## **15007 - 15008** Double Drive Tube 1278.4 grams

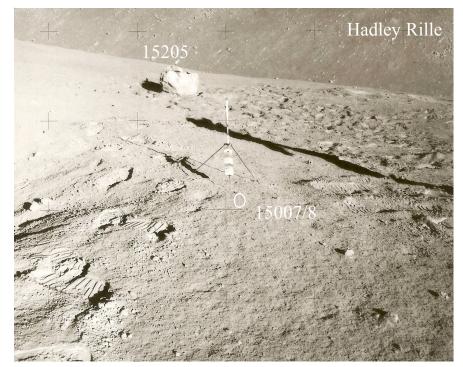


Figure 1: Location of double drive tube 15008/7, station 2, Apollo 15. AS15-85-11443.

## **Introduction**

The total depth of the 15007/8 double drive tube is 56.6 cm. It was collected from the rim of a 10 meter shallow crater at the highest point on the Apennine Front at station 2, Apollo 15 (figure 1) close to the location of reference soil (15220), and other soil samples including 15091, 15101, 15201, 15211, 15231 collected nearby (see section on 15100). Station 2 was on the outer flank of St. George Crater which punched deep into the Apennine Front material.

15008 is the top segment of the double drive tube, and 15007 is the bottom segment. 15007 was found to contain immature, glass-rich and more aluminous soil at about 55 cm depth.

The 15008/7 double drive tube was the first taken with the new 4.13 cm diameter assembly (Mitchell et al. 1972). The soil material did not fall out and was less disturbed, because it did not have to "flow" past a lip as was the case in previous missions. Papike et al. (1982) included 15007/8 in their set of reference soils.

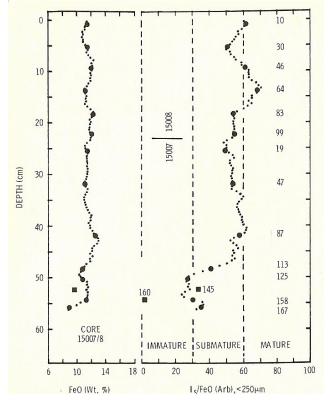
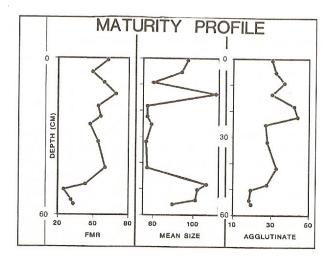


Figure 2: Maturity of soil as function of depth in double drive tume 15008 - 15007 (Bogard et al. 1982).



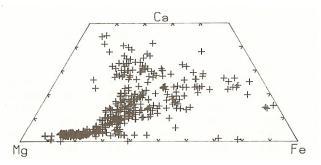
*Figure 3: Maturity, grain size and agglutinate content as function of depth in 15008/7 drive tube (Basu et al. 1982).* 

## **Petrography**

The maturity index (I<sub>s</sub>/FeO) of 15008/7 varies as a function of depth (figure 2). There are more agglutinates in the top of the core than the bottom and the average grain size is also variable (figure 3). Graf (1992) reported the grain size studies (figure 10). Nagle (1980, 1981a,b) gives a description of the core during dissection and Basu et al. (1988) give the detailed model mineralogy (see table). The only notable result is that there was a lot (26 %) of green glass in the 55 cm horizon.

During dissection, Nagle (1981) separated several apparently pristine rock fragments from the bottom of 15007. Portions of these were studied by Warren et al. (1983). Some of these separated particles were KREEP basalt, while others were "anorthosite".

Mutispectral imaging by Pieters et al. (1981) revealed a feldspathic fragment-rich zone with a chaotic fabric that occurs between 10 and 18 cm depth. Figure 13 is



*Figure 4: Composition of individual pyroxene grains in 15008/7 (Basu et al. 1982).* 

a low-magnification photomicrograph of thin section 15008,6023 illustrating fragments of breccias incorporated into the core at about 18 cm depth.

The composition of individual pyroxene grains shows that KREEP basalt is a major component of the Apennine Front (figure 4).

### <u>Chemistry</u>

The bottom of the core (15007) is more aluminous and less mafic than the surface soils (Korotev 1987). The composition of the top (15008) has not been reported, but one can assume it is the same as the surface soil (15201).

## Cosmogenic isotopes and exposure ages

Nishiizumi et al. (1989), Jull et al. (1998) and Fruchter et al. (1982) studied <sup>36</sup>Cl, <sup>14</sup>C and <sup>26</sup>Al (respectively) as function of depth (figures 6 - 8).

