12022 Ilmenite Basalt 1864.3 grams



Figure 1: PET photo of 12022,0 with caked on dirt (before dusting). NASA photo # S69-64108.

Introduction

Neal et al. (1994) classify 12022 as an ilmenite basalt while James and Wright (1972) had termed it an ilmenite-bearing olivine basalt. It has been dated at 3.2 b.y. All surfaces, except perhaps the flat end, were apparently covered with micrometeorite craters (Hörz et al. 1971), making this rock less than ideal for cosmicray depth profiles.

Petrography

The petrography of 12022 is discussed in Weill et al. (1971), Brett et al. (1971), McGee et al. (1977). McGee et al. describe 12022 as "a medium grained porphyritic basalt characterized by subhedral olivine (0.3 mm) and pyroxene (1-2 mm) phenocrysts. Several olivine phenocrysts are epitaxially overgrown with pyroxene. The matrix consists of feathery intergrowths of parallel feldspar tablets (0.05-1 mm), subrounded ilmenite laths

(0.03-0.2 mm), anhedral pyroxene crystals (0.6-0.8 mm) and minor glassy mesostasis".

Ilmenite in 12022 has an interesting cross-cutting, parallel, skeletal habit.

Mineralogy

Olivine: Butler (1972) determined the minor element content of olivine in 12022.

Pyroxene: Weill et al. (1971) and Brett et al. (1971) determined that the cores of pyroxene phenocrysts in 12022 are Ca-rich, zoning outward to Fe-rich (figure 4). The zonation of minor elements (Al, Ti, Cr) in pyroxene in 12022 is discussed in Bence and Papike (1972).



Figure 2: Refected light photo of thin section 12022,10 showing alignment of ilmenite, clumps of pyroxene phenocrysts in fine-grained variolitic groundmass. Field of view 1 cm. NASA photo # S70-24742.

Plagioclase: Plagioclase is An₉₁₋₈₅.

Opaques: Ilmenite laths occur in groups, cutting trough the matrix but not intersecting the phenocrysts (McGee et al. 1977). Subrounded octahedra of Cr-spinel, with or without rims of Ti-spinel, occur in the matrix and as inclusions in olivine and pyroxene phenocrysts.

Mineralogical Mode of 12022

_	McGee et	Neal et	Brett et
	al. 1977	al. 1994	al. 1971
Olivine	16-33	19.5	16.5
Pyroxene	30-59	56	58.6
Plagioclase	12-26	12.2	12
Opaques	9-23	~9	11.2
"silica"		0.2	
mesostasis	1	2.3	1.6

Metal: Brett et al. (1971) determined the Ni content of minute metallic iron grains in 12022 (figure 5).

Chemistry

Kushiro et al. (1971), Engel et al. (1971) and Snyder
et al. (1997) determined relatively high TiO₂ for 12022 (figure 7). Numerous investigators determined trace elements (table 1, figure 6). Kaplan and Petrowski (1971) determined about 20 ppm carbon and 800 ppm sulfur in 12022. Moore et al. (1971) reported about 40 ppm carbon and 44 ppm nitrogen. Rees and Thode (1972) reported 914 ppm sulfur.

Radiogenic age dating

Alexander et al. (1972) determined an age for 12022 of 3.18 ± 0.04 b.y. by the Argon plateau method. Snyder

Lunar Sample Compendium C Meyer 2011



Figure 3: Photomicrographs of thin section 12022,9 (plane-polarized, crossed-nicol). Field of view 2.6 mm. NASA phot #s S70-49455-456

et al. (1997) reported the isotopic composition of Sr and Nd.

Cosmogenic isotopes and exposure ages

Not reported !

Other Studies

Bogard et al. (1971) reported the content and isotopic composition of rare gases in 12022. Barber et al. (1971) determined the track density and erosion rate (figure 8).



Figure 4: Pyroxene and olivine composition for 12022 (adapted from Weill et al. 1971, Brett et al. 1971).



Figure 5: Histogram of Ni conentrations of metal grains in 7 lunar samples (lifted from Brett et al. 1971).

