

**14305**  
Crystalline Matrix Breccia  
2497 grams

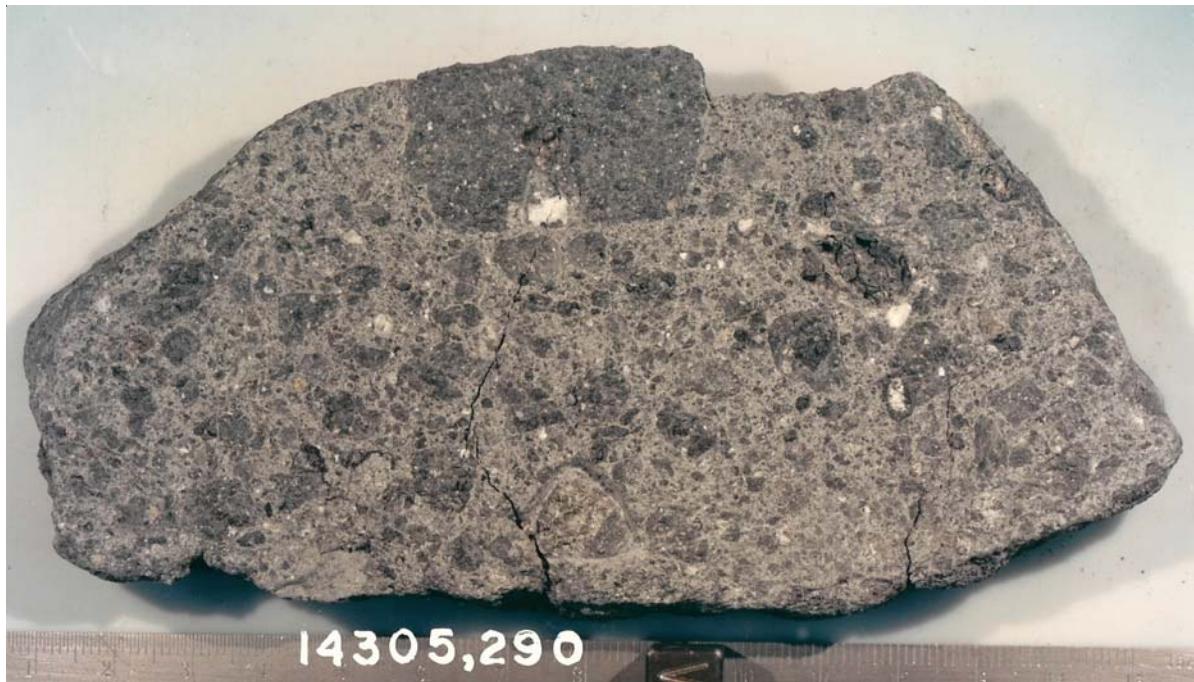


Figure 1: Sawn surface of 14305 showing dark clast in lighter matrix. Cube and scale are in cm.  
NASA S86-30442.

### Introduction

Lunar sample 14305 is a Fra Mauro breccia similar in texture and clast distribution to 14321 (figure 1). This “football-sized” rock was collected about 100 m from the LM landing site at Apollo 14. A piece broke off during transit and was originally numbered 14302. When both pieces were found to exactly fit together, 14302 was renumbered 14305,18.

Since 14305 has zap pits on both sides and has higher  $^{26}\text{Al}$  on the bottom, it has had a complicated exposure history (i.e. has turned over). Using  $^{53}\text{Mn}$ , Herpers et al. (1974) showed that the rock has “tumbled” (figure 2). Eugster et al. (1984) concluded that 14305 did not originate from Cone Crater, mainly because its exposure age is slightly higher than the age of Cone Crater.

One large, glass-lined, micrometeorite pit records the location on rock 14305 that almost ruptured the rock (see figure 3 in Gault et al. 1972).

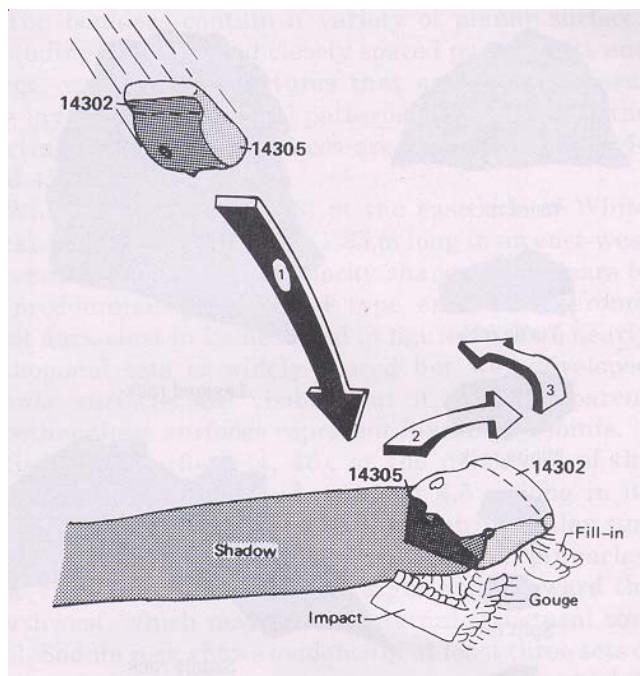


Figure 2: Illustration of emplacement and tumbling of 14305-14302 on lunar regolith.

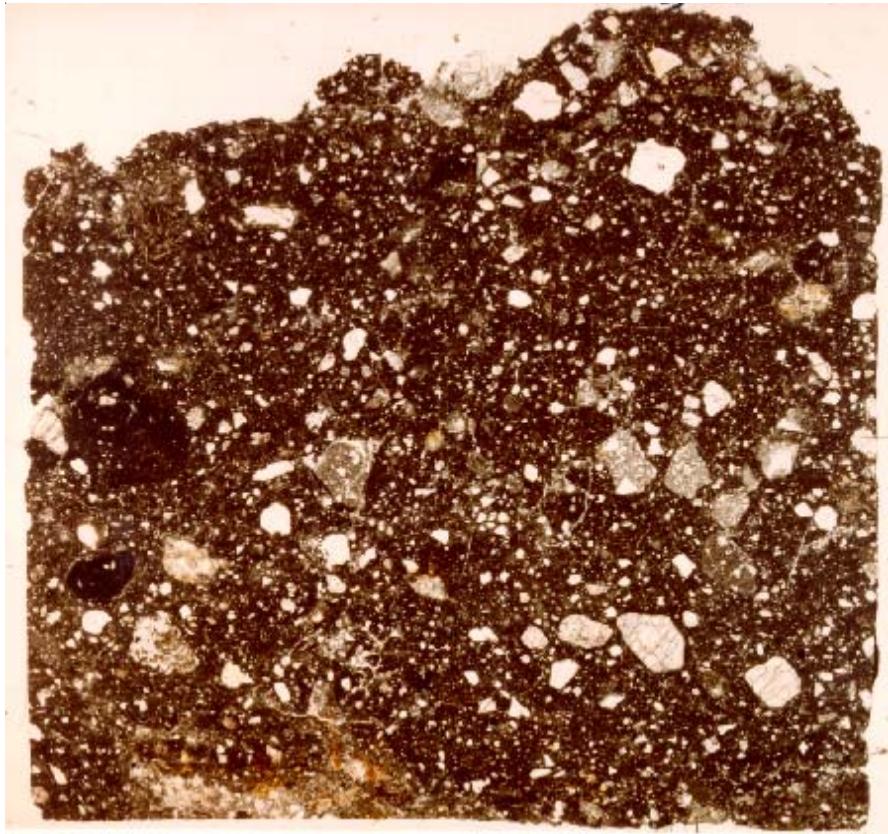


Figure 3: Thin section photomicrograph of 14305, 255. NASA photo #S80-42342. Field of view about 2 cm.

The crystallization age of 14305 was found to be 3.92 b.y. with an exposure to cosmic rays of about 28 m.y. However, various clasts in this breccia are older, and zircons are found to be extremely old (4.35 b.y.).

### Petrography

Petrographic descriptions of breccia 14305 are found in Juan et al. (1972), Klein and Drake (1972), Dence and Plant (1972), Lovering et al. (1972), Carlson and Walton (1978), Shervais and Taylor (1983). According to the classification schemes of Simonds et al. (1977) and Stöffler et al. (1980), 14305 is a clast-rich, crystalline matrix breccias (figure 4). The seriate matrix (down to <25 microns) consists of intergrown, blocky decussate pyroxenes with plagioclase laths and granular opaques (figure 3). Lithic clasts >1 mm comprise about 30% of the sample (Wilshire and Jackson 1972). Dark gray microbreccia is the dominant clast type (figure 1), with basalt, plagioclase-olivine cumulates and granites as minor clast types (figure 8). Juan et al. found that the fine-grain basalt clasts were like 14310. They also reported gabbroic anorthosite from the highlands. Shervais et al. (1985) also studied

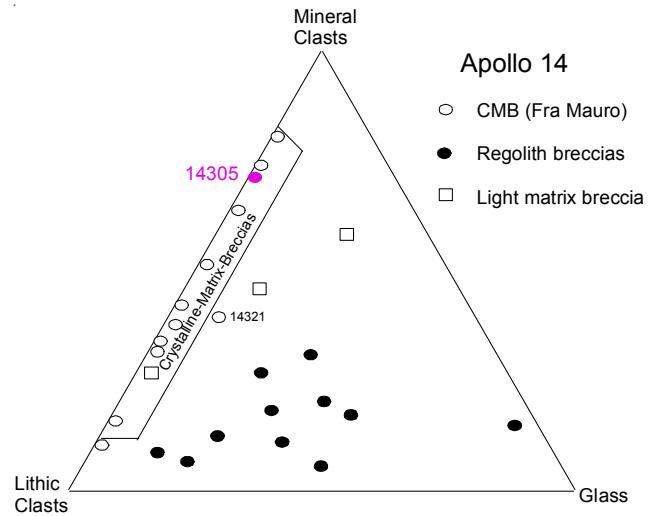


Figure 4: 14305 is a crystalline matrix breccia (Simonds et al. 1977).

the plagioclase rich basalt clasts in 14305 and termed them VHK basalts.

Twedell et al. (1978) mapped some of the surfaces of 14305 and Carlson and Walton (1978) summarized what was known up to that time. Shervais and Taylor (1983) documented what was known about breccia

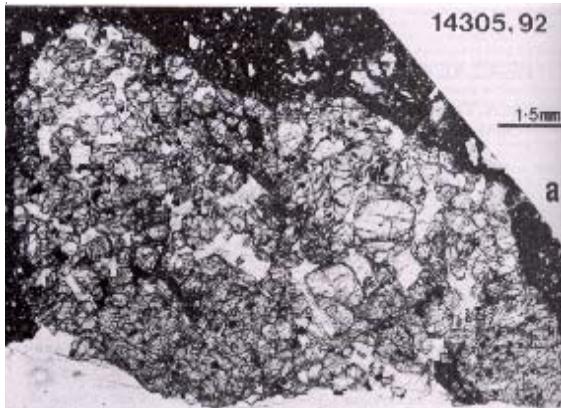


Figure 5: Thin section photomicrograph of basalt clast in 14305 (from Taylor et al. 1983). This is the olivine gabbronorite that proved to be 4.2 b.y.