## **Other Studies**

Bogard et al. (1982) determined the rare gas content and isotopic ratios as function of depth. They concluded that the top 18 cm was deposited by the adjacent 10 meter crater, while the material 18 to 49 cm was the preexisting regolith, mass-wasted off of

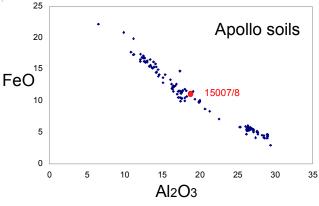
Mineralogical Mode 15007/8 (f	from Basu et al.	1988)
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Summar	ry (90 – 1	50 micron	)	-		,		
Depth	(cm)	9.3	25	32	42	50	55	Ave
Agglutin	nate	36.9%	43.9	27.9	32.7	19.3	18.5	29.9
Breccia		11.3	13.8	12.2	11.1	13.8	9.1	12.4
Lithic		8	6.3	8	6.6	7	8.7	7.8
Glass		12.5	14.2	15.1	12.7	28.6	33.8	15.2
Mineral	s	31.3	38.5	36.3	36.8	30.5	29.5	34.6
	plag	12.5	18.5	15.4	18.4	18.1	15.4	16.7
	pyx.	17.2	18.8	19.4	16.5	11.5	13.5	15.9
	ol.	1.3	0.9	0.9	1.9	0.3	0.6	1.4
Note S	ee Rasu I	(table 2) fa	r detail					

Note: See Basu (table 2) for detail.

Table 1	. Chem	ical con	npositi	on	of 150	008/7.	
	15007		15201		15007 (	particles)	)
reference	Korotev87				Warren8		
<i>weight</i> SiO2 %	55-57cm ave. 3	42-43cm ave. 3	surface		,290 49.4	,296 43.9	$(\mathbf{a})$
TiO2 //	1.24	ave. 5 1.5	1.3	(a)		0.03	(a) (a)
AI2O3	20	18.4	17.1	(a)	19.8	35.1	(a)
FeO	10	11.6	11.5		7.85	0.22	(a)
MnO	0.144	0.162	0.16	(a)	0.11	0.005	(a)
MgO	10.3	10.7	10.7	1.1	7.63	0.33	(a)
CaO	11.4	10.6	10.7	(a)		19.9	(a)
Na2O K2O	0.467	0.423	0.42	(a)	0.83 0.45	0.31 0.07	(a) (a)
P2O5					0.40	0.07	(a)
S %							
sum							
0	40 <del>-</del>		<b>00</b> 4		45.0		
Sc ppm V	18.5	22	22.1		15.2	0.84	(a)
v Cr	65 2013	85 2263	76 2220	(a) (a)	1810	75	(a)
Co	30.2	38.2	37.7	(a)	17	0.7	(a)
Ni	162	238	225		28	4	(a)
Cu							
Zn					4.7	0.4	(a)
Ga					6	3.1	(a)
Ge ppb As					340	106	(a)
Se							
Rb					16		
Sr	145	113	140	(a)	230	168	(a)
Y							
Zr	363	277	320	(a)	870	10	(a)
Nb Mo							
Ru							
Rh							
Pd ppb							
Ag ppb							
Cd ppb							
In ppb							
Sn ppb Sb ppb							
Te ppb							
Cs ppm	0.25	0.2	0.24	(a)			
Ва	256	226	218	(a)	660	30	(a)
La	22.9	20.6	20.4	(a)	53	0.3	(a)
Ce	59	54.3	53	(a)	139	0.66	(a)
Pr Nd	34.3	32.7	30	(a)	88	0.46	(a)
Sm	10.8	9.8	9.48		23.2	0.11	(a)
Eu	1.36	1.29	1.27		2.63	0.74	(a)
Gd							
Tb	2.02	1.91	1.78	(a)	4.7	0.021	(a)
Dy					30	0.107	(a)
Ho Er							
Tm							
Yb	7.67	6.93	6.6	(a)	15.8	0.068	(a)
Lu	1.04	0.96	0.99		2.2	0.008	(a)
Hf	8.7	7.7	7.6		18.6	0.111	(a)
Ta	1	0.93	0.9	(a)	2.04	0.006	(a)
W ppb Re ppb					0.025	0.005	(b)
Os ppb					0.020	0.000	(0)
lr ppb	4.4	7.2	7.6	(a)	0.073	0.026	(b)
Pt ppb				. ,			. ,
Au ppb	1.9	3.3	2.2		0.31	0.041	(b)
Th ppm	3.8	3.4	4.3		9.2		(a)
U ppm	1.27 : (a) INAA,	0.87 (b) RNAA	0.86	(a)	2.4		(a)
leonnque	. (a) IIVAA,						

Table 1. Chemical composition of 15008/7.



*Figure 5: Chemical composition of 15008/7 drive tube compared with other Apollo soils.* 

a) the Apennine Front. However, they ignored the
a) contribution St. George Crater must have made.

## $(a) a) \frac{Processing}{15007}$

a) 15007 and 15008 were returned in ALSRC#1 (sealed).
15008 was dissected, described and first distributed in
a) 1980 (Nagle 1981a,b). Thin sections are available for
the complete core (e.g. figure 13). Pb contamination
a) was found to be severe.

