

**73275**  
Micropoikilitic Impact melt Breccia  
429.6 grams

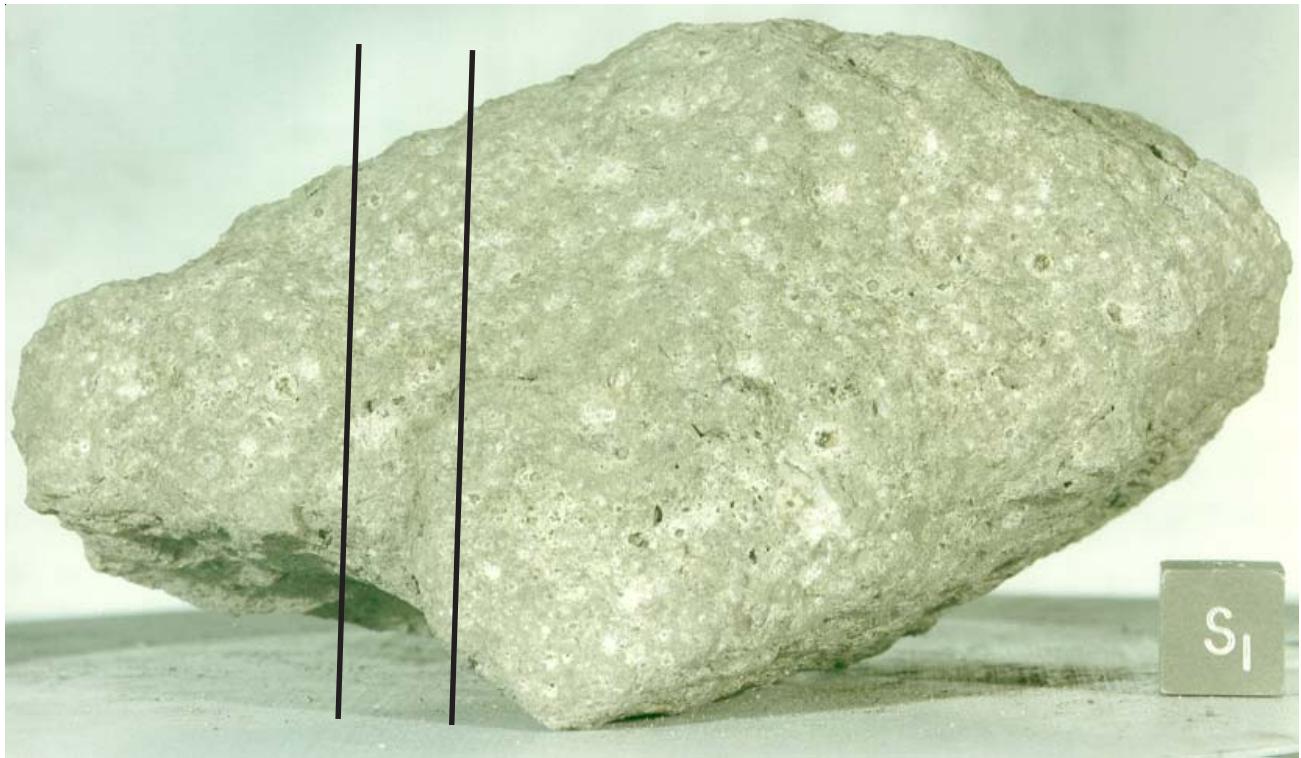


Figure 1: Photo of 73275. NASA S73-16925. Cube is 1cm. Note numerous zap pits and smooth-rounded surface..

### Introduction

Lunar sample 73275 was collected from station 3, Apollo 17 (Wolfe et al. 1981). It was found sitting perched on the regolith near the rim of a small crater (figures 2, 3). It is rounded and covered with micrometeorite craters on all sides except where a piece broke off (figure 1).

73275 is an impact melt breccia with fine-grained poikilitic matrix and minor vesicularity (figure 4). It has a crystallization age of about 3.9 b.y. with exposure to cosmic rays about 140 m.y.

### Petrography

Ryder (1993) describes 73275 as a clast-bearing micropoikilitic impact melt breccia with a composition similar to other samples thought to be from the Serenitatis impact event.