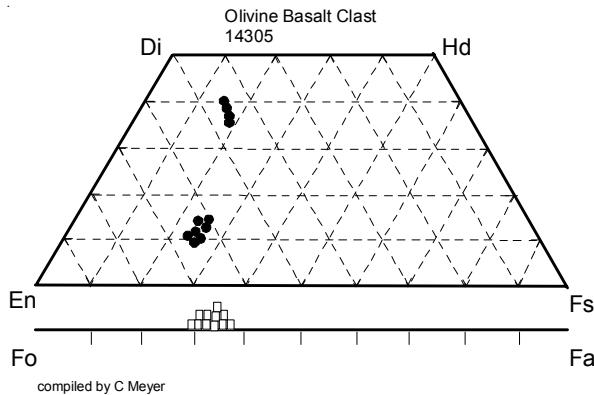


Figure 6: Composition of mafic minerals in olivine basalt clast in 14305, 92 (from Taylor et al. 1983).

sample 14305 in 1983 and began a search for rock clasts as part of a “pull-apart” project. They “mapped” some of the new surfaces created by new saw cuts. The most important finding from the study of 14305 was the discovery a large clast of very old mare basalt (see below).

## Mineralogy

**Pyroxene:** Klein and Drake (1972) present pyroxene analyses – including orthopyroxene in the matrix. Pyroxene compositions of the clasts in this breccia are reproduced in figures 5, 6 and 10.

**Spinel:** Dence and Plant (1972) give an analysis of pink spinel in 14305, 5.

**Phosphate:** Brown et al. (1972) and Warren et al. (1983) give analyses of whitlockite. Analyses for apatite from clasts can be found in Shervais et al. (1984).

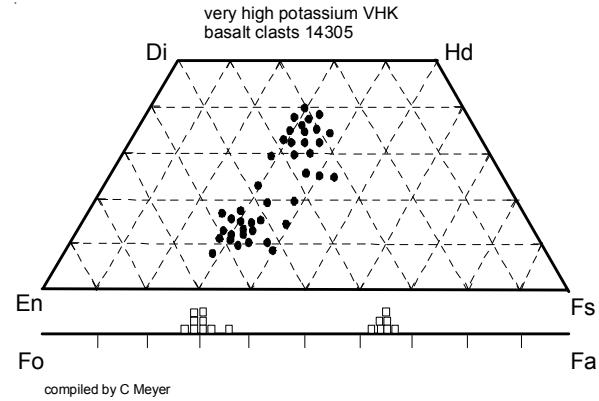


Figure 7: Olivine and pyroxene compositions of VHK basalt clasts in 14305 (after Shervais et al. 1983).

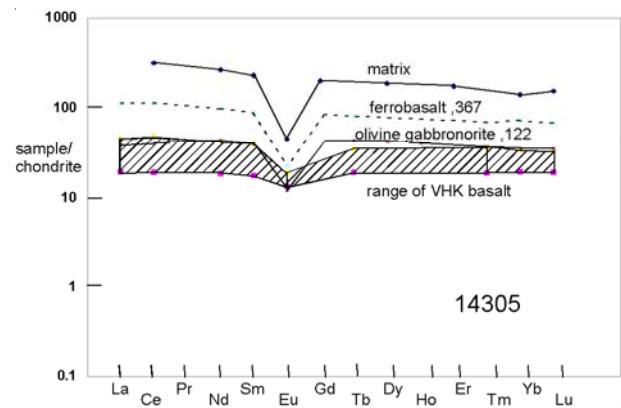


Figure 8: Normalized rare-earth-element diagram for 14305 matrix and some of its basalt clasts. (data from table 2)

**Ternary Feldspar:** Shervais and Taylor (1983) reported feldspars with roughly equal amounts of K, Na and Ca in granophryic clasts in 14305.

**Zircon:** Lovering et al. (1972) found U-rich zircon in a monzonite clast. Meyer et al. (1988, 1991) studied zircons in two norite clasts in 14305. Taylor et al. (2009) studied zircons from sawdust and Nemchin et al. (2008) dated 6 zircons in thin section.

## Significant clasts

### **,122 Olivine Gabbronorite (mare basalt)**

Hunter and Taylor (1983) and Taylor et al. (1983) describe a large clast of olivine gabbronorite (basalt) in thin section 14305, 92 (table 2, figure 5, 6 and 14). They dug the remainder of this clast out of the adjacent epoxy mount and determined the age by Rb-Sr internal mineral isochron 4.2 b.y. (see below). This clast is a mare basalt very similar to 12005 (see also Ari et al. 2006).

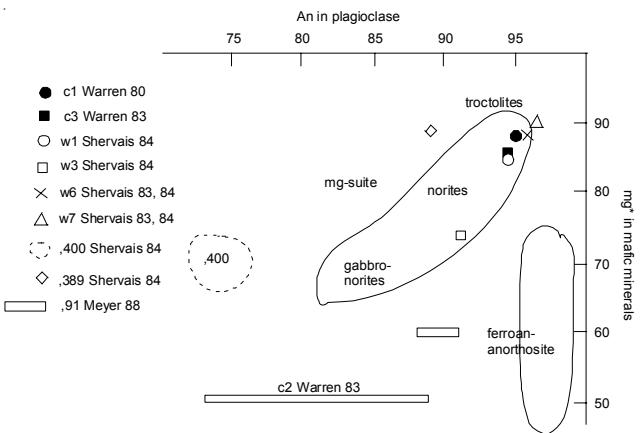


Figure 9: Composition of plagioclase and mafic minerals for white clasts in 14305 (see text for explanation).

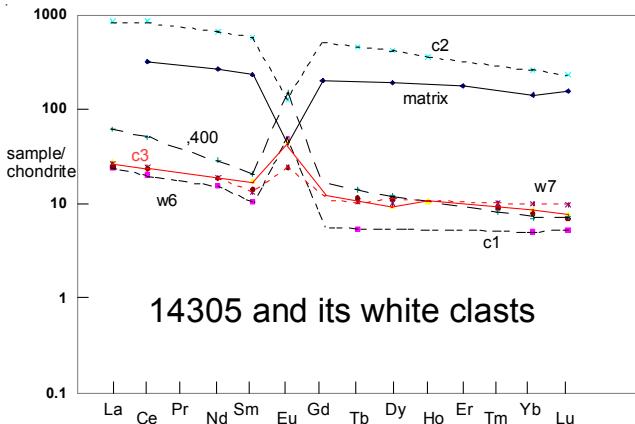


Figure 10: Normalized rare-earth-element diagram for 14305 matrix and its white clasts (data from table 2).

### ,304 VHK Basalt

Shervais et al. (1985) studied several high potassium basalt clasts (table 2, figure 7). One of these was dated by Shih et al. (1983) (figure 15).

### ,367 ferrobasalt

An analysis of an “ilmenite ferrobasalt” is presented in Shervais et al. (1985), but no details are given (figure 8).

### c1 from ,18

Warren and Wasson (1980) analyzed a white troctolite clast from the top corner of 14305,18. Thin section ,268 showed that this clast was 30% olivine ( $Fo_{87}$ ), 70% plagioclase ( $An_{95}$ ). This clast is similar to 14172 (which might be a piece of same rock). The clast is pristine ( $Ir < 0.05$  ppb).

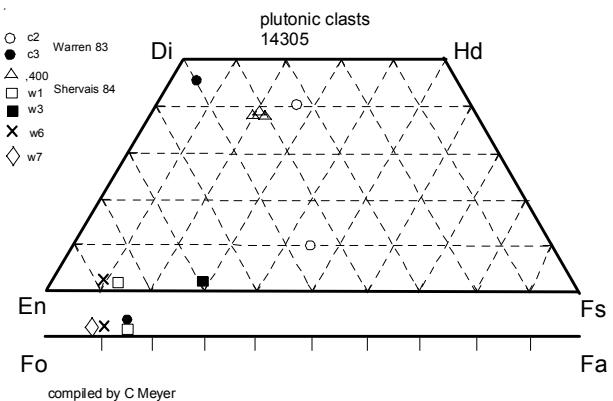


Figure 11: Pyroxene and olivine compositions of troctolite, anorthosite and norite clasts in 14305 (from Warren et al. 1983 and Shervais et al. 1983).

### c2 trace-element-rich “anorthosite”

Warren et al. (1983) analyzed a whitlockite-rich anorthosite clast 14305,283 (table 2, figure 10). The plagioclase ranged widely in composition ( $An_{68-90}$ ). Shervais et al. (1984) also studied an alkali anorthosite clast (.400) with plagioclase  $An_{71-77}$ .

### c3

Warren et al. (1983) also found a troctolite clast ( $7 \times 6$  mm) on 14305,27 with 90% plagioclase ( $An_{94}$ ), 10% mafic minerals ( $Fo_{85}$ ) – see table 2 and figure 11.

### w1

Shervais et al. (1984) identified a troctolitic anorthosite ,394 (no analysis) with 90% plagioclase ( $An_{95}$ ), 10% olivine ( $Fo_{85}$ ) and trace pyroxene (figures 9, 10 and 11).

### w3

Shervais et al. (1984) studied a rare norite clast in thin section (no analysis). It was 92% plagioclase ( $An_{91}$ ) and 8% enstatite ( $En_{73}$ ).

### w6 from ,27 and ,290

Shervais et al. (1983, 1984) analyzed (in duplicate) an enstatite troctolite clast ( $9 \times 5$  mm) that was dissected by the saw cut. The thin section showed 25% maskelynite ( $An_{96}$ ), 25% olivine ( $Fo_{89}$ ), and 5% enstatite ( $En_{90}$ ).

### w7 from ,27

Shervais et al. (1983, 1984) analyzed (in duplicate) a magnesium anorthosite clast that is made up mostly of plagioclase ( $An_{97}$ ) with trace olivine ( $Fo_{90}$ ). The chemical composition is similar to w6 (figure 9).

