

**78505** - 506.3 grams

**78506** - 56 grams

**78507** - 23.4 grams

**78509** - 8.7 grams

Ilmenite Basalt



Figure 1: Photo of 78505. Scale and cube are in cm. S73-15386. White line shows approximate position of saw cut.

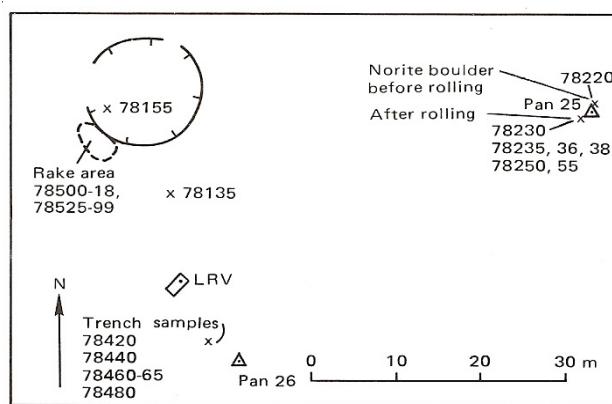


Figure 2: Map of station 8 showing location of soil 78500 and rake sample 78525-78599.

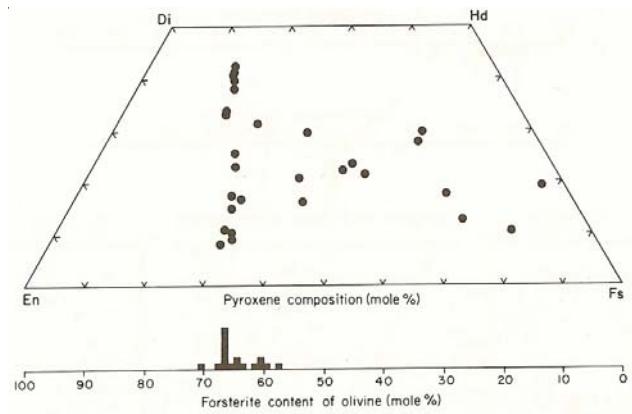


Figure 3: Composition of pyroxene and olivine in 78505 (Warner et al. 1978).



Figure 4: Photo of 78506. Scale in cm. S73-15468.

### **Mineralogical Mode**

	Warner	Brown
Olivine	0.5	0.5
Pyroxene	45.7	47.7
Plagioclase	35.7	27.7
Opaques	15.3	21
Silica	1.5	1.9
Meostasis	1.1	1.2

### **Introduction**

According to Warner et al. (1978) 78505 is a plagioclase-poikilitic ilmenite basalt similar to 71565, 71567 and 71509. It has a coarse grain size (1 mm) with abundant pigeonite and blocky ilmenite (figures 6 and 7).

78505 – 78509 were collected as part of a comprehensive soil sample (78500) at station 8, Apollo 17. A large rake sample collected adjacent to 78500 also returned abundant mare basalt fragments.

### **Petrography**

Brown et al. (1975) and others studied the Apollo 17 mare basalts including 78505 finding that they are all ilmenite-rich because of their high Ti content. 78505 is slightly different because of its proportion of low-Ca pigeonite (figure 3).



Figure 5: Photo of 78509. About 2 cm across.  
S73-18608

**Table 2: Zr-rich mineral 78506.**

(Brown et al. 1975)

TiO <sub>2</sub>	28.8
FeO	9.71
MgO	0.22
CaO	3.15
ZrO <sub>2</sub>	35.13
Y <sub>2</sub> O <sub>3</sub>	8.31
NbO <sub>5</sub>	2.7

Warner et al. (1978) reported tranquillityite and zirconolite in 78505. Brown et al. (1975) reported a Zr-rich mineral in 78506 (table 2). Roedder (1979) discusses the low-K inclusions in ilmenite in 78505.

Usselman et al. (1975) and others experimentally reproduced the textures and mineral chemistries of high-titanium mare basalts from Apollo 17 and discussed cooling rates.

### **Chemistry**

The chemical compositions of 78505, 78506, 78507 and 78509 are tabulated in table 1. These samples are typical of other Apollo 17 mare basalts (figures 10, 11 and 12).

### **Cosmogenic isotopes and exposure ages**

Drozd et al. (1977) determined the cosmic ray exposure age to be 121 m.y.

