

15311
Soil (rake)
464 grams

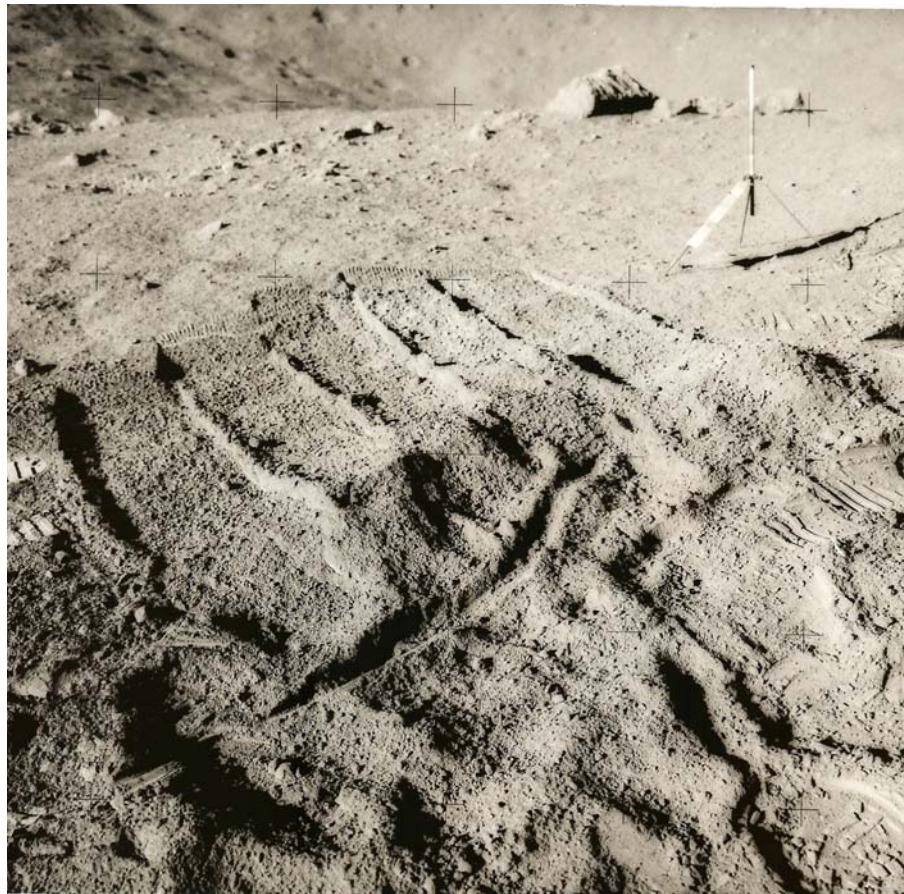


Figure 1: Raked area from rim of Spur Crater. AS15-90-12234. The feet on the Gnomon are 50 cm apart.

Introduction

Soil sample 15310 was collected as part of a large rake sample (figure 1) and may contain material rubbed off of the friable rocks collected by the rake. On the other hand, it is certain to also contain some soil collected by the rake and/or attached to the rocks. In any case 15311 is a large soil sample that goes along with 15301 and it has about the same composition..

Petrography

The fines from this sample have not been characterized. An unpublished mode by Heiken and McKay (1972) reported an abundance of green glass particles. Indeed several friable “green glass” clods were included in the rake sample and in the coarse-fines.

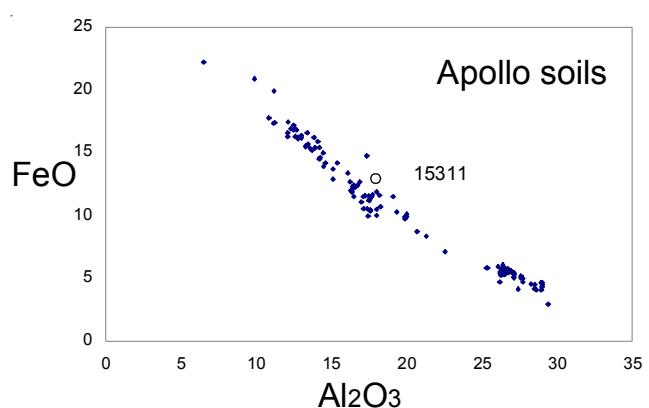


Figure 2: Composition of 15311 compared with other Apollo soil samples.