**Table 1. Chemical composition of 14305.**

reference	Keith72	Fields72	Yokoyama 72	LSPET72	Wanke 72	Willis 72	Philpotts72	Palme78	Wiik73
weight	380 g						sawdust		
SiO <sub>2</sub> %				49.2	(a)	48.35	(c)	47.92	(d) 48.2 47.92
TiO <sub>2</sub>				1.67	(a)	1.52	(c)	1.71	(d) 1.71 1.78
Al <sub>2</sub> O <sub>3</sub>				16.1	(a)	16.25	(c)	15.47	(d) 15.16 16.64
FeO				9.52	(a)	10.42	(c)	11.34	(d) 10.88 10.8 (e) 10.59
MnO				0.18	(a)	0.13	(c)	0.14	(d) 0.13 0.14 (e) 0.14
MgO				13	(a)	10.28	(c)	11.14	(d) 11.12 10.37
CaO				7.4	(a)	9.93	(c)	9.96	(d) 10.13 12.12
Na <sub>2</sub> O				0.85	(a)	0.76	(c)	0.73	(d) 0.87 0.8 (e) 0.89
K <sub>2</sub> O	0.64	(b)		1.2	(a)	0.64	(c)	0.68	(d) 0.73 0.64 (e) 0.62
P <sub>2</sub> O <sub>5</sub>								0.57	(d) 0.64 0.65
S %								0.094	(d) 0 0.09
sum								99.75	99.57
Sc ppm					24	(e)			25.3 (e) 24
V				52	(a)				30
Cr				1200	(a)	1330	(c)		1395 (e) 1369
Co				32	(a)	31	(c)		33.4 (e) 20
Ni				205	(a)	200	(c)		300 (e) 210
Cu				13	(a)	10.9	(c)		20
Zn						2.1	(c)		
Ga						5	(c)		4.9 (e)
Ge ppb						44	(c)		
As						0.077	(c)		
Se									
Rb				31	(a)	25		17.9	(d) 19 (e)
Sr				200	(a)	190		162	(d) 166 (e) 185 (e) 160
Y				210	(a)			238	(d) 200
Zr				900	(a)			1158	(d) 1060 (e) 840
Nb				49	(a)			70.7	(d)
Mo									
Ru									
Rh									
Pd ppb									
Ag ppb									
Cd ppb									
In ppb									
Sn ppb									
Sb ppb									
Te ppb									
Cs ppm					1.36	(c)			
Ba				930	(a)	830	(c)	913	(d) 924 (e) 950 (e) 640
La				54	(a)	109	(c)		111 (e) 83
Ce						200	(c)		193 (e) 220 (e)
Pr						26	(c)		
Nd						140	(c)		121 (e) 137 (e)
Sm						23	(c)		34.4 (e) 32.4 (e)
Eu						2.6	(c)		2.56 (e) 2.6 (e)
Gd						38	(c)		40.1 (e)
Tb						7.4	(c)		6.9 (e)
Dy						43	(c)		46.7 (e) 43.9 (e)
Ho						6.5	(c)		
Er						32	(c)		28.3 (e)
Tm									
Yb				28	(a)	24.2	(c)		23.3 (e) 26 (e) 23
Lu						3.5	(c)		3.82 (e) 3.62 (e)
Hf						26	(c)		28.6 (e) 26.5 (e)
Ta						3.2	(c)		3.2 (e)
W ppb						1940	(c)		2.6 (e)
Re ppb									
Os ppb									
Ir ppb						10	(c)		11 (e)
Pt ppb									
Au ppb						6.7	(c)		
Th ppm	13.9	16.4	14.6	13.3	14.3	(b)		17.4	(c)
U ppm	3.8	4.13	3.8	3.8	3.8	(b)		5.15	(c)
technique	(a) Optical Emission, (b) Radiation counting, (c) INAA, (d) XRF, (e) IDMS								

**Table 2. Chemical composition of 14305 clasts.**

	monzonite Lovering 72 ts only	troctolite Warren 80 c1	anor? Warren 83 c3	troctolite Shervais 83 c2	anor. ,317 ,322	basalt Taylor 83 ,122	VHK bas. Shervais ,304	VHK Shih 86 ,304	K-anor. Shervais 84 ,400	Pxite ,389
SiO <sub>2</sub> %	59.6	43.64	43.4							
TiO <sub>2</sub>	0.6	0.04	0.54	1.69	0.13	0.13	3.6	2.4	0.2	<0.5
Al <sub>2</sub> O <sub>3</sub>	20.6	27.96	28.54		21.9	34	4	9.9	29.9	1.1
FeO	3.3	2.83	2.25	2.55	4.2	0.43	22.4	18.1	0.19	7.3
MnO		0.031	0.026	0.037	0.049	0.007	0.25	0.23	0.003	0.11
MgO	3.3	11.44	8.3		17.6	1.4	22	15	<1	35
CaO	7.9	14.28	15.96	16.8	11.7	20.7	6.5	9	15.3	0.73
Na <sub>2</sub> O	0.9	0.43	0.47	1.48	0.44	0.34	0.16	0.34	2.4	0.04
K <sub>2</sub> O	4.7	0.07	0.07	0.506	0.063	0.055	0.04	0.62	0.46	0.01
P <sub>2</sub> O <sub>5</sub>										
S %										
<i>sum</i>										
Sc ppm		1.78	2.6	8.6	3.9	0.75	56	50	0.54	4.7
V					70		190	120		40
Cr	140	201	111		2395	116	821	3968		
Co	19	12.6	3.6		17.8	2.5	56.5	37	0.4	28.5
Ni	19				20	10	150	60		110
Cu										
Zn		0.58								
Ga										
Ge ppb		12								
As										
Se										
Rb			5.3				0.606		14.18	
Sr		211	400	120	210		39.17		67.42	
Y										
Zr	120		196	<30	20	230	80			
Nb										
Mo										
Ru										
Rh										
Pd ppb										
Ag ppb										
Cd ppb										
In ppb										
Sn ppb										
Sb ppb										
Te ppb										
Cs ppm										
Ba	630	450	1180	140	110	<50	200		600	40
La	5.6	6.6	201	6.6	5.4	9.2	4.6		14.5	10.3
Ce	12.2	14.8	520	15	13	25	11.7		31	24
Pr										
Nd	7.1	8	305	9.4	7.3	20	8.5	8.447	13	15
Sm	1.56	2.16	86	2	1.95	6	2.6	2.803	3.1	2.5
Eu	2.72	2.6	7.2	1.45	1.35	0.7	0.75		8.4	0.13
Gd										
Tb	0.2	0.4	16.7	0.4	0.36	1.6	0.7		0.52	0.32
Dy		2.33	103	2.6	2.5	10.5	4.8		3	<3
Ho		0.6	20.1			2.3				
Er										
Tm						0.9	0.46		0.2	0.2
Yb	0.83	1.44	43	1.7	1	5.8	3.2		1.15	1
Lu	0.13	0.19	5.7	0.25	0.12	0.87	0.47		0.18	0.14
Hf	0.25	0.37	2.4	0.88	0.45	6.2	2.5		1.3	0.72
Ta	<0.3	0.07	0.49	0.05	0.07	1.6	0.65		0.22	0.2
W ppb										
Re ppb		0.6								
Os ppb										
Ir ppb		<0.05								
Pt ppb										
Au ppb		0.016								
Th ppm	0.22	0.46	21.6	0.35	0.52	1.2	0.61		1.3	0.94
U ppm	0.17	0.15	3	<0.1	<0.1	<0.4	<0.2		<0.5	<0.4
<i>technique (a) INAA, (b) IDMS</i>										

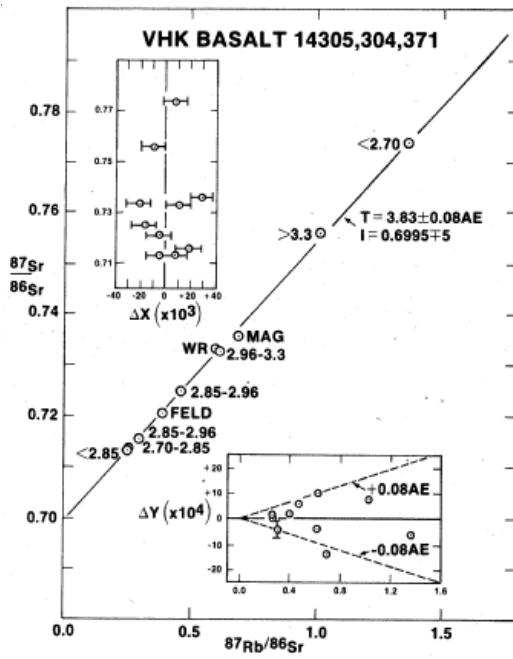


Figure 12: Rb-Sr isochron diagram of high-K basalt clast in 14305 (from Shih et al. 1986).

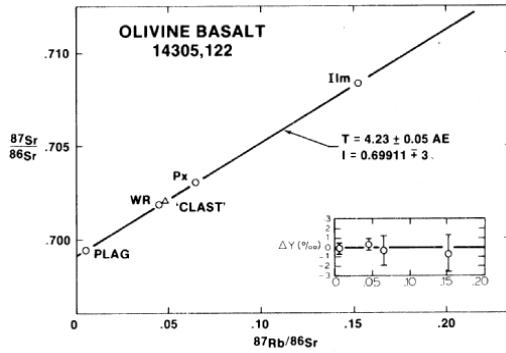


Figure 14: Rb-Sr isochron for clast in 14305 (from Taylor et al. 1983). Note the significant old age.

### Granite

Shervais and Taylor (1983) studied several small clasts of “micrographic granite” in thin sections 14305, 102 and ,111.

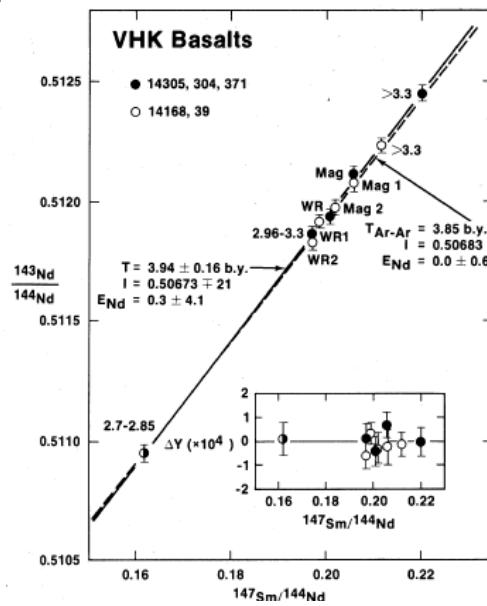


Figure 13: Sm-Nd isochron diagram for 14305 basalts clast (from Shih et al. 1986).

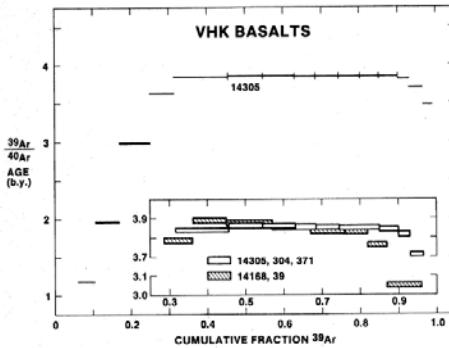


Figure 15: Ar plateau age for 14305 basalt clast (from Shih et al. 1986).

### ,77 Monzonite clast

Lovering et al. (1972) studied a 2 mm white fragment with ~35% plagioclase, 18% Ca-poor pyroxene in 44% interstitial mesostasis rich in Si and K. Apatite (50 ppm U), whitlockite (75 ppm U) and zircon (140-1300 ppm U) were accessory phases in this clast.