Simonds et al. (1974) describes the matrix of 73275 as subophitic-micropoikilitic, giving the grain size of

plagioclase as 10-50 microns (figure 4). Heuer et al. (1974) and Radcliffe et al. (1974) describe 73275 as a recrystallized breccia with large clasts of plagioclase (0.3-1.0 mm), orthopyroxene (0.1-0.2 mm) and smaller olivine (0.05-0.2 mm). They reported the calcic cores of plagioclase ( $An_{90-97}$ ) have sodic overgrowths ( $An_{85-90}$ ) and the pyroxene has wide exsolution lamellae indicating prolonged annealing at high temperature.

73275 apparently lacks “flow banding” and relic aphanitic clasts that are characteristic of adjacent samples (732xx).

Schneider and Horz (1974) determined the size distribution of micrometeorite craters.

### Mineralogy

**Pyroxene:** Simonds et al. (1974) determined the pyroxene compositions in the matrix and small clasts (figure 5).

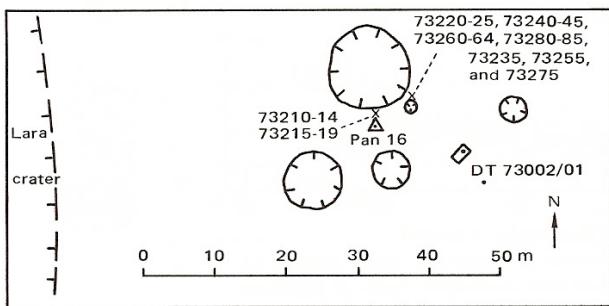


Figure 2: Location of 73275 at Station 3, Apollo 17.

**Plagioclase:** Heuer et al. (1974) reported plagioclase composition ( $\text{An}_{90-97}$ ).

**Metallic iron:** Goldstein et al. (1976) analyzed the Ni, Co and C content of iron grains found interstitial to silicates in 73275 (Ni = ~6%, Co = ~0.6%).

**Cohenite:** Goldstein et al. (1976) found cohenite ((Fe,Ni)<sub>3</sub>C) apparently exsolved from kamacite (figure 6 and 7). The carbon (and Ni) is probably of meteoritic origin.

## Chemistry

The chemical composition of 73275 is generally similar to that of the other station three breccias and to that of 77135 from a boulder on the other side of the Taurus-Littrow valley (figure 8). Morgan et al. (1976) found the meteoritic siderophile elements were a match with other samples from the Serenitatis Basin (table 1).

## Radiogenic age dating

Turner and Cadogen (1975) dated 73215 at  $3.96 \pm 0.05$  b.y. by the Ar 39/40 method (figure 9). Nyquist et al. (1974) and Oberli et al. (1978) determined whole rock Rb-Sr, Sm-Nd and U-Th-Pb isotopic data.

## Cosmogenic isotopes and exposure ages

Turner and Cadogen (1975) determined the cosmic ray exposure age as 160 m.y. by  $^{38}\text{Ar}$  (figure 9). Crozaz et al. (1974) determined the CRE age as 139 m.y. by  $^{81}\text{Kr}$ . Crozaz et al. (1974) and Goswami and Lal (1974) determined cosmic ray track density yielding "ages" in the range of a few m.y.

## Other Studies

Nagata et al. (1974) and Housley et al. (1976) studied the magnetic properties of 73275. Huffman et al. (1974) and Huffman and Dunphyre (1975) studied 73275 by Mossbauer spectroscopy.

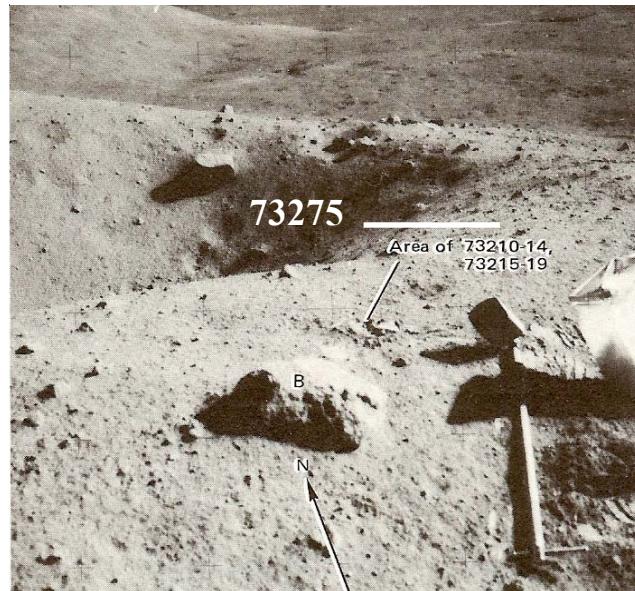


Figure 3: Surface photo of 10 meter crater at station 3.