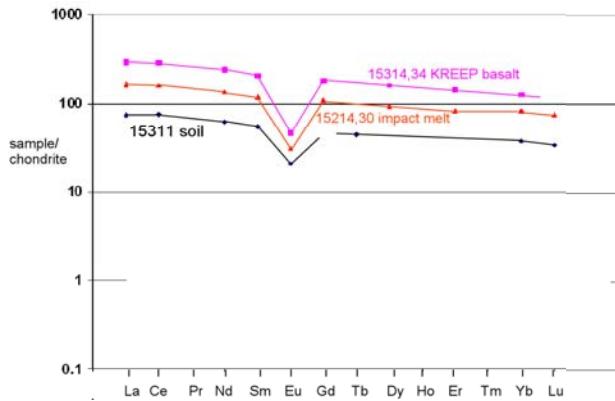


Figure 3: Normalized rare-earth-element diagram for 15311 soil, 15314 impact melt rock and 15314 KREEP basalt (see tables).

The coarse fines (4-10 mm) were cataloged by Powell (1972), Phinney et al. (1972) and Ryder and Sherman (1989). They reported on 5 mare basalt, 1 KREEP basalt, 8 impact melt rocks, 3 green glass clods, 10 regolith breccias and one ropy glass particle in 15434. Compare this with the rake sample which was in the ratio of 7 mare basalt, 2 KREEP basalt, 3 impact melt rocks, 13 green glass clods, 39 regolith breccias and 2 agglutinates (table 2).

Chemistry

Korotev (1987) determined the bulk composition of 15311 (table 1). Several labs reported data on the coarse fines (table 3). Again this “soil” sample reflects the composition of KREEP – either from KREEP basalt or from the impact melt rocks (which themselves carry a high KREEP component) (figure 3).

Modal content of soil 15311

<i>From Morris et al. 1983</i>	
(Heiken and McKay 1972)	
Agglutinates	9 %
Basalt	5.5
Breccia	15
Anorthosite	
Norite	
Gabbro	
Plagioclase	3
Pyroxene	10
Olivine	1
Ilmenite	2.3
Green Glass	46 %
Yellow Glass	4.5
Glass other	2.5