### Summary of Age Data for 14305 (b.y. = billion years)

	Ar release	Rb-Sr	Sm-Nd	U-Pb	
Matrix	$3.92 \pm 0.03$ b.y.				Eugster et al. (1984)
14305, 122		$4.23 \pm 0.05$			Taylor et al. 1983
14305, 304	$3.85 \pm 0.02$	$3.83 \pm 0.08$	$3.95 \pm 0.17$		Shih et al. 1986
14305, 91				$4.2$ b.y.	Meyer et al. (1991)
Zircons				$3.97$ to $4.35$ b.y.	Nemchin et al. (2008)
					Taylor et al. (2007)

**Caution: Not corrected for new radiogenic decay constants.**

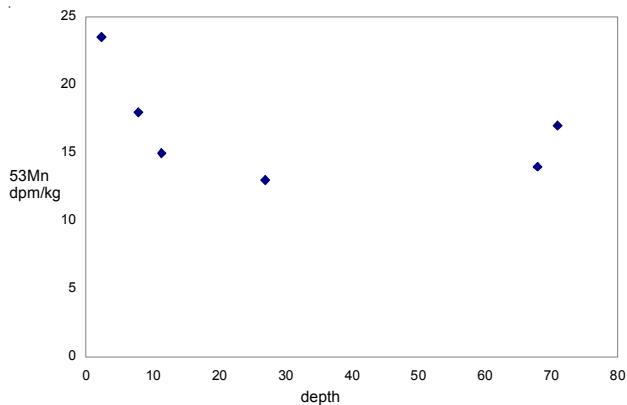


Figure 16: Depth profile for  $^{53}\text{Mn}$  in 14305. This is evidence for “tumbling” (Herpers et al. 1974).

### **,91 Gabbronorite**

Meyer et al. (1988) dated a small zircon included in an igneous rock clast with 48% plagioclase ( $\text{An}_{90}$ ), 48% pyroxene ( $\text{En}_{70}$ ), 2% ilmenite, apatite and whitlockite.

### **,103 Gabbronorite**

Meyer et al. (1988) reported another small clast of gabbronorite (with included zircon) in thin section 14305,103.

### **Chemistry**

Keith et al. (1972) determined  $\text{Th} = 13.9$  ppm for a large piece of 14305 by radiation counting (table 1). This is a good guide for which chemical analysis to use as representative of the bulk rock. Philpotts et al. (1972) analyzed the “sawdust” from 14305, which should be representative of the rock as a whole (figures 8 and 10). Analyses by Wanke et al. (1972), Wiik et al. (1973) and Palme et al. (1978) are in agreement and show very high K, P and REE content. Ni, Ir and Au are high, indicating a high content of meteoritic material. Jovanovic and Reed (1976) determined Ru and Os.

Warren et al. (1980, 1983) and Shervais et al. (1983, 1984) analyzed a large number of clasts in 14305 (table 2, figures 8 and 10).

### **Radiogenic age dating**

Eugster et al. (1984) determined an Ar release age of  $3.92 \pm 0.03$  b.y. for several subsamples (presumably matrix material). Taylor et al. (1983) and Ari et al. (2006) found the oldest mare basalt as a clast in this breccia (figure 14). Shih et al. (1986) dated one of the high K basalt fragments (figures 11, 12, 14). Meyer et al. (1991) were able to date a zircon in a gabbronorite

clast by the ion microprobe method. Taylor et al. (2009) dated zircons from sawdust of 14305, while Nemchin et al. (2008) dated zircons found in thin section.

### **Cosmogenic isotopes and exposure ages**

Eugster et al. (1984) determined an exposure age of  $27.6 \pm 1.5$  for several subsamples of 14305 by the  $^{81}\text{Kr}$  method. Since 14305 has higher  $^{26}\text{Al}$  on the bottom side ( $160 \pm 60$  dpm) than the top side ( $63 \pm 34$  dpm/kg), it is thought that it has turned over in the regolith (Yokoyama et al. 1972). Herpers et al. (1974) also demonstrated that the rock had “tumbled” while on the lunar surface. They found that the  $^{53}\text{Mn}$  depth profile was at a minimum in the middle of the rock (figure 16).

### **Other Studies**

Eugster et al. (1984) determined the rare gas content of 14305.

Herr et al. (1972) determined the activity of  $^{53}\text{Mn}$ . Keith et al. (1972) determined activity of  $^{26}\text{Al}$ ,  $^{22}\text{Na}$ ,  $^{54}\text{Mn}$ ,  $^{56}\text{Co}$  and  $^{46}\text{Sc}$ .

Herpers et al. (1974) studied the density of cosmic ray tracks in olivine and plagioclase as a function of depth in this rock. They also determined the enrichment of  $^{187}\text{Re}$  due to neutron capture.

### **Processing**

14302 was re-labeled 14305,18 when it was found to be a broken piece of 14305. The main mass of 14305 was cut thru the center to yield a large slab for allocations (figures 17 - 19). In 1982, 14305,27 was again sawn to produce slab ,290 and again in 1985 to create slab ,483. 14305,29 was cut to produce slab ,459. Numerous “clasts” were extracted from these new slabs as part of a “breccia-pull-apart” project by Larry Taylor.

Shervais and Taylor (1983) produced a guidebook for this rock, summarizing the work up to 1983. There are over 100 thin sections of this rock (figure 26). This rock is featured in the Lunar Petrographic Educational Thin Section Package (Meyer 2003).

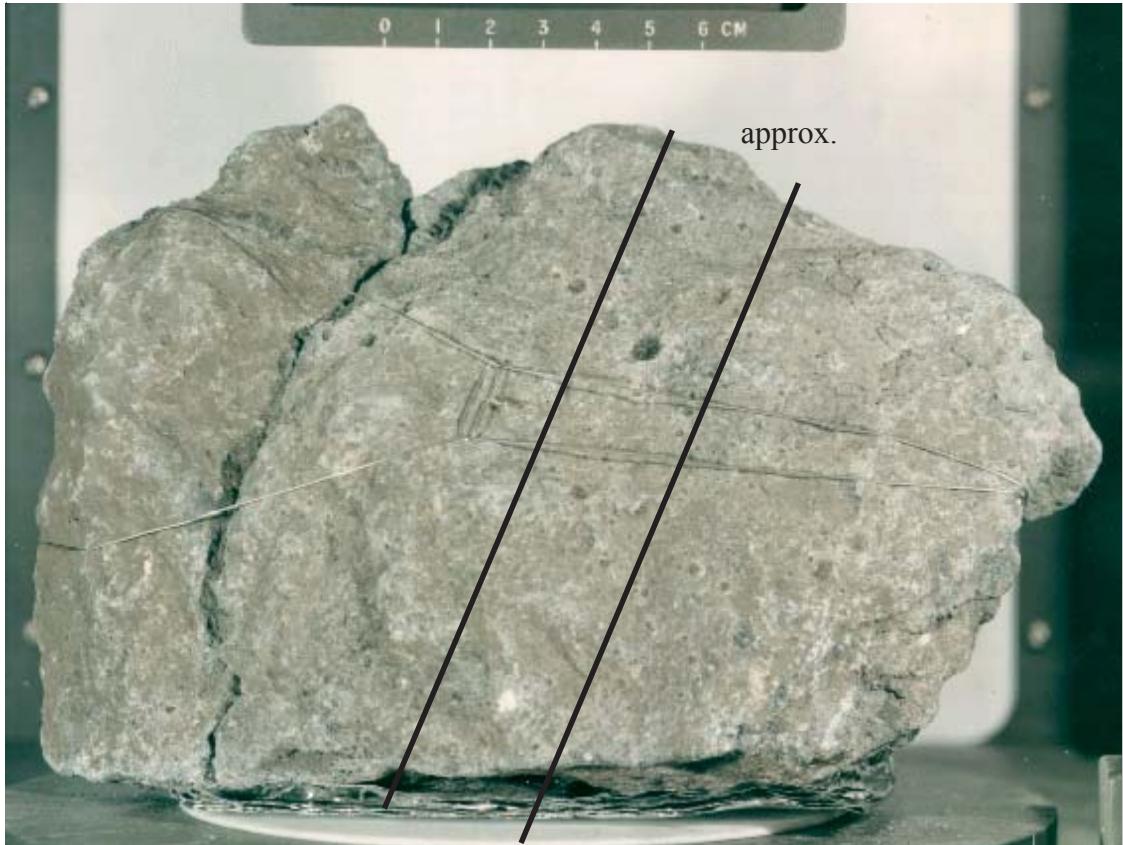


Figure 167: NASA photo of 14305 with 14302 attached. #S71-31391. Scale at top is in cm.

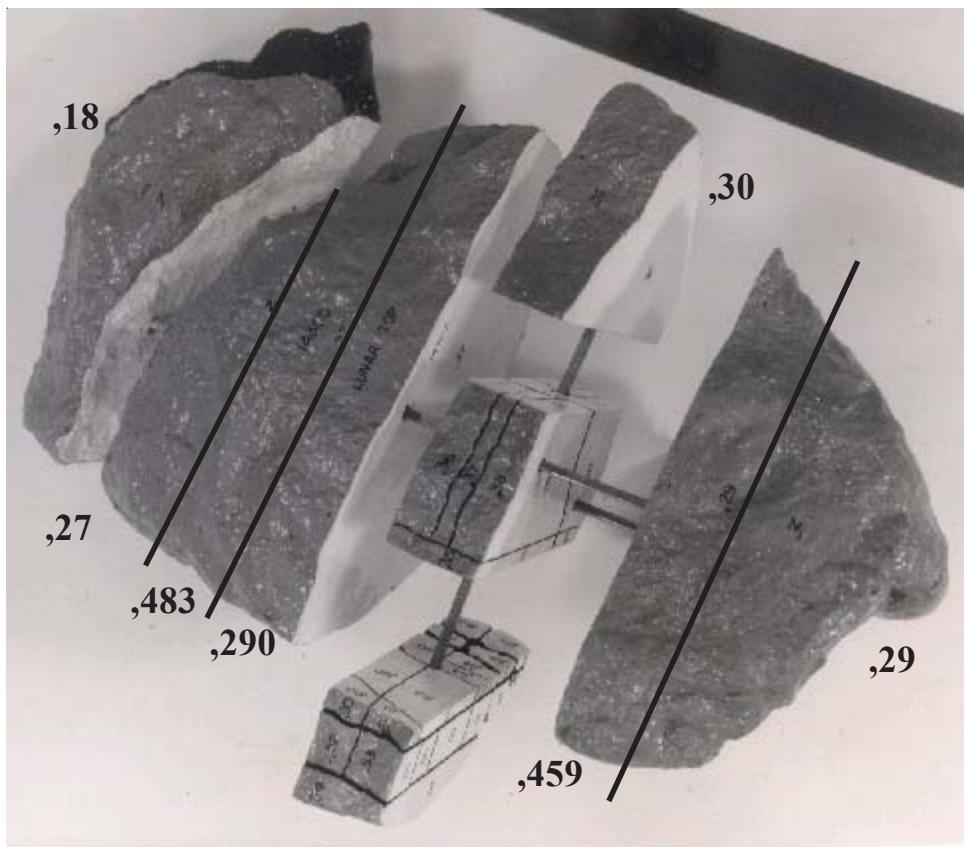


Figure 18: Exploded parts diagram for 14305 and 14302 (now called 14305,18)