"Sampling was easy (figure 9). The double drive tube was pushed almost to the depth of the lower drive tube (AS15-86-11577), then driven 2" per hammer blow to the full depth possible. The rammer was inserted 6" after sampling. There was no indication of spillage during uncoupling of the drive tubes or placing them into the return container SCB1. Because the drive tube was returned in as SCB, it was subjected to spacecraft cabin atmosphere for approximately 7 days. The core was placed in dry gN, on August 10, 1971. It was Xrayed October 1979. There was slumping approximately 8 cm at the top of 15008. About 5.5 cm of compaction of 15008 occurred during extrusion. Bulk density of 15008 is 1.65 gm/cm<sup>3</sup> and 15007 is 1.7 gm/cm<sup>3</sup>." a)

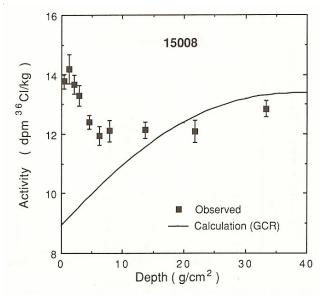
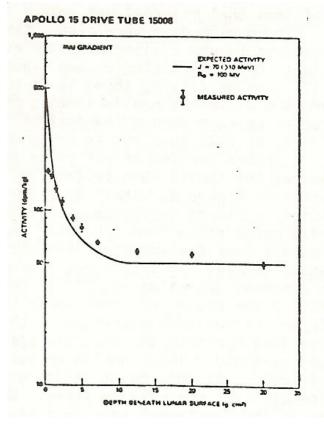


Figure 6: 36Cl as fucnction of depth in 15008 (Nishiizumi et al. 1989).



*Figure 8: 26Al as function of depth in top of 15008 drive tube (Fruchter et al. 1982).* 

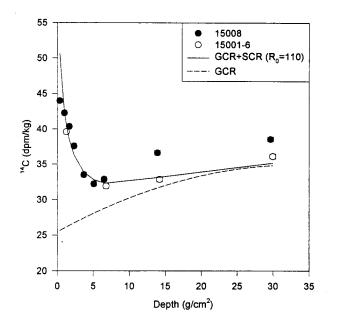


Figure 7: 14C as function of depth in 15008/7 and deep drill 15001 (Jull et al. 1998).

## Transcript core

*CDR Okay* – *next thing on the agenda is a double core.* 

*LMP* Yes. Okay: I'm going to go over and configure for it.

CDR Oh, we've got a good place here. We've got a fairly deep crater: it must be about 10 meters across, and a meter and a half or so deep, and we'll pick the rim of that – There's a fresh impact crater in – in the rim anyway, which like it pulled out some

CDR Is that as far as you can push it, Jim?

*LMP* That's as far as I can push it. I got the picture: go ahead.

*CDR* Okay. It's a – We've got one full core, second core is going in about 2 inches per hammer stroke.

CDR And we've got almost a second core. Got a couple of inches to go, Jim. Doing good. Okay: that's good, man. All the way in, Good show. Okay. Pull it out slowly. Nice. Nice. Easy does it. That's nice. Coming out very clean. Hold it steady. Got a good one. Okay. Come on over this way a little. Cap for it.

*LMP* Give me the cap. I'll put it on, Dave

CDR Okay. Good idea. Okay Rammer went in about 6 inches.

*CC And Dave, we're standing by for a number on the core.* 

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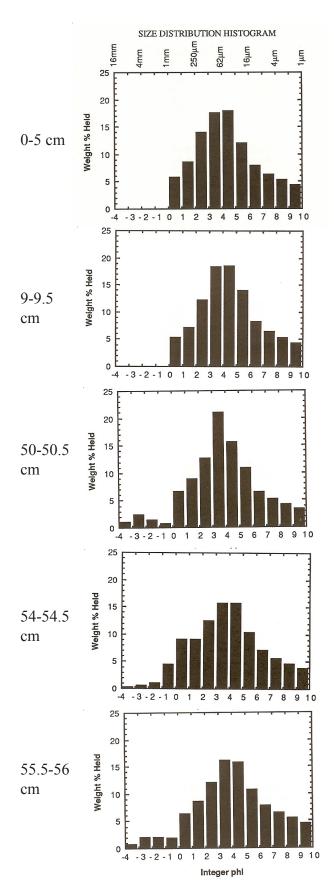
*Figure 9: Doube drive tube 15008/7 during insertion into soft rim of small crater.* 

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*Figure 10: Grain size distribution of several layers of 15008/7 double drive tube (from Graf 1992).* 