## Processing

A slab was cut from 73275 (figure 10). There are 16 thin sections of 73275.

## References for 73275

Butler P. (1973) Lunar Sample Information Catalog Apollo 17. Lunar Receiving Laboratory. MSC 03211 Curator's Catalog. pp. 447.

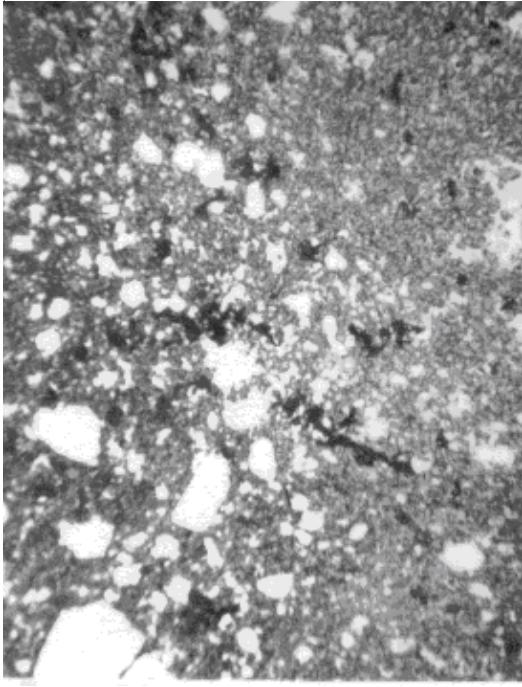
Crozaz G., Drozd R., Hohenberg C., Morgan C., Ralston C., Walker R. and Yuhas D. (1974a) Lunar surface dynamics: Some general conclusions and new results from Apollo 16 and 17. *Proc. 5<sup>th</sup> Lunar Sci. Conf.* 2475-2499.

Dalrymple G.B. and Ryder G. (1996) Argon-40/Argon-39 age spectra of Apollo 17 highland breccia samples by laser step heating and the age of the Serenitatis basin. *J. Geophys. Res.* 101, 26,069-26,084.

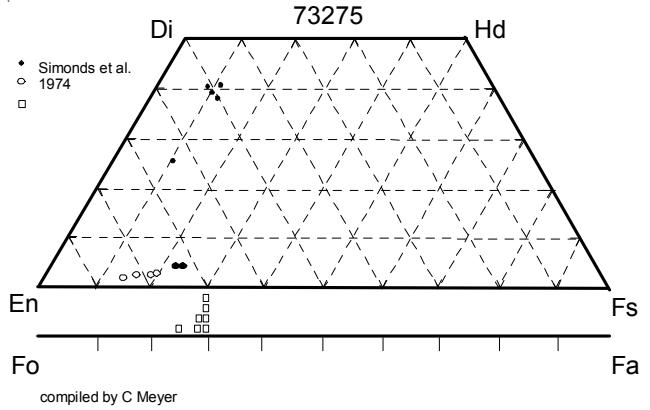
Eldridge J.S., O'Kelley G.D. and Northcutt K.J. (1974a) Primordial radioelement concentrations in rocks and soils from Taurus-Littrow. *Proc. 5<sup>th</sup> Lunar Sci. Conf.* 1025-1033.

Goldstein J.I., Hewins R.H. and Romig A.D. (1976a) Carbides in lunar soils and rocks. *Proc. 7<sup>th</sup> Lunar Sci. Conf.* 807-818.

Goswami J.N. and Lal D. (1974) Cosmic ray irradiation pattern at the Apollo 17 site: implications to lunar regolith dynamics. *Proc. 5<sup>th</sup> Lunar Sci. Conf.* 2643-2662.