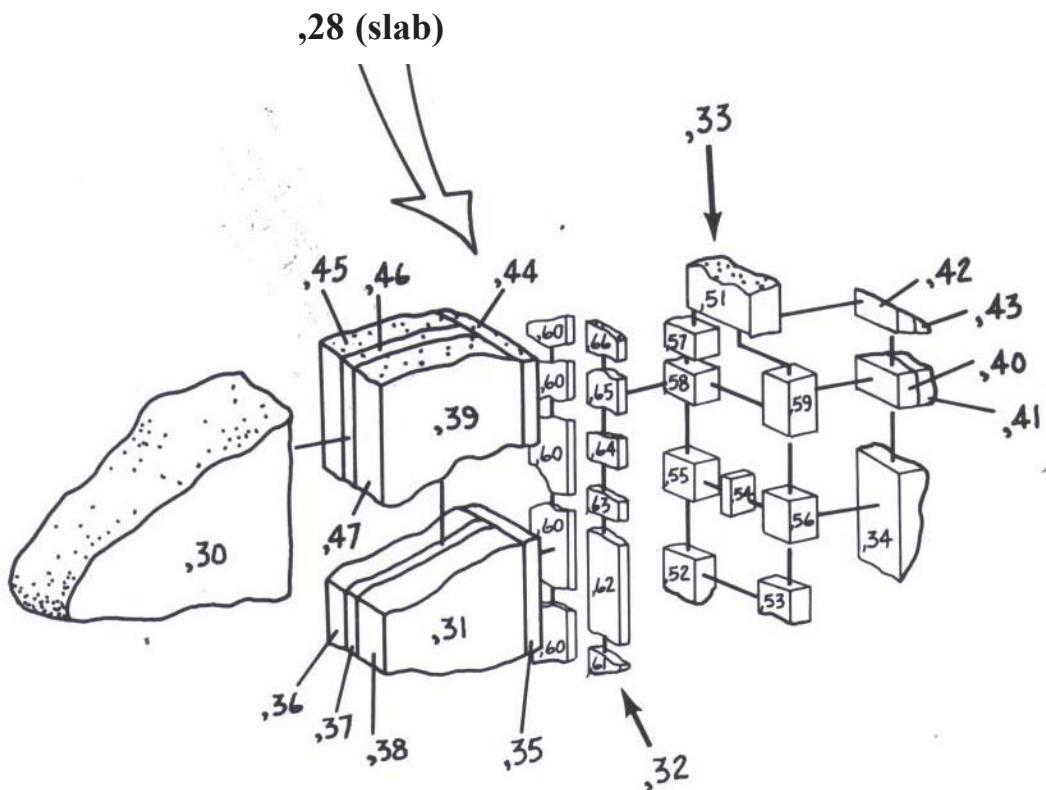
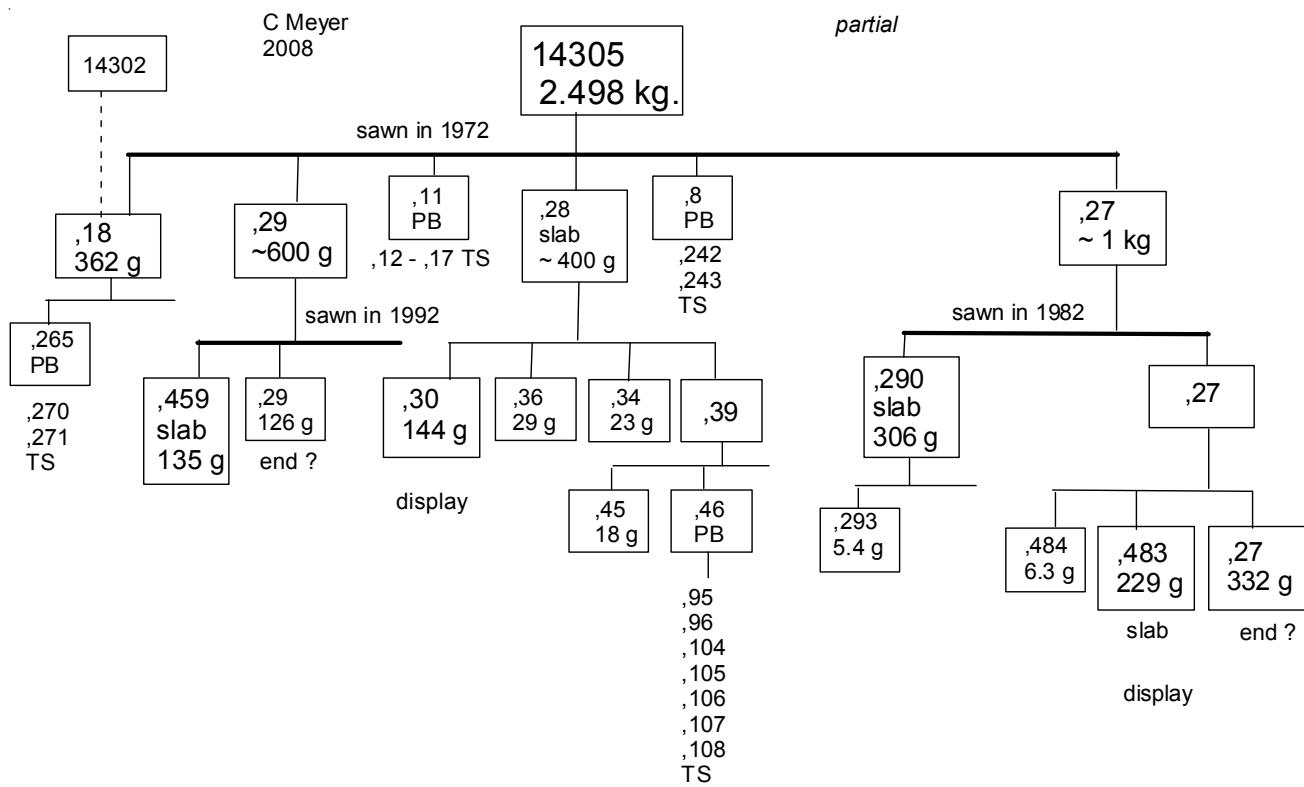


Figure 19: Exploded parts diagram for first slab.

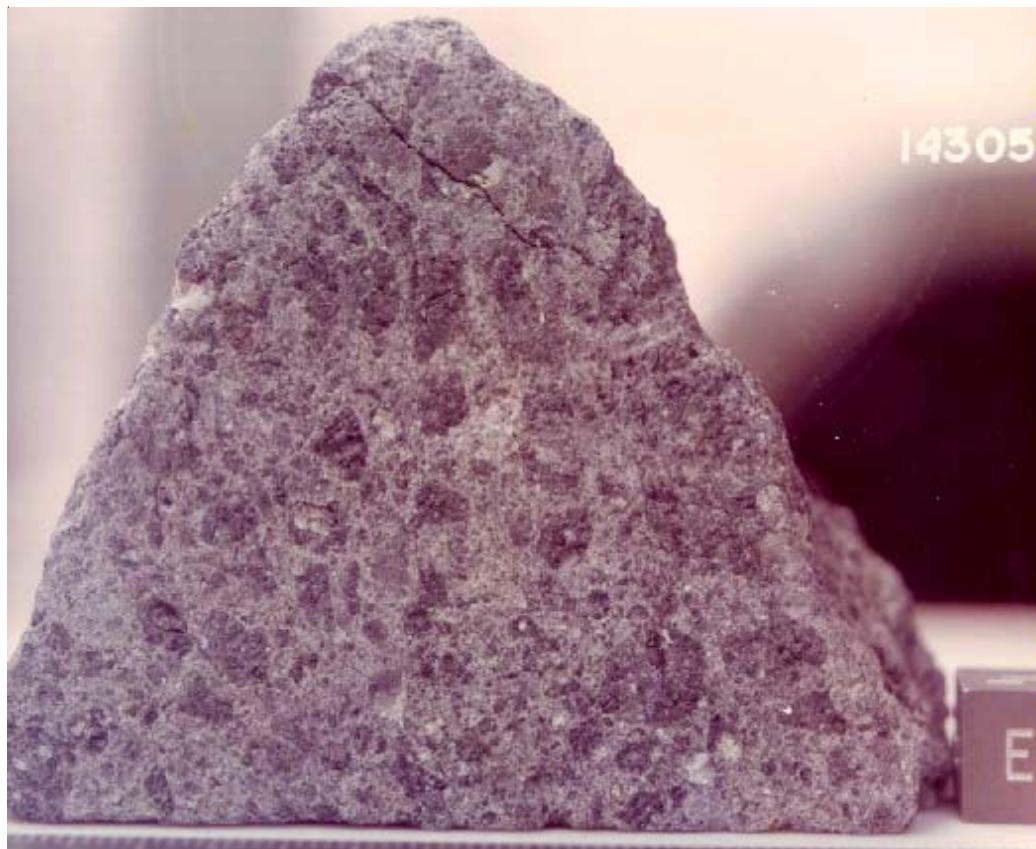


Figure 20: Sawn interior surface of 14305,30 (display piece) illustrating dark microbreccia clasts in light matrix. Cube is 1 inch. NASA photo #S77-21471.