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Unit		stance below unar surface (linear cm)	Fine < 1m Sample No.	m Fraction Sample ∀t(gm)	Coarse > 1 Sample No.	Wit(gm)	Special Sample No.	samples Sample Wt(gm)	Sample interval	Sr mple type
	0000	- 0.5	.8	1.566	,9	0.116	,20	0.255	0 - 1 cm	bulk, slumped
	\$ 080 G	- 1.0 -	,10	1,694	.11	0,357				
	0 0 000	- 1.5 -	,12	1.992	,13	0,305	1	Contra de la contr		
	083	- 2.0 -	,14	1.750	,15	0.530	11		No. of Concession, Name	
	000	- 2.5 -	,16	1.770	,17	0.142	11			
9	00000	- 3.0 -	,18	1.614	,19	0.175	11			
۰.	and and	- 3.5 -	,22	2.146	,23	0.210	11			and the second
••	0 9 9 0	- 4.0 -	,24	1,803	,25	0.125	11			
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	a Colu	- 6.0 -	,32 .	2.004	• ,33	0.248				
	1 m Carl	- 6.5 -	. ,34	1,983	,35	0.219	╫───			
8	a C D O		, 36	2.319	,37	0.055				
	45	- 7.0 -	,38	2,145	,39	0.091	<del>il</del>			
8	Bos	- 7.5 -	,40	2.194	,41	0.045	#			
-	690	8.5	,42	2.117	,43	0.04				
7	1 00 00	and a state of a state	,44	2.047	,45	0.185				
Y	Bo Bo	- 9.0 -	.46	2.144	.47	0.369				
	and a	- 9.5 -	,48	2.051	,49	0.068	H			
	കംക	- 10.0 -	,50	2.278	,51	0.096	H			
	0_ 0_0	- 10.5 -	,52	1,490	.53	0.050	H			
	0 000	- 11.0 -	,54	2.929	,55	0.087				
	10 00000	- 12.0 -	,56	2.104	,57	0.365				·
6	6000	- 12.5 -	,58	2.013	,59	0.140				
6	15) 0.00	- 13.0 -	,60	2.070	,61	1.353				
6	Do D	- 13.5 -	,62	1.942	,63	0.113				Territoria and a second statements
		- 14.0 -	,64	1.839	,65	0.128				
		- 14.5 -	,66	2.174	,67	0.335				
		- 15.0 -	,68	2.461	,69	0.196				·····
	(1) (1)	- 15.5 -	,70	2.338	.71	0.255				
		- 16.0 -	,72	1.665	,73	0.263				
_	3-078-00	- 16.5 -	.74	1.896	,75	0,599	,104	0.972	15,5-16.3 c	n large agglutinate
5		- 17.0 -	,76	2.297	.77	0.212	,78	0.021		n light clast
gular	40 @	- 17.5 -	,79	2.361	,80	0.437				
boundary		- 18.0	.81	1.914	.82	0.154				
	9 8	- 18.5 -	,83	2.381	,64	0.120				
5		- 19.0	,85	2.336	,86	0.045				
	co	- 19.5 -	,87	2.067	,88	0.057				
с П	~~~~~~~~~~~~~~~~~~~~~~~~~~~~~~~~~~~~~~	- 20.0 -	,89	2.454	,90	0.161				
	000 00	- 20.5	.91	2.333	,92	0.070				
		- 21.0	.93	2.674	.94	0.100		•		
	<b>*</b> @	- 21.5	,95	2.453	,96	0.056				
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	LOGIC SAMBOLS USED IN						Contraction of the owner of the		Statement of the second	

## Figure 11: First dissection layer of 15008 (top)(Nagle 1981b).

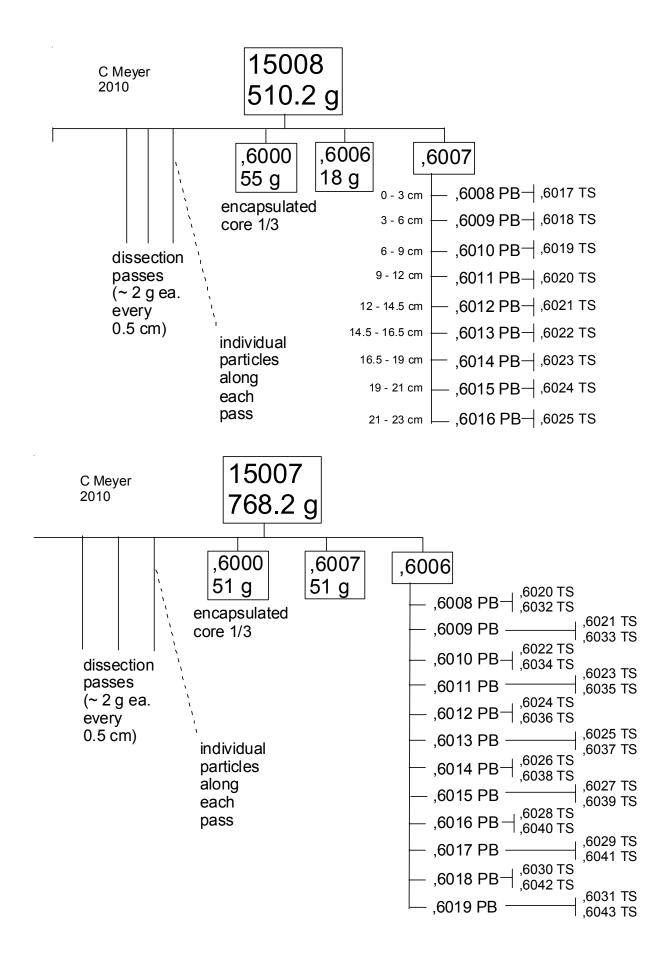
oriented vesicular glass particles or clumps light soil clasts (basalt, ANT) 5.0 Atalline-matrix breccia . 0 orange-brown glass #X fragmented vesicular glass 🔿 cataclastic anorthosite