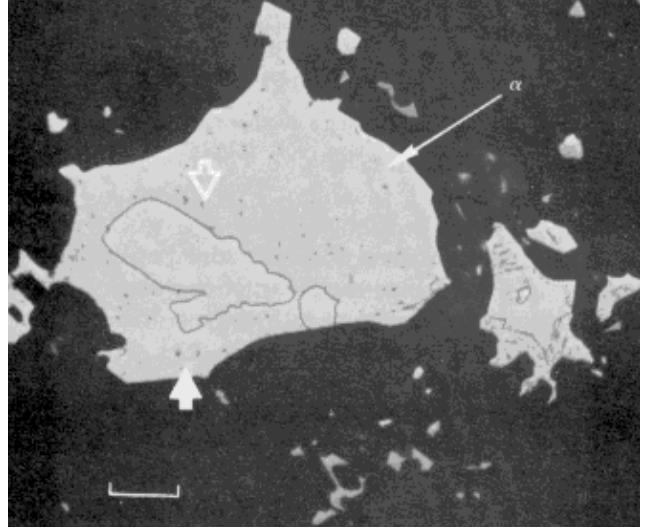


*Figure 4: Photomicrograph of thin section of 73275 showing “micropoikilitic” texture (from Ryder 1993). Field of view is 2 mm.*



compiled by C Meyer

*Figure 5: Pyroxene and olivine composition of 73275 (matrix and clasts) from Simonds et al. (1974).*



*Figure 6: Reflected light photo of metallic iron grain in 73275, 68 showing cohenite included in kamacite (Goldstein et al. 1976). Scale bar is 16 microns.*

Hartung J.B., Horz F., Aitken F.K., Gault D.E. and Brownlee D.E. (1973) The development of microcrater populations on lunar rocks. *Proc. 4<sup>th</sup> Lunar Sci. Conf.* 3213-3234.

Heuer A.H., Christie J.M., Lally J.S. and Nord G.L. (1974) Electron petrographic study of some Apollo 17 breccias. *Proc. 5<sup>th</sup> Lunar Sci. Conf.* 275-286.

Housley R.M. and Grant R.W. (1976) ESCA studies of the surface chemistry of lunar fines. *Proc. 7<sup>th</sup> Lunar Sci. Conf.* 881-889.

Housley R.M., Cirlin E.H., Goldberg I.B. and Crowe H. (1976) Ferromagnetic resonance studies of lunar core stratigraphy. *Proc. 7<sup>th</sup> Lunar Sci. Conf.* 13-26.

Huffman G.P., Schwerer F.C., Fisher R.M. and Nagata T. (1974) Iron distribution and metallic-ferrous ratios for Apollo lunar samples: Mossbauer and magnetic analyses. *Proc. 5<sup>th</sup> Lunar Sci. Conf.* 2779-2794.

Huffman G.P. and Dunmyre G.R. (1975) Superparamagnetic clusters of Fe+2 spins in lunar olivine: Dissolution by high-temperature annealing. *Proc. 6<sup>th</sup> Lunar Sci. Conf.* 757-772,

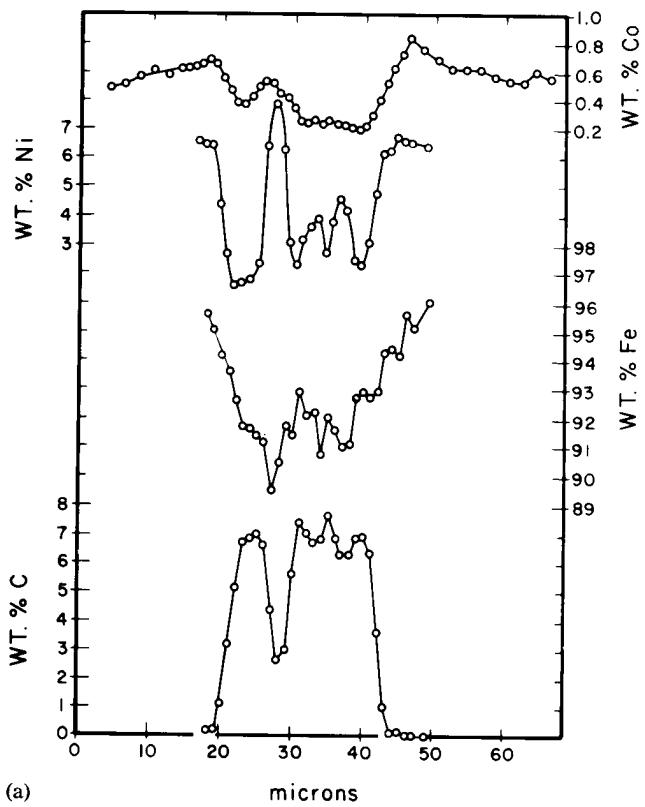
LSPET (1973) Apollo 17 lunar samples: Chemical and petrographic description. *Science* **182**, 659-672.