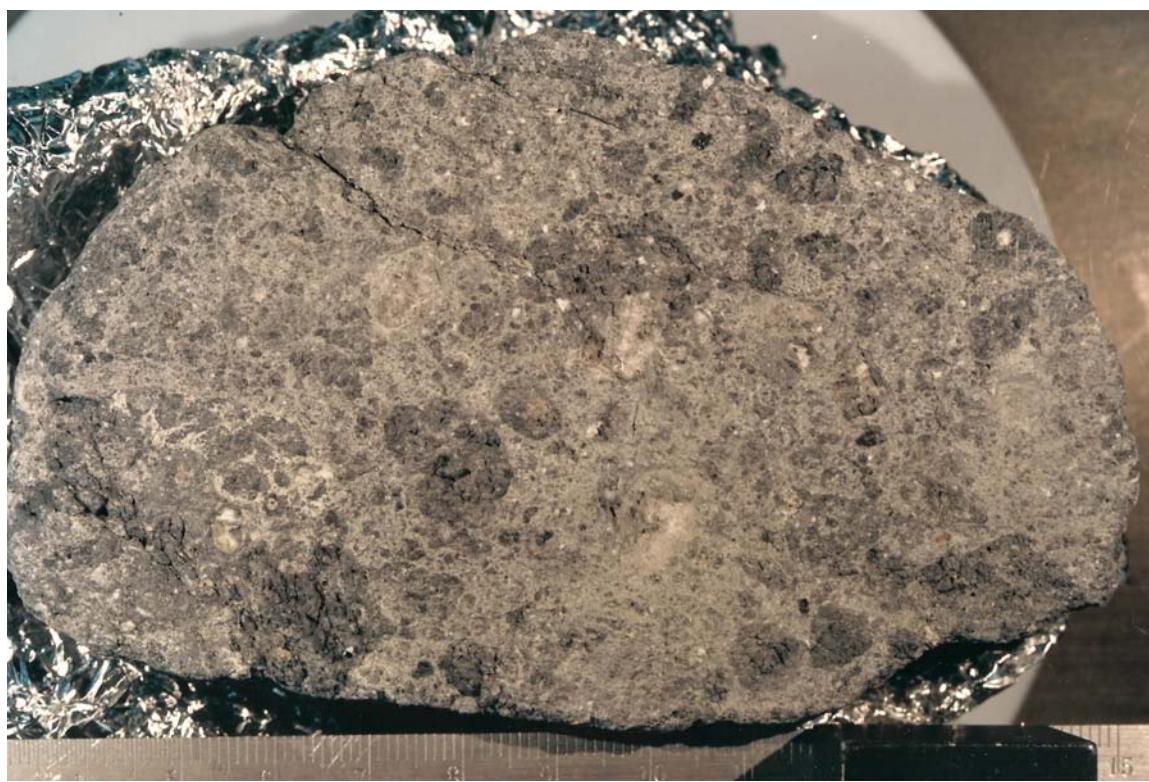


Figure 21: Interior surface of 14305,459. Scale is in cm. no number

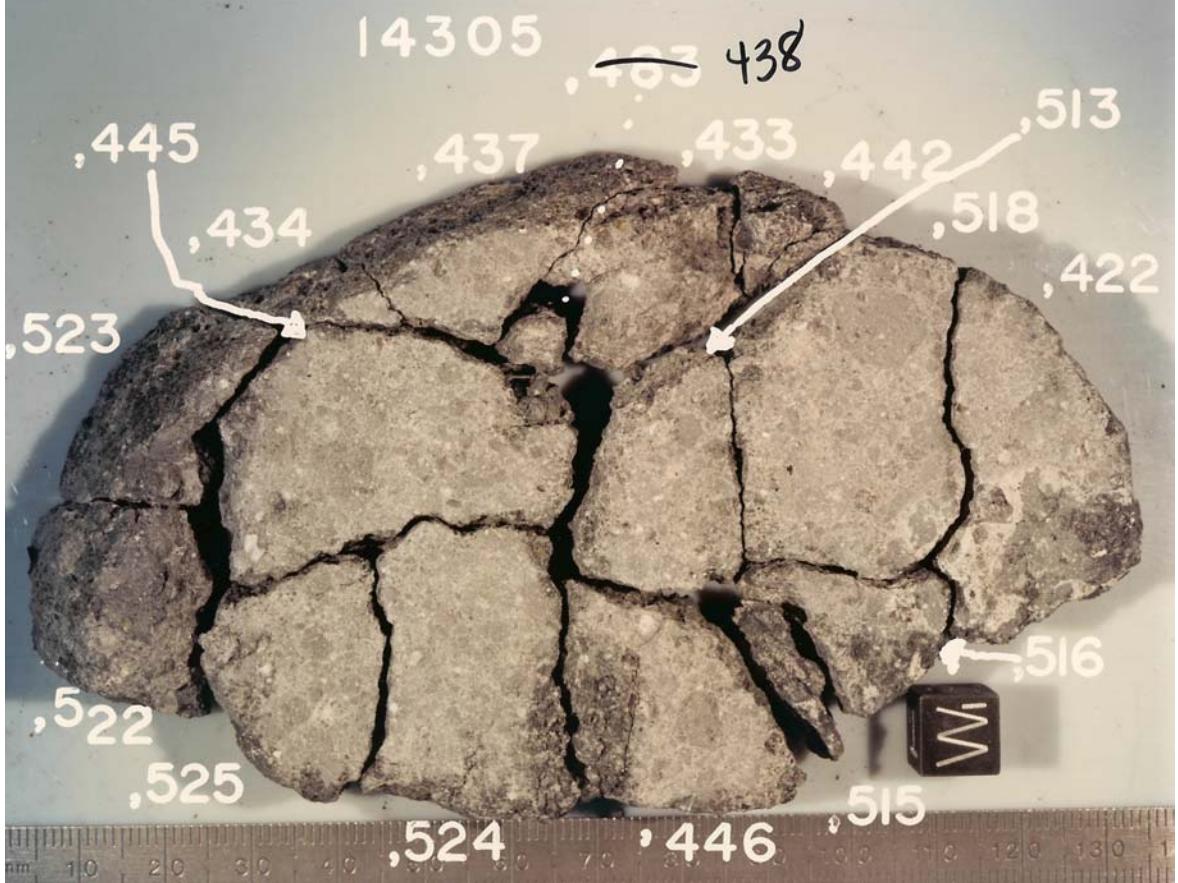
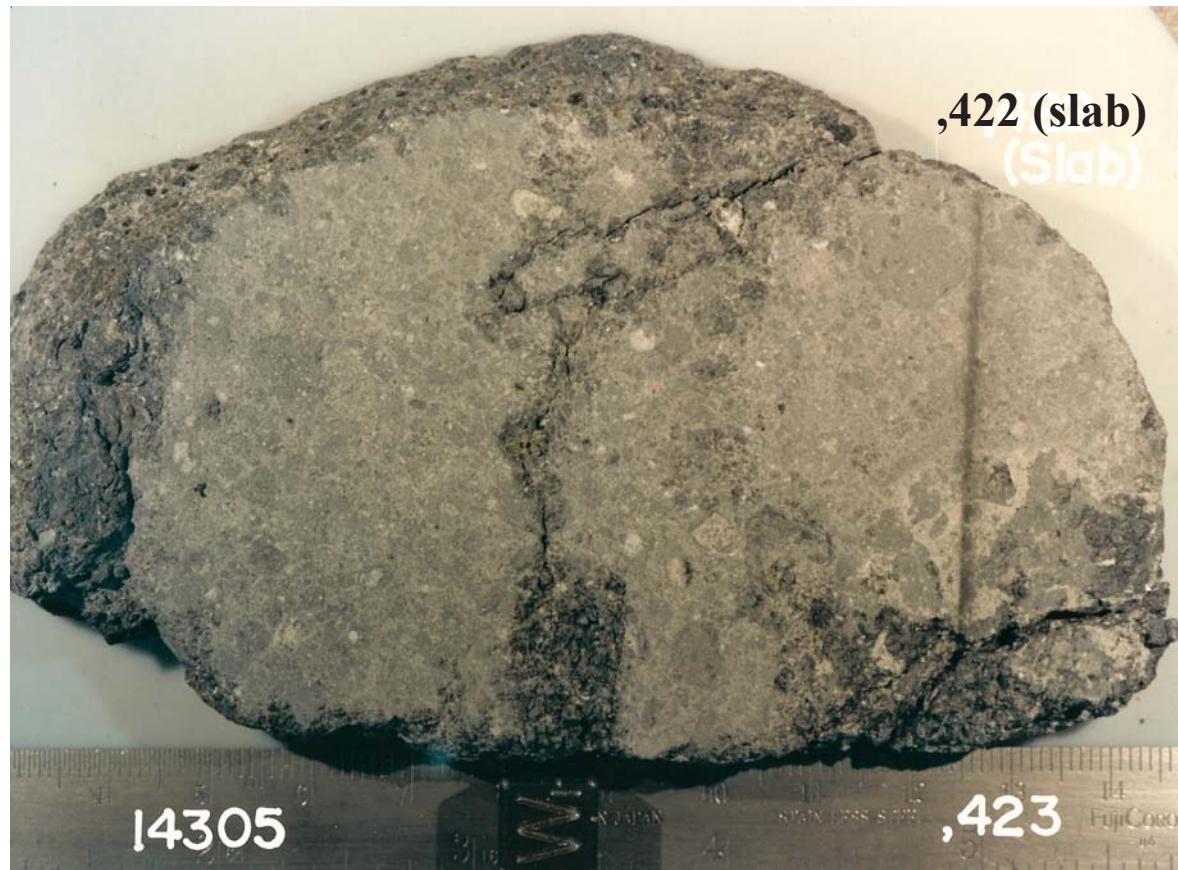


Figure 22: Photos of slab 14305,422 before and after processing. Cube is 1 cm. NASA S85-38255 and S87-29851.

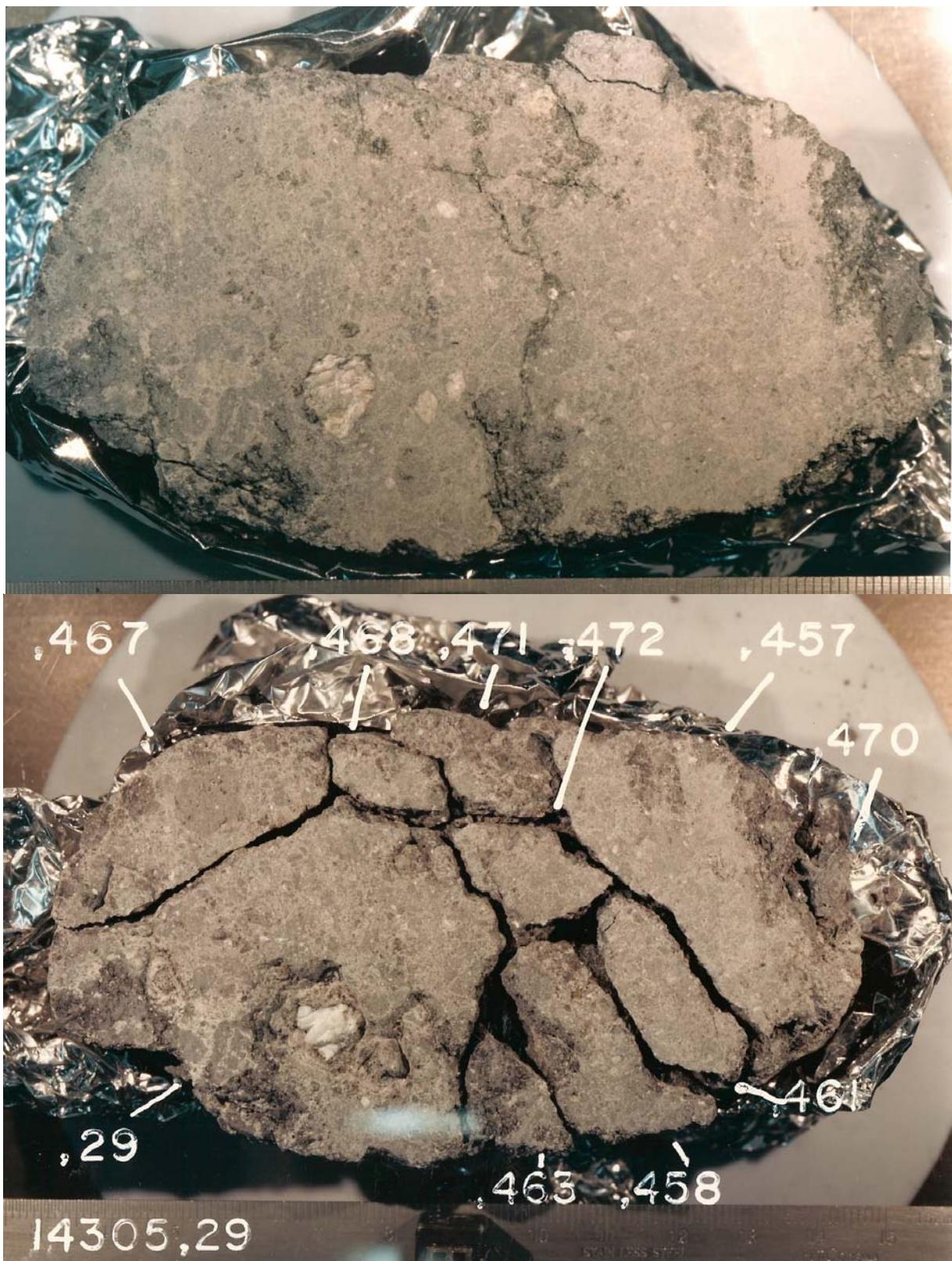


Figure 23: Photo of end piece 14305,29 showing processing chips. Scale is 1 cm. NASA S85-38259 and 39002.