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Unit Dis	tance below lunar sur	face (linear c	Fine (< imm) fractio n) Sample Sample No. Wt.(um)	and the second of the	
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	0 4000	- 25.1 -	,17 2.080	,18 0.131	of 15007 (top)(Nag
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	Be B	28.6 -	,33 1.827	,34 0.072	
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	19 ° ~ 8	- 32.1 -	.47 2.212	,48 0.270	
	00 00	32.6 -	49 2.303	,50 0.089	
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		- 33.6 -	,55 2.181	,54 0.081 ,56 0.094	
		34.6	,57 2.157	,58 0.142	
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	60 799970482	- 35.6 -	,61 2.202	,62 0.045	0
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	-	36.6 -	,67 2.24B	,66 0.045 ,68 0.030	
		37.6 -	,69 2,158	.70 0.100	
		- 38.1 -	,71 2.238	,72 0.062	
4	00 *	- 38.6 -	.73 2.310 .75 2.409	,74 0.043 ,76 0.076	
		39.1 -	,77 2.105	,76 0.076	
	000	- 40.1 -	,79 1.998	,80 0.095	interest in the states.
		40.6	.81 2.256 ,83 1.997	,82 0.074	
		- 41.1 -	,83 1.997 ,85 2.118	,84 0.079 ,86 0.028	
	880	41.6 -	,87 2.315	.88 0.095	
		- 42.6 -	,89 1.992	,90 0.094	
	CB CB	43.1 -	,91 2.224 ,93 2.193	,92 0.043	
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	0000°	- 48.1 -	111 2.553	,112 0.134	
	1 5	- 48.6 -	,113 2.124	,114 0.209	
transition 49.1-	The Os	- 49.1 +	,115 1.855 ,117 1.658	,116 0.187	
43.0		- 49.6 -	,119 2.447	,118 0.309 * ,120 0.222 *	
	UPPOD	- 50.6	,125 2,421	,126 0.436 *	
a break in a share		- 51,1 -	,128 2.068	,129 0.156 *	
	an the	- 51.6 -	,134 2.231	,135 0.366 *	
2	Bra Caro	- 52.1 -	,137 2.070 ,142 2.144	,138 0,116 *	
	~ 52	- 52.6 -	,147 2.194	,143 0.173 * ,148 0.076 *	An and the set of the
	a cal 3 J	- 53.1	,150 2,166	,148 0.076 * 	
	1 Same	- 54,1	,154 2.274 •	,155 0.120 *	
· · · · · · · · · · · · · · · · · · ·	85	- 54.6	,158 2,175	,159 0.102 *	
transition 55.1-	a dene so	- 55.1	,162 2.111	,163 0.137 *	
STANSTERM 55.6		- 55.6	,165 2.522	,166 0.164	
1	V NO H	- 56.1	,167 1.699	,168 0.116	
56.8	and the second	56.8	,169 2,582	,170 0.524	
Total dept	h after extrusion: 56.	8 cm below lur		onal distinctive material	
LITHOLOGIC S	MBOLS IN COLUMNAR SE	CTION		aced in special samples	
					15008/7
S Ktalli (basa)	ne rock fragments	119	ht soil clasts an ad	By oriented vesicular glass	p. 5 of 8
	t, ANT, norite etc.) ne-matrix breccia		nge-brown glass 🖉 🖵	particles or clumps fragmented vesicular glass	
O Casaci	astic anorthosite	gree	an glass clods . 🥶	soll breccle	

*Figure 12: First dissection layer* of 15007 (top)(Nagle 1981b).



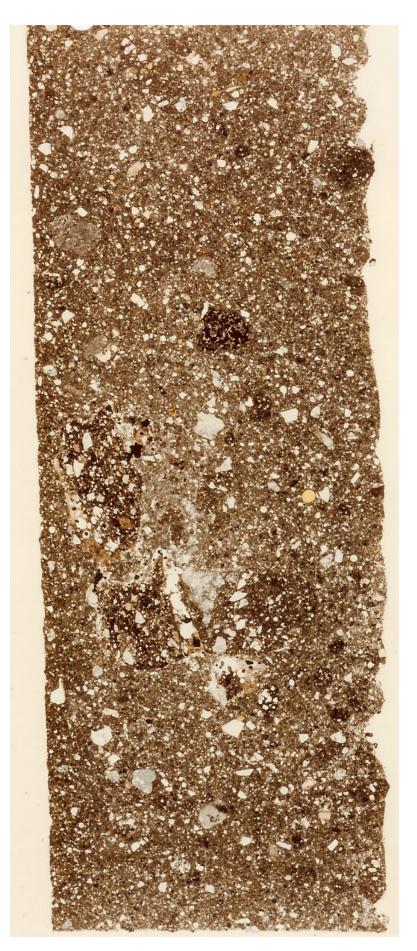
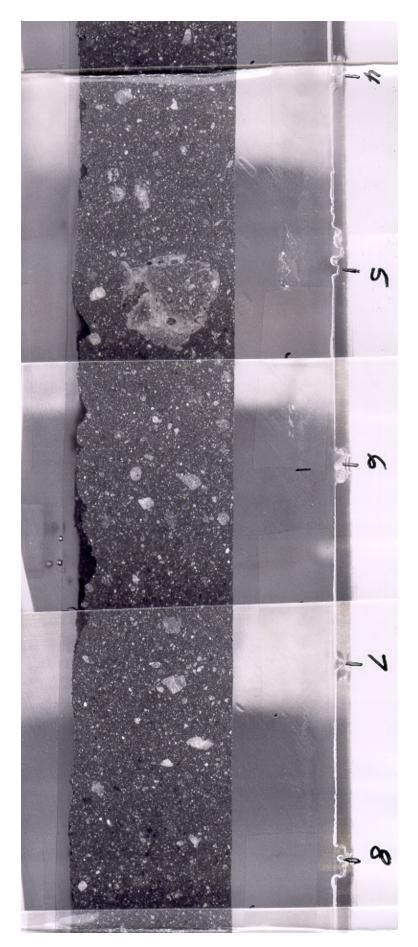


