d)
technique: (a) XRF, (b) INAA, (c) elec. Probe, (d) ICP-MS			

References for 10092

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James O.B. and Jackson E.D. (1970) Petrology of the Apollo 11 ilmenite basalts. *J. Geophys. Res.* **75**, 5793-5824.

Kramer F.E., Tweddell D.B. and Walton W.J.A. (1977) **Apollo 11 Lunar Sample Information Catalogue** (revised). Curator's Office, JSC 12522

LSPET (1969) Preliminary examination of lunar samples from Apollo 11. *Science* **165**, 1211-1227.

Neal C.R. (2001) Interior of the moon: The presence of garnet in the primitive deep lunar mantle. *J. Geophys. Res.* **106**, 27865-27885.

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Schmitt H.H., Lofgren G., Swann G.A. and Simmons G. (1970) The Apollo 11 samples: Introduction. *Proc. Apollo 11 Lunar Science Conf.* 1-54.