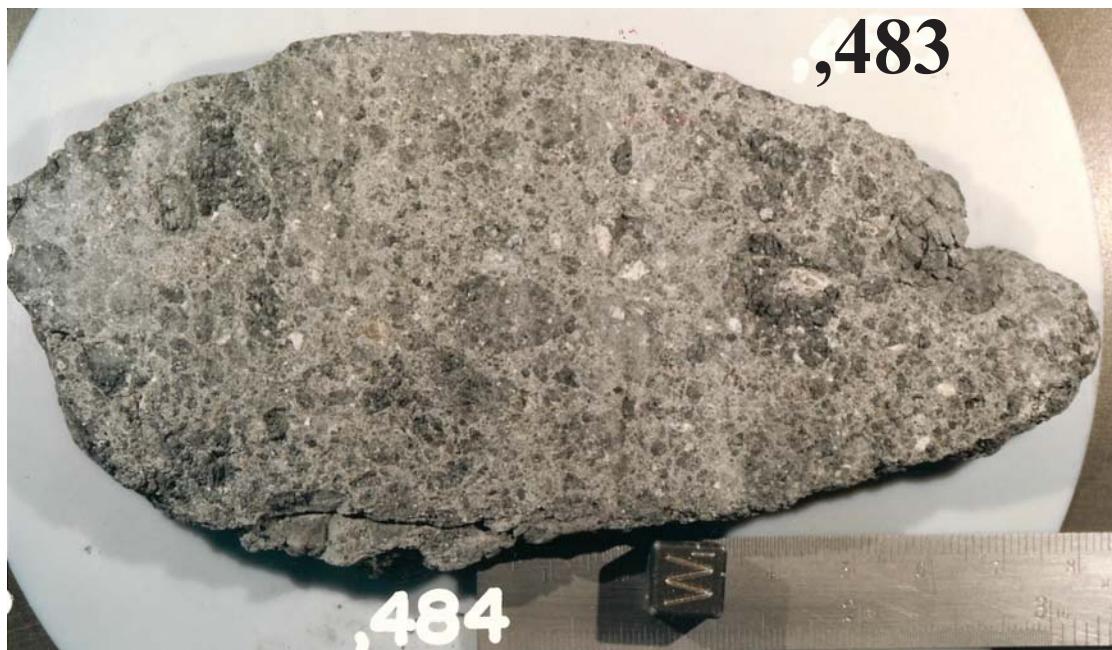
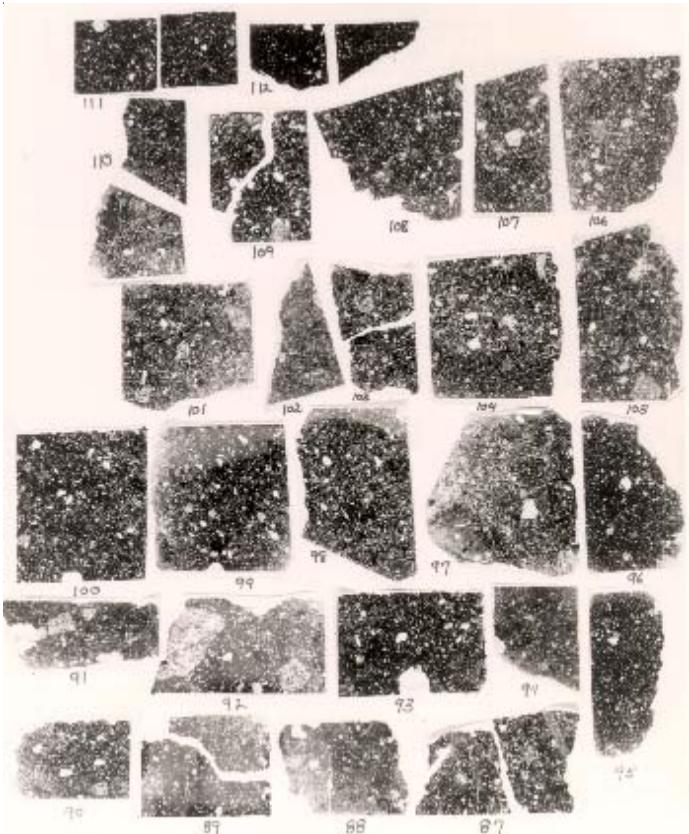


Figure 24: Sawn surface of slab 14305,483. Cube is 1 cm. NASA photo S86-25842.



Figure 25: Sawn surface of 14305,27 (end piece). Scale bar in cm. NASA S86-25845.



*Figure 26: Numerous thin section of 14305.  
Note the large basalt clast in ,92. Each section is  
about 2 cm across.*

#### Partial List of Photo #s for 14305

S71-31391-9	color mug shots
S75-33045	,29 sawn surface
S85-39002	,29 processing
S77-21470-3	,30
S77-21474-5	,18
S77-21476	,27 sawn surface
S83-34589	
S85-38259	,29 end
S85-38255	,422 slab
S86-25842	,483 slab
S86-25845	,27 end
S86-30442	,290 slab
S87-29851	,438 slab

#### Clast correlations for 14305

	from	size mm	analyses	duplicate	thin section	age	authors
c1	,18	7 x 5	,264		,268		Warren 80
c2		6 x 5	,283	,303			Warren 83
c3	,27	7 x 6	,279	,301			Warren 83
w1	,29	8 x 3			,394		Shervais 84
w3	,29	2 x 1.5			,396		Shervais 84
w6	,27	9 x 5	,358	,317	,377 ,347		Shervais 83, 84
w7	,27	6 x 4	,361	,322	,362 ,320		Shervais 83, 84
,400	,51	10 x 5	,400		,412		Shervais 84
M1	,29E	4 x 3	,389		,405		Shervais 84
,122	,92				,92		Taylor 83
M1	,18		,304	,370	,343 ,344	,371	Shervais 85, Shih 85
M2	,18		,385	,384	,383		Shervais 85
M3	,18		,382		,388		Shervais 85
M11	,27		,373		,380		Shervais 85
,390			,390		,393		Shervais 85
,367			,367				Shervais 85

## References for 14305

- Ari T., Takeda H., Miyamoto M. and Kojima H. (2006) Apollo 14 oldest mare basalt revisited: Possible petrogenetic connection between mg gabbronorite and VHK basalt (abs#2387). *Lunar Planet. Sci.* XXXVII Lunar Planetary Institute, Houston.
- Ari T., Kaiden H., Misawa K. and Kojima H. (2006) Ion microprobe study of Apollo 14 oldest basalt (abs). *Antarctic Meteorites* XXX, 3-4.
- Carlson I.C. and Walton W.J.A. (1978) **Apollo 14 Rock Samples**. Curators Office. JSC 14240
- Chao E.C.T., Minkin J.A. and Best J.B. (1972) Apollo 14 breccias: General characteristics and classification. *Proc. 3<sup>rd</sup> Lunar Sci. Conf.* 645-659.
- Cisowski S.M., Collinson D.W., Runcorn S.K., Stephenson A. and Fuller M. (1983) A review of lunar paleointensity data and implications for the origin of lunar magnetism. *Proc. 13<sup>th</sup> Lunar Planet. Sci. Conf.* A691-A704.
- Collinson D.W., Runcorn S.K., Stephenson A. and Manson A.J. (1972) Magnetic properties of Apollo 14 rocks and fines. *Proc. 3<sup>rd</sup> Lunar Sci. Conf.* 2343-2361.
- Dence M.R. and Plant A.G. (1972) Analysis of Fra Mauro samples and the origin of the Imbrium Basin. *Proc. 3<sup>rd</sup> Lunar Sci. Conf.* 379-399.
- Dran J.C., Duraud J.P., Maurette M., Durrieu L., Jouret C. and Legressus C. (1972) Track metamorphism in extraterrestrial breccias. *Proc. 3<sup>rd</sup> Lunar Sci. Conf.* 2883-2903.
- Eugster O., Eberhardt P., Geiss J., Grogler N., Jungck M., Meier F., Morgell M. and Niederer F. (1984a) Cosmic ray exposure histories of Apollo 14, Apollo 15 and Apollo 16 rocks. *Proc. 14<sup>th</sup> Lunar Planet. Sci. Conf.* in *J. Geophys. Res.* 89, B498-B512.
- Fields P.R., Diamond H., Metta D.N., Rokop D.J. and Stevens C.M. (1972)  $^{237}\text{Np}$ ,  $^{236}\text{U}$  and other actinides on the moon. *Proc. 3<sup>rd</sup> Lunar Sci. Conf.* 1637-1644.
- Gault D.E., Horz F. and Hartung J.B. (1972) Effects of microcratering on the lunar surface. *Proc. 3<sup>rd</sup> Lunar Sci. Conf.* 2713-2734.
- Goel P.S. and Kothari B.K. (1972) Total nitrogen contents of some Apollo 14 lunar samples by neutron activation analysis. *Proc. 3<sup>rd</sup> Lunar Sci. Conf.* 2041-2050.
- Herpers U., Herr W., Kulus H., Michel R., Thiel K and Woelfle R. (1973) Manganese-53 profile, particle track studies and rhenium-187 isotopic anomaly of breccia 14305. *Proc. 4<sup>th</sup> Lunar Sci. Conf.* 2157-2169.
- Herr W., Herpers U. and Woelfle R. (1972) Study on the cosmic ray produced long-lived Mn53 in Apollo 14 samples. *Proc. 3<sup>rd</sup> Lunar Sci. Conf.* 1763-1769.
- Hughes T.C., Keays R.R. and Lovering J.F. (1973) Siderophile and volatile trace elements in Apollo 14, 15 and 16 rocks and fines: Evidence for extralunar component and Tl-, Au- and Ag-enriched rocks in the ancient lunar crust. (abs) *LS IV*, 400-402.
- Hunter R.H. and Taylor L.A. (1983) The magma ocean from the Fra Mauro shoreline: An overview of the Apollo 14 crust. *Proc. 13<sup>th</sup> Lunar Planet. Sci. Conf.* in *J. Geophys. Res.* 88, A591-A602.
- Jovanovic S. and Reed G.W. (1976a) Chemical fractionation of Ru and Os in the Moon. *Proc. 7<sup>th</sup> Lunar Sci. Conf.* 3437-3446.
- Jovanovic S. and Reed G.W. (1977) Trace element geochemistry and the early lunar differentiation. *Proc. 8<sup>th</sup> Lunar Sci. Conf.* 623-632.
- Juan V.C., Chen J.C., Huang C.K., Chen P.Y. and Wang Lee C.M. (1972) Petrology and chemistry of some Apollo 14 lunar samples. *Proc. 3<sup>rd</sup> Lunar Sci. Conf.* 687-705.
- Keith J.E., Clark R.S. and Richardson K.A. (1972) Gamma-ray measurements of Apollo 12, 14 and 15 lunar samples. *Proc. 3<sup>rd</sup> Lunar Sci. Conf.* 1671-1680.
- Klein C. and Drake J.C. (1972) Mineralogy, petrology and surface features of some fragmental material from the Fra Mauro site. *Proc. 3<sup>rd</sup> Lunar Sci. Conf.* 1095-1113.
- Lindstrom M.M., Knapp S.A., Shervais J.W. and Taylor L.A. (1984) Magnesian anorthosites and associated troctolites and dunite in Apollo 14 breccias. *Proc. 15<sup>th</sup> Lunar Planet. Sci. Conf.* in *J. Geophys. Res.* 89, C41-C49.
- Lovering J.F., Wark D.A., Geadow A.J.W. and Sewell D.K.B. (1972) Uranium and potassium fractionation in pre-Imbrian lunar cratral rocks. *Proc. 3<sup>rd</sup> Lunar Sci. Conf.* 281-294.
- LSPET (1971) Preliminary examination of lunar samples from Apollo 14. *Science* 173, 681-693.
- Meyer C., Galindo C. and Yang V. (1991) Lunar Zircon (abs). *Lunar Planet. Sci.* XXII, 895-896. Lunar Planetary Institute, Houston.
- Meyer C., Williams I.S. and Compston W. (1989)  $^{207}\text{Pb}/^{206}\text{Pb}$  ages of zircon-containing rock fragments indicate continuous magmatism in the lunar crust from 4350 to 3900