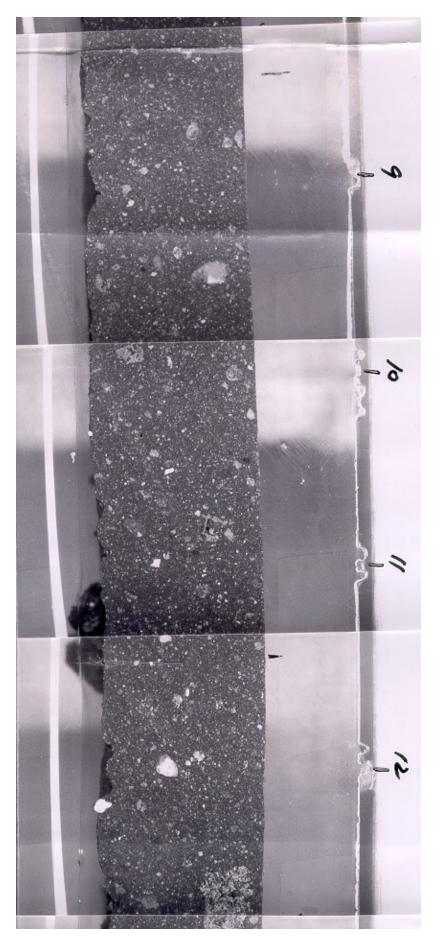
Figure 13: Photomicrograph of thin section 15008,6023 from depth 16.5 - 19 cm. 15008,6006 epoxy encapsulated core