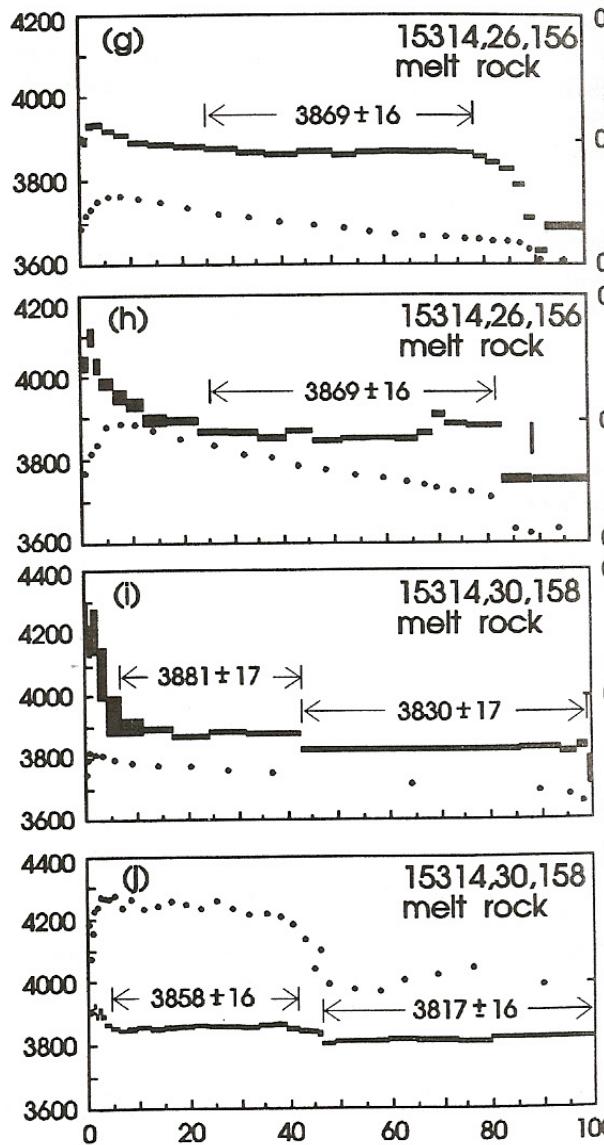


Figure 4: Ar/Ar plateau diagrams for two impact melt rocks from the coarse fines from 14311 (Dalrymple and Ryder 1993).

Summary of Age Data for 15314

	Ar/Ar
Dalrymple and Ryder 1991	3869 ± 16 m.y.
	3881 ± 17 m.y.

As is the case for 15301, this soil has high Mg and Fe due to a significant amount of mafic green glass particles found in abundance in and around Spur Crater (see 15425 – 15427).

Radiogenic age dating

Dalrymple and Ryder (1991, 1993) determined the age of two impact melt rocks found in the coarse-fines using the Ar/Ar dating technique (figure 3), concluding that

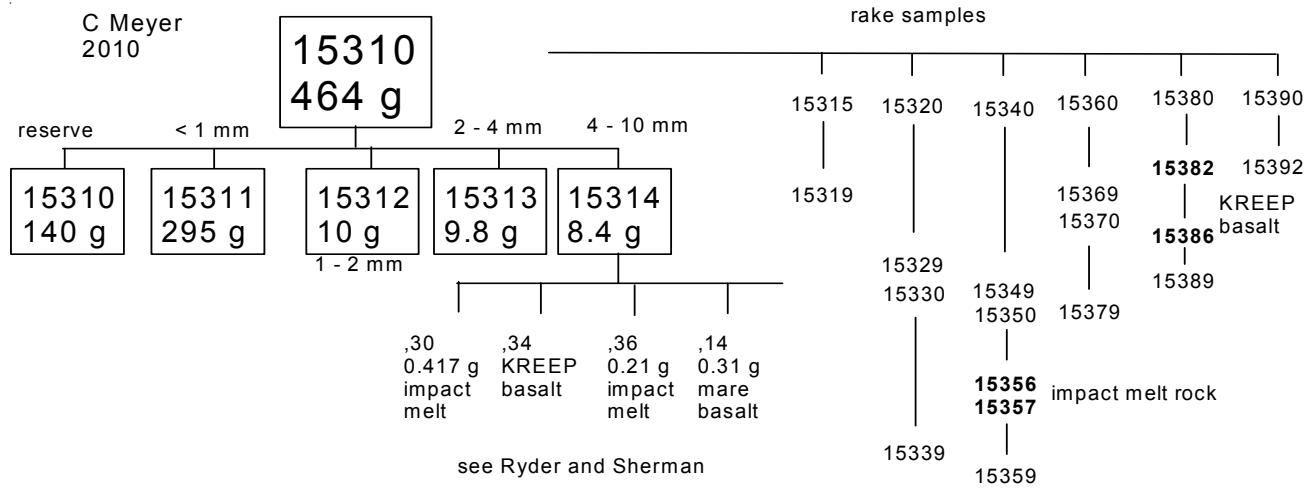


Table 2: Rake samples from 15310.