LSPET (1973) Preliminary Examination of lunar samples. Apollo 17 Preliminary Science Rpt. NASA SP-330. 7-1 – 7-46.

Meyer C. (2010) Lunar Sample Compendium (abs#1016). The 41<sup>st</sup> *Lunar Planet. Sci. Conf.* @ The Woodlands

Morgan J.W., Gros J., Takahashi H. and Hertogen J. (1976) Lunar breccia 73215: siderophile and volatile elements. *Proc. 7<sup>th</sup> Lunar Sci. Conf.* 2189-2199.

Muehlberger et al. (1973) Documentation and environment of the Apollo 17 samples: A preliminary report. Astrogeology 71 322 pp superceeded by Astrogeology 73 (1975) and by Wolfe et al. (1981)

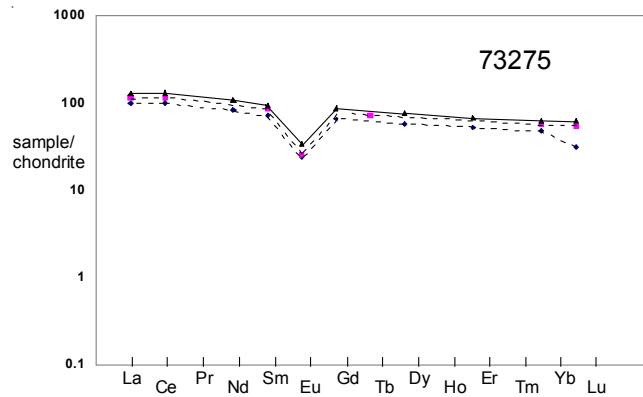


*Figure 7: Electron microprobe traverse across metallic iron grain (with carbon) in 73275,68 (from Goldstein et al. 1976).*

### Summary of Age Data for 73275

	Ar/Ar	U/Pb
Turner and Cadogen 1975	$3.96 \pm 0.05$ b.y.	
Oberli et al. 1978		$\sim 4.42$ b.y.

**Caution: Data is with old decay constants.**

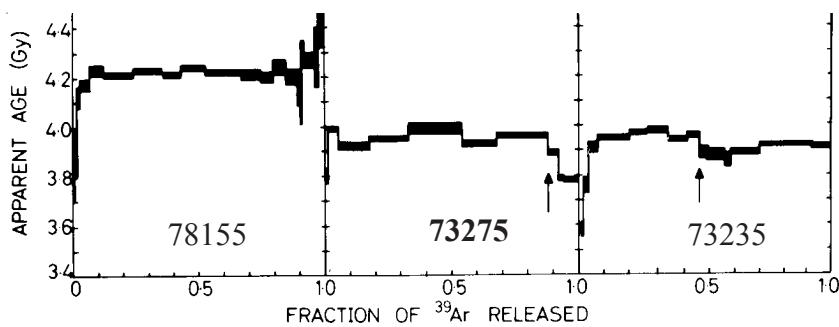


*Figure 8: Normalized rare-earth-element diagram for 73275 (solid) compared with that of 73215 and 73235 (dashed). Data from Wiesmann and Hubbard 1975.*

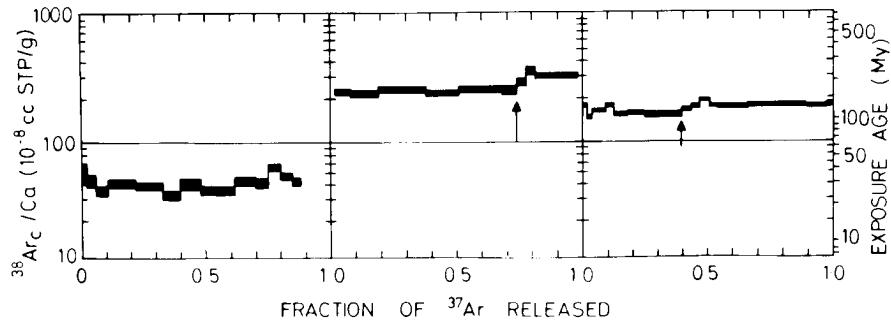
Muehlberger W.R. and many others (1973) Preliminary Geological Investigation of the Apollo 17 Landing Site. In **Apollo 17 Preliminary Science Report**. NASA SP-330.