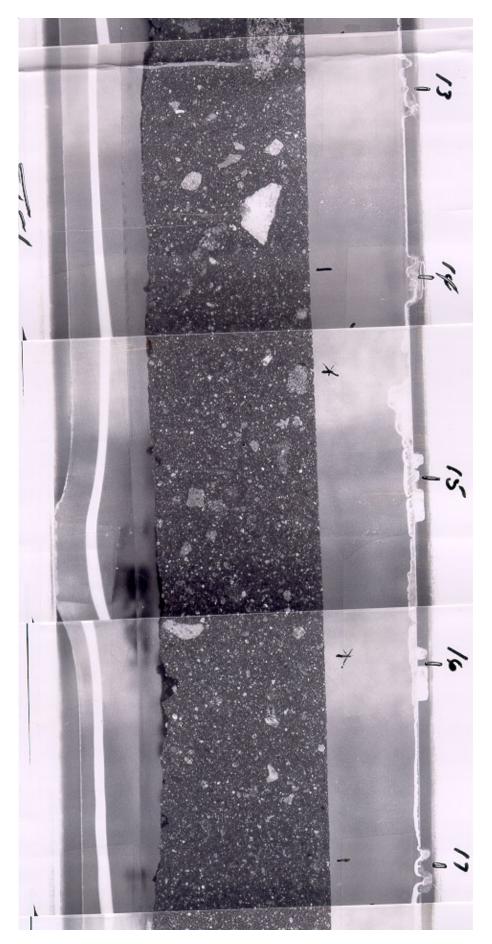
Ica N S

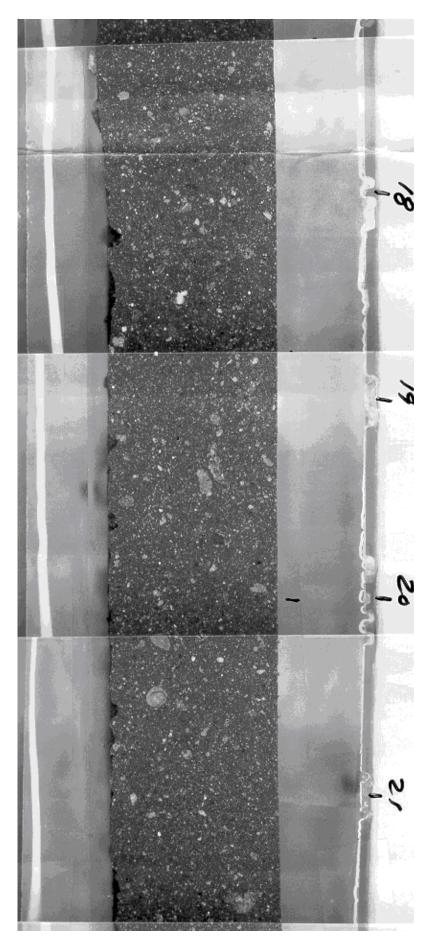
top

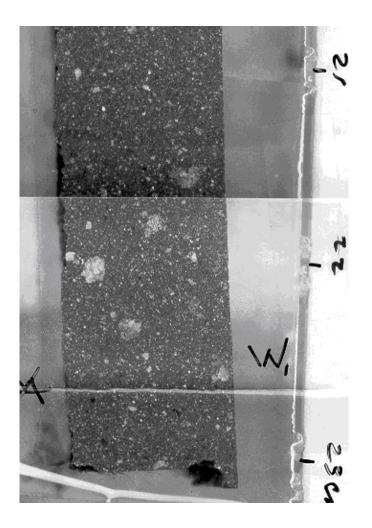
 $W_1$ 

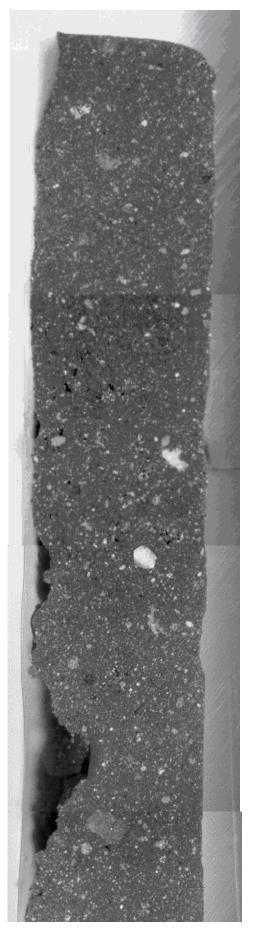










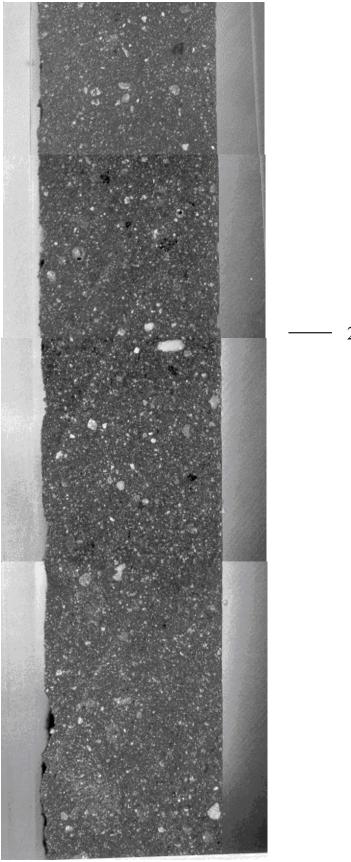


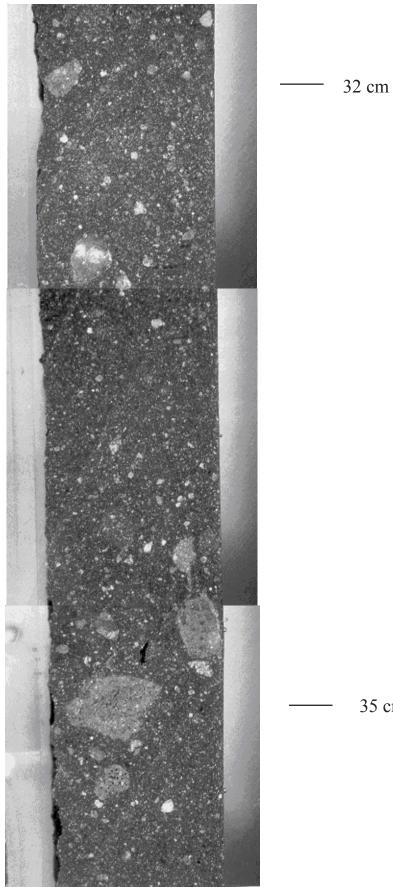