	grams	Ryder's name.		
15315	35.6	regolith breccia	15354	0.3
15316	6.1	regolith breccia	15355	5.2
15317	0.6	regolith breccia	15356	2
15318	5.4	regolith breccia	15357	11.8
15319	8	regolith breccia	15358	14.6
15320	4.7	regolith breccia	15359	4.2
15321	0.3	regolith breccia	15360	9.3
15322	8.4	regolith breccia	15361	0.9
15323	4.4	regolith breccia	15362	4.2
15324	32.3	regolith breccia	15363	0.5
15325	57.8	dense regolith breccia	15364	1.5
15326	2.5	regolith breccia	15365	2.9
15327	12.4	clast-rich glassy melt	15366	3.3
15328	0.3	regolith breccia	15367	1.1
15329	2.2	regolith breccia, glass-coated	15368	0.4
15330	57.8	regolith breccia	15369	2.5
15331	2.6	regolith breccia	15370	2.9
15332	2.3	agglutinate	15371	0.5
15333	0.3	regolith breccia	15372	0.8
15334	7.5	regolith breccia	15373	0.6
15335	6	regolith breccia	15374	1
15336	0.2	regolith breccia	15375	0.4
15337	4.3	regolith breccia	15376	1
15338	11.1	regolith breccia	15377	0.5
15339	0.4	regolith breccia	15378	3.3
15340	0.9	regolith breccia	15379	64.3
15341	1.6	regolith breccia	15380	5.2
15342	7.5	regolith breccia	15381	0.3
15343	6.9	regolith breccia	15382	3.2
15344	7.9	regolith breccia	15383	1.4
15345	12.3	vesicular glass with breccia	15384	1.4
15346	3.1	regolith breccia	15385	8.7
15347	3.2	regolith breccia	15386	7.5
15348	0.3	regolith breccia	15387	2
15349	2.3	regolith breccia	15388	9
15350	2.9	regolith breccia	15389	2.8
15351	4.2	regolith breccia	15390	3.5
15352	2.9	regolith breccia	15391	0.3
15353	10.6	regolith breccia	15392	0.4

Table 1. Chemical composition of 15311

reference Korotev87

weightSiO₂ %

TiO ₂	1.42	(a)
Al ₂ O ₃	16.9	(a)
FeO	12.7	(a)
MnO	0.17	(a)
MgO	11.9	(a)
CaO	10.4	(a)
Na ₂ O	0.43	(a)

K₂OP₂O₅

S %

sum

Sc ppm	24.5	(a)
V	89	(a)
Cr	2370	(a)
Co	45.9	(a)
Ni	215	(a)

Cu

Zn

Ga

Ge ppb

As

Se

Rb

Sr	110	(a)
Y		

Zr	260	(a)
Nb		

Mo

Ru

Rh

Pd ppb

Ag ppb

Cd ppb

In ppb

Sn ppb

Sb ppb

Te ppb

Cs ppm	0.19	(a)
Ba	190	(a)
La	17.1	(a)
Ce	45	(a)

Pr

Nd	28	(a)
Sm	8.16	(a)
Eu	1.15	(a)

Gd

Tb	1.6	(a)
Dy		

Ho

Er

Tm

Yb	6	(a)
Lu	0.81	(a)
Hf	6.9	(a)
Ta	0.8	(a)

W ppb

Re ppb

Os ppb

Ir ppb	4.8	(a)
Pt ppb		

Au ppb	4.5	(a)
Th ppm	2.8	(a)
U ppm	0.86	(a)

technique: (a) INAA

3.87 b.y. was the age of the Imbrium impact ! Wow !
So much information from such small particles !

Processing

15311 was returned in sample collection bag 3 placed in ALSRC#2 (which did not seal).

It is the soil collected and returned along with the rake sample. As such, it probably contains material abraded off of friable rake samples.

PS Samples greater than 1 cm are called "walnuts" and samples in the 4 – 10 mm size range are called :coarse-fines, or "peanuts".

Table 3. Composition of some particles from 15310.