Nagata T., Sugiura N., Fisher R.M., Schwerer F.C., Fuller M.D. and Dunn J.R. (1974a) Magnetic properties of Apollo 11-17 lunar materials with special reference to effects of meteorite impact. *Proc. 5<sup>th</sup> Lunar Sci. Conf.* 2827-2839.

Norman M.D., Bennett V.C and Ryder G. (2002) Targeting the impactors: Siderophile element signatures of lunar impact melts from Serenitatis. *Earth Planet Sci. Lett.* 202, 217-228.



*Figure 9: Ar release pattern for 73235, 73275 and 78155 (from Turner and Cadogen 1975).*



**Table 1. Chemical composition of 73275.**

reference	Eldridge 1974	Wiesmann75	Rhodes74	Morgan74	Oberli78	Norman2002	
<i>weight</i>							
SiO <sub>2</sub> %		46.16	(c )		46.8	(e)	
TiO <sub>2</sub>		1.43	(c )		1.24	(e)	
Al <sub>2</sub> O <sub>3</sub>		18.49	(c )		18.5	(e)	
FeO		9.05	(c )		7.84	(e)	
MnO		0.13	(c )		0.12	(e)	
MgO		11.54	(c )		11.7	(e)	
CaO		11.3	(c )		11.1	(e)	
Na <sub>2</sub> O		0.67	(c )		0.6	(f)	
K <sub>2</sub> O	0.27	(b) 0.27	(a) 0.27	(c )		0.17	(f)
P <sub>2</sub> O <sub>5</sub>				0.26 (c )			
S %				0.08 (c )			
<i>sum</i>							
Sc ppm					17.6	(d)	
V					51	(d)	
Cr	1221	(a)			1442	(d)	
Co					26.7	(d)	
Ni			182	(c )	214	(d)	
Cu					14.5	(d)	
Zn			2.5	(c )	13.3	(d)	
Ga					4.7	(d)	
Ge ppb			265	(c )			
As							
Se							
Rb	6.625	(a)	6.9	(c ) 9.11	(a) 6.8	(d)	
Sr	172	(a)		185	(a) 166	(d)	
Y					106	(d)	
Zr	427	(a)			482	(d)	
Nb					32	(d)	
Mo							
Ru					11.8	(d)	
Rh							
Pd ppb			0.74	(c )	11.8	(d)	
Ag ppb			4.1	(c )			
Cd ppb							
In ppb							
Sn ppb							
Sb ppb			1.19	(c )			
Te ppb			5.5	(c )			
Cs ppm			0.27	(c )	0.26	(d)	
Ba	333	(a)			328	(d)	
La	30.2	(a)			27	(d)	
Ce	78.6	(a)			68	(d)	
Pr					9.47	(d)	
Nd	49	(a)		50.7	(a) 43.7	(d)	
Sm	13.9	(a)		14.3	(a) 12.4	(d)	
Eu	1.89	(a)			1.69	(d)	
Gd	17	(a)			13.8	(d)	
Tb					2.46	(d)	
Dy	18.3	(a)			15.6	(d)	
Ho					3.4	(d)	
Er	10.7	(a)			9.71	(d)	
Tm							
Yb	9.91	(a)			8.8	(d)	
Lu	1.48	(a)			1.28	(d)	
Hf					9.94	(d)	
Ta					1.37	(d)	
W ppb					0.64	(d)	
Re ppb			0.494	(c )	0.67	(d)	
Os ppb							
Ir ppb			5.71	(c )	6.49	(d)	
Pt ppb					13.8	(d)	
Au ppb			3.34	(c )			
Th ppm	4.53	(b) 4.26	(a)		4.97	(a) 5.69	(d)
U ppm	1.2	(b) 1.36	(a)	1.36	(c) 1.31	(a) 1.56	(d)

technique: (a) IDMS, (b) radiation counting, (c) RNAA, (d) ICP-MS, (e) fused-bead elec. probe, (f) INAA