reference weight	Kushiro 1 g	071	LSPET70	Hubbar 188 mg	d71	Weismani 188 mg	n75	Murthy	71	Haskin	71	Engel7	1	Taylor	71 Tats71
SiO2 % TiO2 Al2O3 FeO MnO MgO CaO Na2O K2O P2O5 S % sum	42.33 4.54 9.12 22.06 0.26 11.58 9.37 0.29 0.07 0.02	 (a) 	36 5.1 11 22 0.17 13 11 0.36 0.068	0.24 0.065	(b)	0.065	(b)	0.051	(b)			43.2 5.16 9.04 21.44 0.25 10.43 9.56 0.47 0.07 0.13	 (a) 		
Sc ppm V Cr Co Ni Cu Zn Ga Ge ppb As	3831	(a)	52 65 2650 36 40									55 180 3800 44 42 8		55 150 3300 52 42 5	(d) (d) (d) (d) (d) (d)
Se Rb Sr Y Zr Nb Mo Ru Rh Pd ppb Ag ppb Cd ppb In ppb Sn ppb			0.17 160 62 160	0.738 143	(b) (b)	0.738 143	(b) (b)	0.819 138	(b) (b)			130 68 180		140 64 135 6	(d) (d) (d)
Sb ppb Te ppb														0.39	(u)
Cs ppm Ba La			38			59.5	(b)	148	(b)	5.81	(c)	70		0.03 60 6.3	(d) (d) (d)
Ce Pr				17.4	(b)	17.4	(b)			16.7 (c)	(c)			19 (3 ((d) (d)
Nd Sm Eu Gd				14.4 5.38 1.26 7.71	(b) (b) (b) (b)	14.4 5.38 1.26 7.71	(b) (b) (b) (b)			19 6.31 1.32 9.2	(c) (c) (c) (c)	(c) (c) (c) (c)	16 (d) 6.4 (d) 1.4 (d) 9.7 (d)	(d) (d) (d) (d)	
Dy				9.37	(b)	9.37	(b)			1.56 (C) 10.8 (C)	(c) (c)			12	(d) (d)
Er				5.42	(b)	5.42	(b)			5.8	8 (c)	8.2 (d) 1.3 (d) 7.2 (d)	(d) (d)		
Yb				5.69	(b)	5.06	(b)			5.34 (c)			10		(d) (d)
Hf Ta W ppb Re ppb Os ppb Ir ppb Pt ppb						0.18	(b)			0.707	(0)			5.2	(d)
Au ppb Th ppm U ppm <i>technique:</i>	(a) con	vent	tional wet, (k) IDMS,	(c)	INAA, (d) S	SSM	S, (e) R	NAA					0.75 0.17	(d) 0.71 (d) 0.198

(b) (b)

Table 1a. Chemical composition of 12022.

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Table 1b. Chemical composition of 12022.

reference	Baedecke	Snyder97	7	Neal2001		
weight SiO2 % TiO2 Al2O3 FeO MnO MgO CaO Na2O K2O P2O5 S % sum			43.2 5.16 9.04 21.44 0.25 10.43 9.56 0.47 0.07 0.13			
Sc ppm V Cr Co Ni			3300 52.6 47.4	(f) (f) (f)	59.6 156 3004 29.1	(f) (f) (f) (f)
Cu Zn Ga Ge ppb As	2 3.9	(e) (e)	17.6 11.5 4.24	(f) (f) (f)	15.7 11.3 3.64	(†) (f) (f)
Rb Sr Y Zr Nb Mo Ru Rh			0.841 142.5 64.2 160.1 7.04	(f) (f) (f) (f) (f)	0.78 147.5 51 122 6 0.11	(f) (f) (f) (f) (f) (f)
Pd ppb Ag ppb Cd ppb In ppb Sn ppb Sb ppb Te ppb	6.4 1.6	(e) (e)	105	(f)		
Cs ppm Ba La Ce Pr Nd Sm Eu Gd Tb Dy Ho Er Tm			0.058 58.4 6.5 17.9 3.35 19.2 6.58 1.45 7.94 1.66 10.39 2.1 5.81 0.82	(f) (f) (f) (f) (f) (f) (f) (f) (f) (f)	0.03 57 5.59 16.9 2.93 15.3 5.63 1.3 8.74 1.48 10.2 2.09 6.13 0.84	(f) (f) (f) (f) (f) (f) (f) (f) (f) (f)
Yb Lu Hf Ta			0.82 5.36 0.78 0.381	(f) (f) (f)	0.84 5.51 0.77 4.25 0.39	(f) (f) (f) (f) (f)
W ppb Re ppb Os ppb Ir ppb Pt ppb Au ppb	0.09	(e)			160	(f)
Th ppm U ppm <i>technique:</i>	(e) RNAA,	(f) IC	0.987 0.28 CP-MS	(f) (f)	0.63 0.19	(f) (f)



Figure 6: Comparison of rare-earth-element composition of 12022 by neutron activation analysis (Haskin et al. 1971) and spark-source mass spectroscopy (Taylor et al. 1971) with isotope dilution mass spectroscopy (line, Hubbard and Gast 1971, Wiesman et al. 1975).

Helsley (1971) found that 12022 has significant magnetic remanance and Chung et al. (1971) determined the dielectric properties.

Gibson and Hubbard (1972) experimentally studied the volatile depletion for 12022.

Processing

A thick slab (B,14) was cut from the middle of 12022 with a circular saw (figure 9) and a column (,17) was cut from the slab with a wire saw (figure 12). End piece (A,13) was also subdivided with a wire saw (figure 10). A large piece of 12022 is on public display in Wales, England (figure 13).



Figure 7: Composition of 12022 compared with that of other lunar basalts.



Figure 8: Track density as function of depth in 12022 (from Barber et al. 1971).

List of Photo #s for 12022

S69-64083	color mug
S69-64108	color mug
S70-16784 - 785	TS color
S70-20956	TS color
S70-49560 - 561	
S70-49455 - 460	TS color
S74-24900 - 902	display
S79-27121 - 122	TS color







Figure 9: Group photo of 12022,0 after 1.5 mm slab was cut.

Figure 10: Group photo of 12022,13. Scale is in mm.

THE CUTTING AND CHIPPING OF SLICE 'B' NO. 12022,14 DRAWING COMPLETED MAY 25, 1970





Figure 11: End piece of slab 12022, 14.



Figure 12: Column (12022,17) cut from slab (12022,14). Scale in mm.





Figure 13: Display case for 12022,92. NASA S74-24901.

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