15007,6000 epoxy encapsulted core

# $W_1$

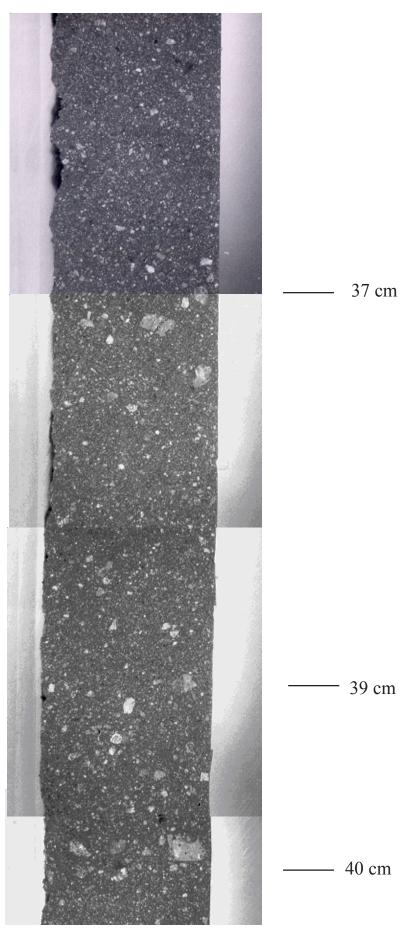
— 25 cm

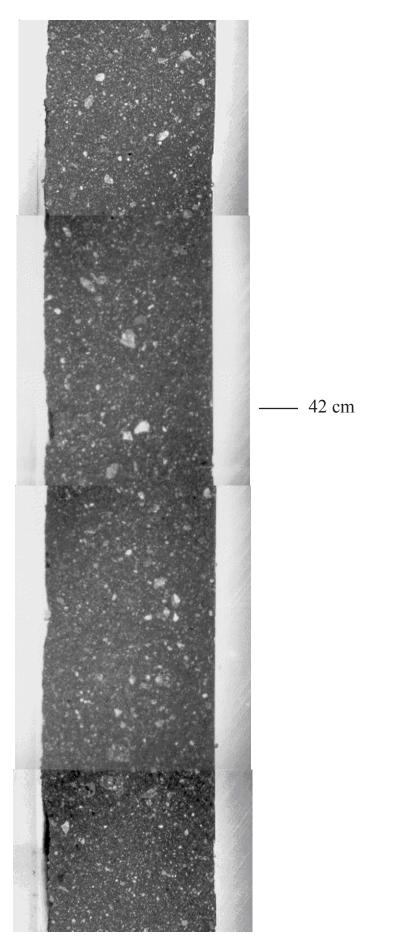
27 cm

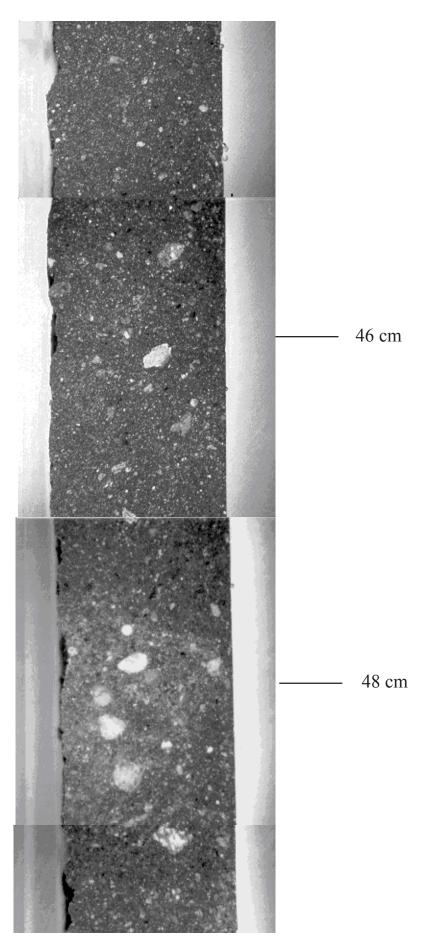


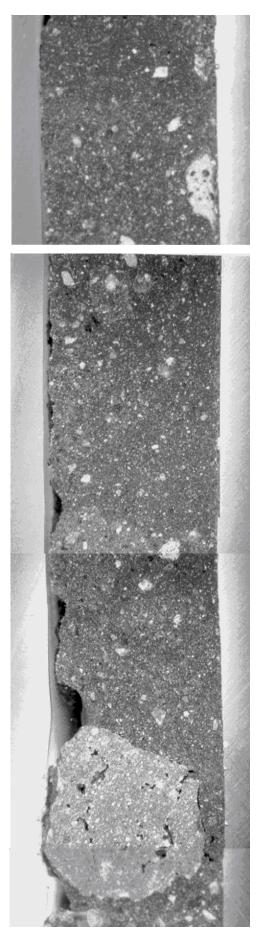




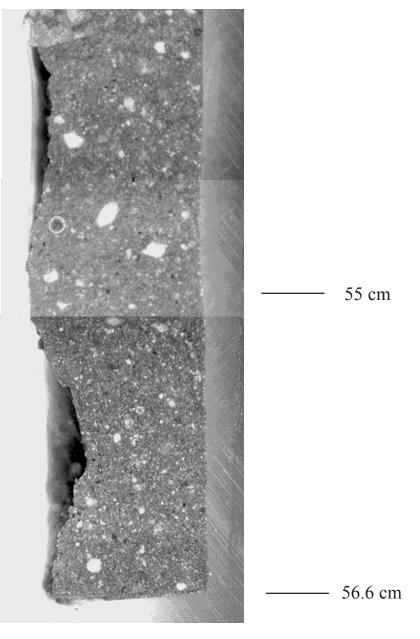








50 cm



bottom