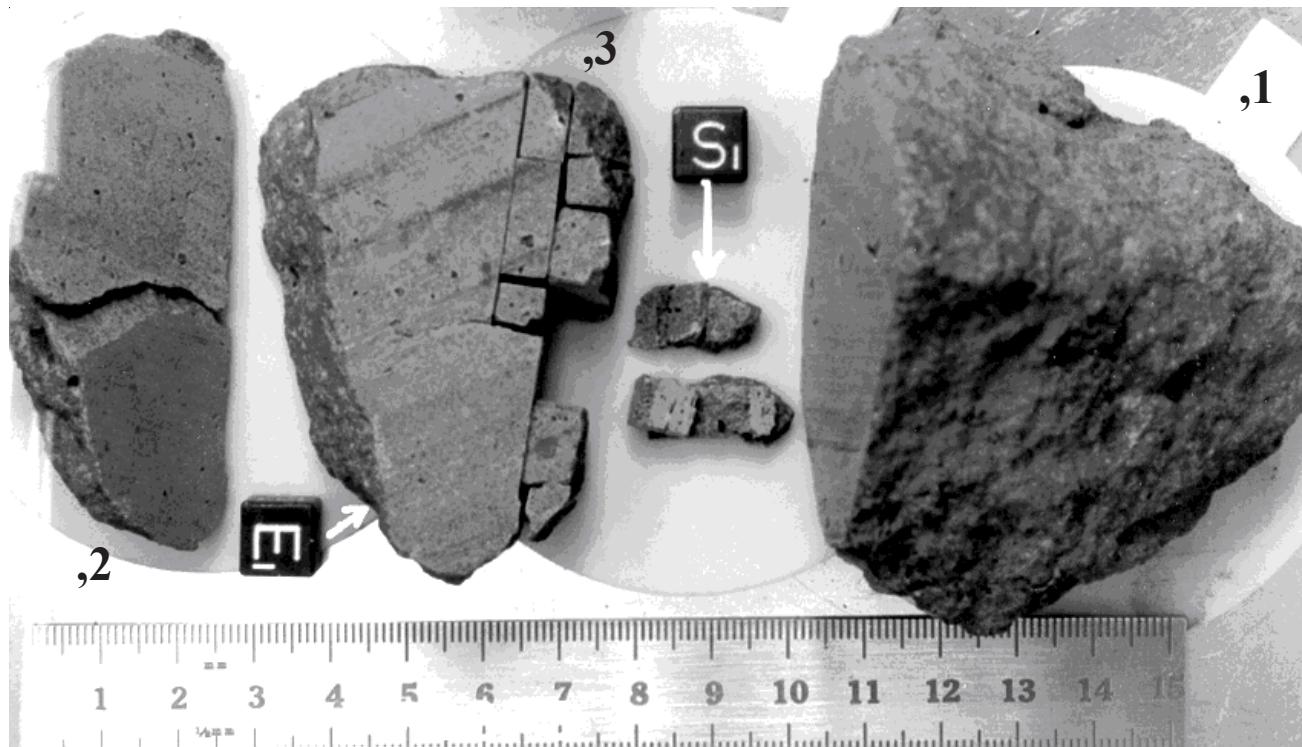
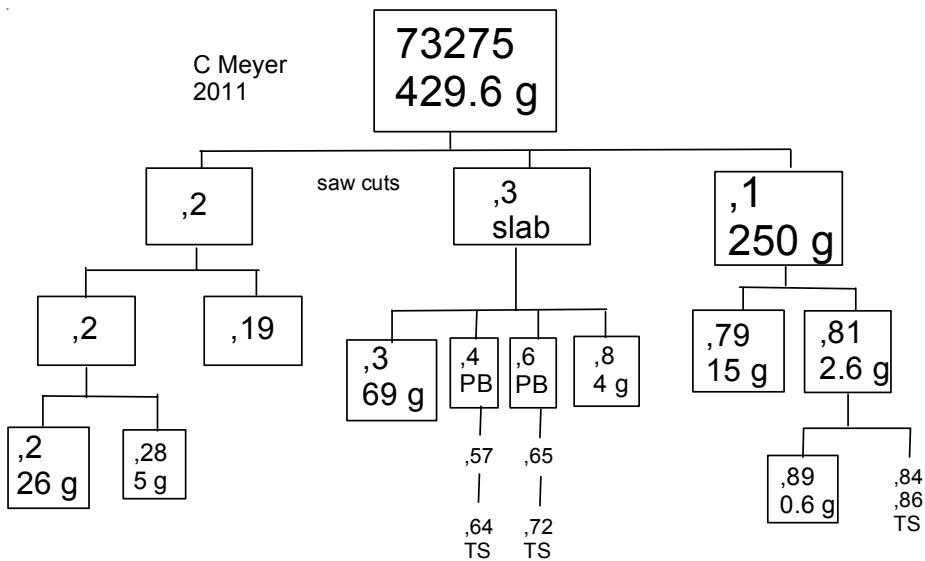