15314 reference	,29,149 Laul87	,26,97 Ryder87	,30 Wiesman76	,30 same	,32 Laul87	,34 Hubbard73
weight	impact melt	impact melt	impact melt	same	impact melt	KREEP basalt
SiO ₂ %		48 (b)		48.5 (b)		
TiO ₂	1.2 (a)	1.37 (b)	1.09 (c)	1.31 (b)	0.97 (a)	2.02 (c)
Al ₂ O ₃	15.2 (a)	17.3 (b)		16.1 (b)	16.2 (a)	16.7
FeO	10.2 (a)	8.6 (b)		8.6 (b)		10.4
MnO	0.125 (a)	0.143 (b)		0.15 (b)	8.7 (a)	
MgO	16.8 (a)	12.8 (b)		13.8 (b)	0.122 (a)	8.56 (c)
CaO	9.4 (a)	10.5 (b)		10.2 (b)	13.2 (a)	7.67 (c)
Na ₂ O	0.4 (a)	0.52 (b)		0.58 (b)	9.7 (a)	0.68
K ₂ O	0.18 (a)	0.35 (b)	0.3 (c)	0.35 (b)	0.55 (a)	0.53 (c)
P ₂ O ₅		0.37 (b)		0.34 (b)	0.44 (a)	
S %						
sum						
Sc ppm	16 (a)	21.1 (a)		20.2 (a)	18 (a)	
V	60 (a)				60 (a)	
Cr		1770 (a)				
Co	61 (a)	50.6 (a)		54.4 (a)	30.5 (a)	
Ni	620 (a)	305 (a)		192 (a)	210 (a)	
Cu						
Zn						
Ga						
Ge ppb						
As						
Se						
Rb		15 (a)				
Sr	150 (a)		8.11 (a)	13 (c)	13 (a)	13.6 (c)
Y			132 (c)		180 (a)	185 (c)
Zr	400 (a)	390 (a)	492 (a)	485 (c)	420 (a)	1009 (c)
Nb						
Mo						
Ru						
Rh						
Pd ppb						
Ag ppb						
Cd ppb						
In ppb						
Sn ppb						
Sb ppb						
Te ppb						
Cs ppm		0.37 (a)		0.34 (a)		
Ba	350 (a)	450 (a)	379 (c)	413 (a)	450 (a)	676 (c)
La	41 (a)	45.6 (a)	38.6 (c)	42.2 (a)	40.5 (a)	67.9 (c)
Ce	105 (a)	124 (a)	97.3 (c)	118 (a)	100 (a)	174 (c)
Pr					(a)	
Nd	70 (a)	74 (a)	61.3 (c)	74 (a)	64 (a)	108 (c)
Sm	18.3 (a)	22.1 (a)	17.4 (c)	20.3 (a)	17.2 (a)	30.2 (c)
Eu	1.9 (a)	2.01 (a)	1.7 (c)	1.85 (a)	1.95 (a)	2.58 (c)
Gd	23 (a)		21 (c)		22 (a)	35.9 (c)
Tb	3.7 (a)	4.3 (a)		4.6 (c)	3.5 (a)	
Dy	22 (a)		22.5 (c)		23 (a)	39.2 (c)
Ho	4.9 (a)				5 (a)	
Er			13.2 (c)			22.9 (c)
Tm	1.8 (a)			2 (a)		
Yb	11 (a)	14.9 (a)	13.5 (c)	12.2 (a)	13 (a)	20.5 (c)
Lu	1.7 (a)	2.14 (a)	1.8 (c)	2.03 (a)	1.9 (a)	
Hf	11 (a)	15.6 (a)	12.5 (c)	13.1 (a)		25.7 (c)
Ta	1.5 (a)	1.9 (a)		1.8 (a)		
W ppb						
Re ppb						
Os ppb						
Ir ppb	15 (a)			2.9 (a)		
Pt ppb						
Au ppb	12.7 (a)			4 (a)		
Th ppm	6 (a)		5.89 (c)	7 (a)	6.7 (a)	10.8 (c)
U ppm	1.5 (a)		1.84 (c)	2.5 (a)	1.6 (a)	3.1 (c)

technique: (a) INAA, (b) fused-bead e-probe, (c) IDMS

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