- million years (abs). *Lunar Planet. Sci.* XX, 691-692. Lunar Planetary Institute, Houston.
- Meyer C., Williams I.S. and Compston W. (1989) Zircon-containing rock fragments within Apollo 14 breccias indicate serial magmatism from 4350 to 4000 million years (abs). In *Workshop on Moon in Transition: Apollo 14, KREEP, and evolved lunar rocks*. LPI Tech Rpt. 89-03, 75-78. Lunar Planetary Institute, Houston.
- Moore C.B., Lewis C.F., Cripe J., Delles F.M., Kelly W.R. and Gibson E.K. (1972) Total carbon, nitrogen and sulfur in Apollo 14 lunar samples. *Proc. 3<sup>rd</sup> Lunar Sci. Conf.* 2051-2058.
- Morrison D.A., McKay D.S., Heiken G.H. and Moore H.J. (1972) Microcraters on lunar rocks. *Proc. 3<sup>rd</sup> Lunar Sci. Conf.* 2767-2791.
- Nemchin A.A., Pidgeon R.T., Whitehouse M.J., Vaughan J.P. and Meyer C. (2008) SIMS study of zircons from Apollo 14 and 17 breccias: Implications for the evolution of lunar KREEP. *Geochim. Cosmochim. Acta* 72, 668-689.
- Palme H., Baddehausen H., Blum K., Cendales M., Dreibus G., Hofmeister H., Kmse H., Palme C., Spettel B., Vilcsek E. and Wanke H. (1978) New data on lunar samples and achondrites and a comparison of the least fractionated samples from the earth, the moon, and the eucrite parent body. *Proc. 9<sup>th</sup> Lunar Planet. Sci. Conf.* 25-57.
- Philpotts J.A., Schnetzler C.C., Nava D.F., Bottino M.L., Fullagar P.D., Thomas H.H., Schumann S. and Kouns C.W. (1972) Apollo 14: Some geochemical aspects. *Proc. 3<sup>rd</sup> Lunar Sci. Conf.* 1293-1305.
- Reed G.W., Jovanovic S. and Fuchs L. (1973) Trace element relations between Apollo 14 and 15 and other lunar samples, and the implications of a moon-wide Cl-KREE coherence and Pt-metal coherence. *Proc. 3<sup>rd</sup> Lunar Sci. Conf.* 1989-2001.
- Shervais J.W. and Taylor L.A. (1983) Breccia Guidebook No. 5 14305. JSC 19267 Curators' office.
- Shervais J.W., Taylor L.A. and Laul J.C. (1983) Ancient crustal components in the Fra Mauro breccias. *Proc. 14<sup>th</sup> Lunar Planet. Sci. Conf.* in *J. Geophys. Res.* 88, 77-92.
- Shervais J.W., Taylor L.A., Laul J.C. and Smith M.R.. (1985a) Pristine highland clasts in consortium breccia 14305: Petrology and Geochemistry. *Proc. 15<sup>th</sup> Lunar Planet. Sci. Conf.* in *J. Geophys. Res.* 89, C25-40.
- Shervais J.W., Taylor L.A., Laul J.C., Shih C.-Y. and Nyquist L.E. (1985) Very high potassium (VHK) basalt: Complications in lunar mare petrogenesis. *Proc. 16<sup>th</sup> Lunar Planet. Sci. Conf.* in *J. Geophys. Res.* 90, D3-D18.
- Shih C.-Y., Nyquist L.E., Bogard D.D., Bansal B.M., Wiesmann H., Johnson P., Shervais J.W. and Taylor L.A. (1986) Geochronology and petrogenesis of Apollo 14 very high potassium mare basalts. *Proc. Lunar Planet. Sci. Conf.* 16<sup>th</sup> in *J. Geophys. Res.* 91, D214-D228.
- Simonds C.H., Phinney W.C., Warner J.L., McGee P.E., Geeslin J., Brown R.W. and Rhodes J.M. (1977) Apollo 14 revisited, or breccias aren't so bad after all. *Proc. 8<sup>th</sup> Lunar Sci. Conf.* 1869-1893.
- Sutton R.L., Hait M.H. and Swann G.A. (1972) Geology of the Apollo 14 landing site. *Proc. 3<sup>rd</sup> Lunar Sci. Conf.* 27-38.
- Swann G.A., Trask N.J., Hait M.H. and Sutton R.L. (1971a) Geologic setting of the Apollo 14 samples. *Science* 173, 716-719.
- Swann G.A., Bailey N.G., Batson R.M., Eggleton R.E., Hait M.H., Holt H.E., Larson K.B., Reed V.S., Schaber G.G., Sutton R.L., Trask N.J., Ulrich G.E. and Wilshire H.G. (1977) Geology of the Apollo 14 landing site in the Fra Mauro Highlands. U.S.G.S. Prof. Paper 880.
- Swann G.A., Bailey N.G., Batson R.M., Eggleton R.E., Hait M.H., Holt H.E., Larson K.B., McEwen M.C., Mitchell E.D., Schaber G.G., Schafer J.P., Shepard A.B., Sutton R.L., Trask N.J., Ulrich G.E., Wilshire H.G. and Wolfe E.W. (1972) 3. Preliminary Geologic Investigation of the Apollo 14 landing site. In *Apollo 14 Preliminary Science Rpt.* NASA SP-272. pages 39-85.
- Taylor D.J., McKeegan K.D., Harrison T.M. and McCulloch M. (2007)  $^{176}\text{Lu}/^{176}\text{Hf}$  in lunar zircons: Identification of an early enriched reservoir on the moon (abs#2130). *Lunar Planet. Sci. XXXVIII* Lunar Planetary Institute, Houston.
- Taylor D.J., McKeegan K.D. and Harrison T.M (2009) Lu-Hf zircon evidence for rapid lunar differentiation. *Earth Planet. Sci. Lett.* 279, 157-164.
- Taylor L.A., Shervais J.W., Hunter R.H., Shih C.-Y., Nyquist L.E., Bansal B.M., Wooden J. and Laul J.C. (1983) Pre-4.2 AE mare-basalt volcanism in the lunar highlands. *Earth Planet. Sci. Lett.* 66, 33-47.
- Tweddell D., Feight S., Carlson I. and Meyer C. (1978) Lithologic maps of selected Apollo 14 breccia samples. Curators Office. JSC 13842
- Wanke H., Baddehausen H., Balacecu A., Teschke F., Spettel B., Dreibus G., Palme H., Quijano-Rico M., Kruse H., Wlotzka F. and Begemann F. (1972) Multielement

analysis of lunar samples and some implications of the results. *Proc. 3<sup>rd</sup> Lunar Sci. Conf.* 1251-1268.

Warner J.L. (1972) Metamorphism of Apollo 14 breccias. *Proc. 3<sup>rd</sup> Lunar Sci. Conf.* 623-643.

Warren P.H. and Wasson J.T. (1980a) Further foraging of pristine nonmare rocks: Correlations between geochemistry and longitude. *Proc. 11<sup>th</sup> Lunar Planet. Sci. Conf.* 431-470.

Warren P.H., Taylor G.J., Keil K., Kallemeyn G.W., Shirley D. and Wasson J.T. (1983d) Seventh foray: Whitlockite-rich lithologies, a diopside-bearing troctolitic anorthosite, ferroan anorthosite and KREEP. *Proc. 14<sup>th</sup> Lunar Planet. Sci. Conf.* in *J. Geophys. Res.* 88, B151-B164.

Warren P.H. (1993) A concise compilation of petrologic information on possibly pristine nonmare Moon rocks. *Am. Mineral.* 78, 360-376.

Wiik H.B., Maxwell J.A. and Bouvier J.-L. (1973) Chemical composition of some Apollo 14 lunar samples. *Earth Planet. Sci. Lett.* 17, 365-368.

Willis J.P., Erlank A.J., Gurney J.J., Theil R.H. and Ahrens L.H. (1972) Major, minor, and trace element data for some Apollo 11, 12, 14 and 15 samples. *Proc. 3<sup>rd</sup> Lunar Sci. Conf.* 1269-1273.

Williams R.J. (1972) The lithification of metamorphism of lunar breccias. *Earth Planet. Sci. Lett.* 16, 250-256.

Wilshire H.G. and Jackson E.D. (1972) Petrology and stratigraphy of the Fra Mauro Formation at the Apollo 14 site. U.S. Geol. Survey Prof. Paper 785.

Wlotzka F., Jagoutz E., Spettle B., Baddehausen H., Balaceșcu A. and Wanke H. (1972) On lunar metallic particles and their contribution to the trace element content of Apollo 14 and 15 soils. *Proc. 3<sup>rd</sup> Lunar Sci. Conf.* 1077-1084.

Wlotzka F., Spettel B. and Wanke H. (1973) On the composition of metal from Apollo 16 fines and the meteoritic component. *Proc. 4<sup>th</sup> Lunar Sci. Conf.* 1483-1491.