Figure 10: Exploded parts diagram for 73275. NASA S73-34460. Small cube and scale are 1 cm.

Nyquist L.E., Bansal B.M., Wiesmann H. and Jahn B.-M. (1974a) Taurus-Littrow chronology: some constraints on early lunar crustal development. *Proc. 5<sup>th</sup> Lunar Sci. Conf.* 1515-1539.

Oberli F., McCulloch M.T., Tera F., Papanastassiou D.A. and Wasserburg G.J. (1978) Early lunar differentiation constraints from U-Th-Pb, Sm-Nd and Rb-Sr model ages

(abs). *Lunar Planet. Sci. IX*, 832-834. Lunar Planetary Institute, Houston.

Radcliffe S.V., Christie J.M., Nord G.L., Lally J.S., Heuer A.H., Griggs D.T. and Fisher R.M. (1974) Electron petrographic evidence concerning the origin and lithification of the lunar breccias (abs). *Lunar Sci. V*, 613-615. Lunar Planetary Institute, Houston.

Rhodes J.M., Rodgers K.V., Shih C., Bansal B.M., Nyquist L.E., Wiesmann H. and Hubbard N.J. (1974a) The relationships between geology and soil chemistry at the Apollo 17 landing site. *Proc. 5<sup>th</sup> Lunar Sci. Conf.* 1097-1117.

Ryder G. (1992) Lunar highlands totality from bits and pieces: A whole-rock-geochemistry-free characterization of an evolved hypabyssal igneous gabbro schlieren from the Apollo 17 landing site. (abs) *LPS XXIII*, 1195-1196.

Ryder G. (1993c) Catalog of Apollo 17 rocks: Stations 2 and 3. Curators Office JSC#26088.

Schneider E. and Hörz F. (1974) Microcrater populations on Apollo 17 rocks. *Icarus* **22**, 459-473.

Simonds C.H., Phinney W.C. and Warner J.L. (1974) Petrography and classification of Apollo 17 non-mare rocks with emphasis on samples from the Station 6 boulder. *Proc. 5<sup>th</sup> Lunar Sci. Conf.* 337-353.

Steele I.M., Hutcheon I.D. and Smith J.V. (1980) Ion microprobe analysis and petrogenetic interpretations of Li, Mg, Ti, K, Sr, Ba in lunar plagioclase. *Proc. 11<sup>th</sup> Lunar Planet. Sci. Conf.* 571-590.

Turner G. and Cadogan P.H. (1975a) The history of lunar bombardment inferred from  $^{40}\text{Ar}$ - $^{39}\text{Ar}$  dating of highland rocks. *Proc. 6<sup>th</sup> Lunar Sci. Conf.* 1509-1538.

Wiesmann H. and Hubbard N.J. (1975) A compilation of the Lunar Sample Data Generated by the Gast, Nyquist and Hubbard Lunar Sample PI-Ships. Unpublished. JSC

Wolfe E.W., Bailey N.G., Lucchitta B.K., Muehlberger W.R., Scott D.H., Sutton R.L and Wilshire H.G. (1981) The geologic investigation of the Taurus-Littrow Valley: Apollo 17 Landing Site. US Geol. Survey Prof. Paper, 1080, pp. 280.