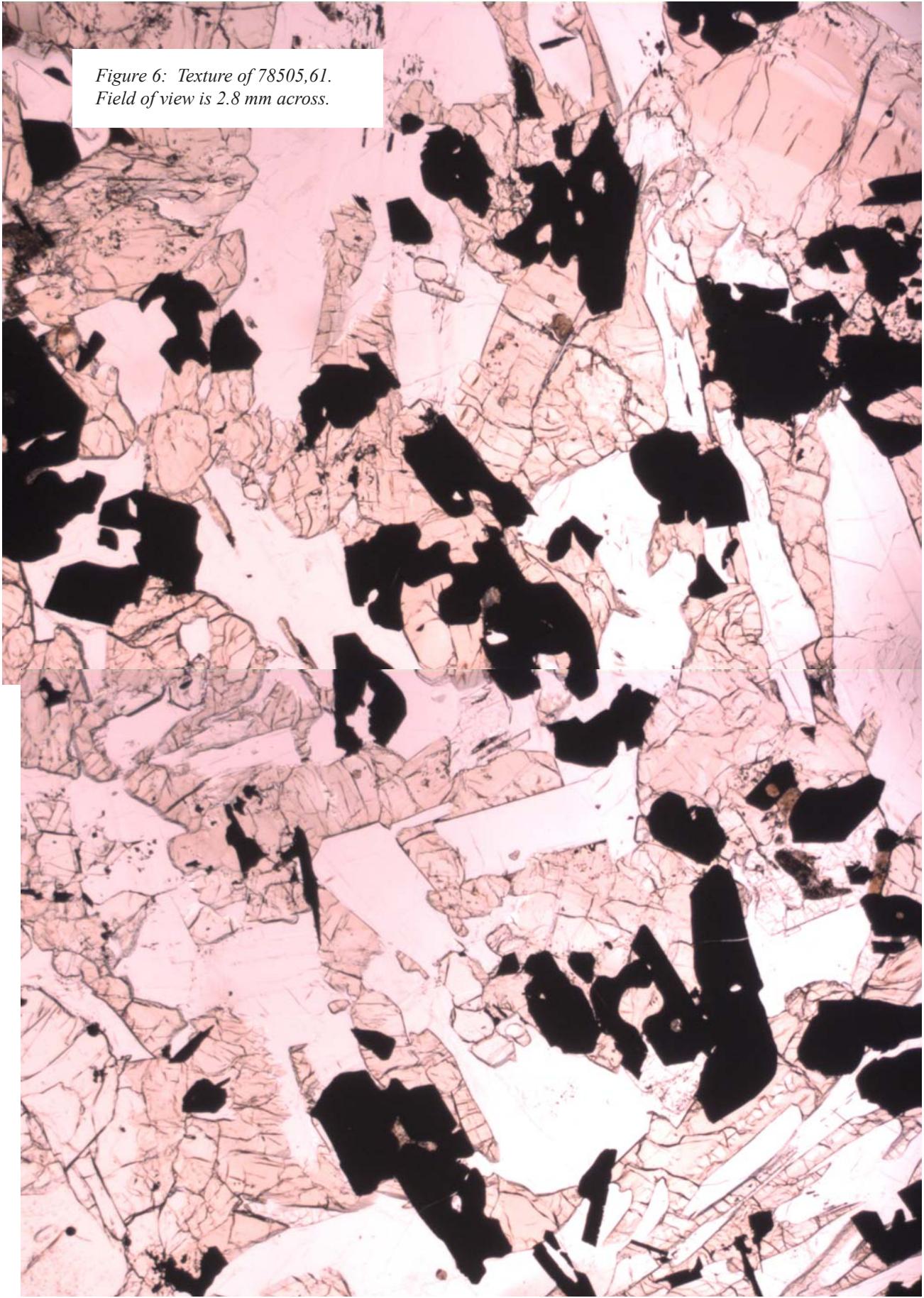
Keith et al. (1974) reported the cosmic-ray-induced activity of 78505.

### **Other Studies**

Stephenson et al. (1975) reported studies of the magnetic properties of 78505.

### **Processing**

78505 was sawn in half (figure 8) but it broke in pieces. There are 12 thin sections of 78505, 3 for 78506 and one each for 78507 and 78509.



*Figure 6: Texture of 78505, 61.  
Field of view is 2.8 mm across.*

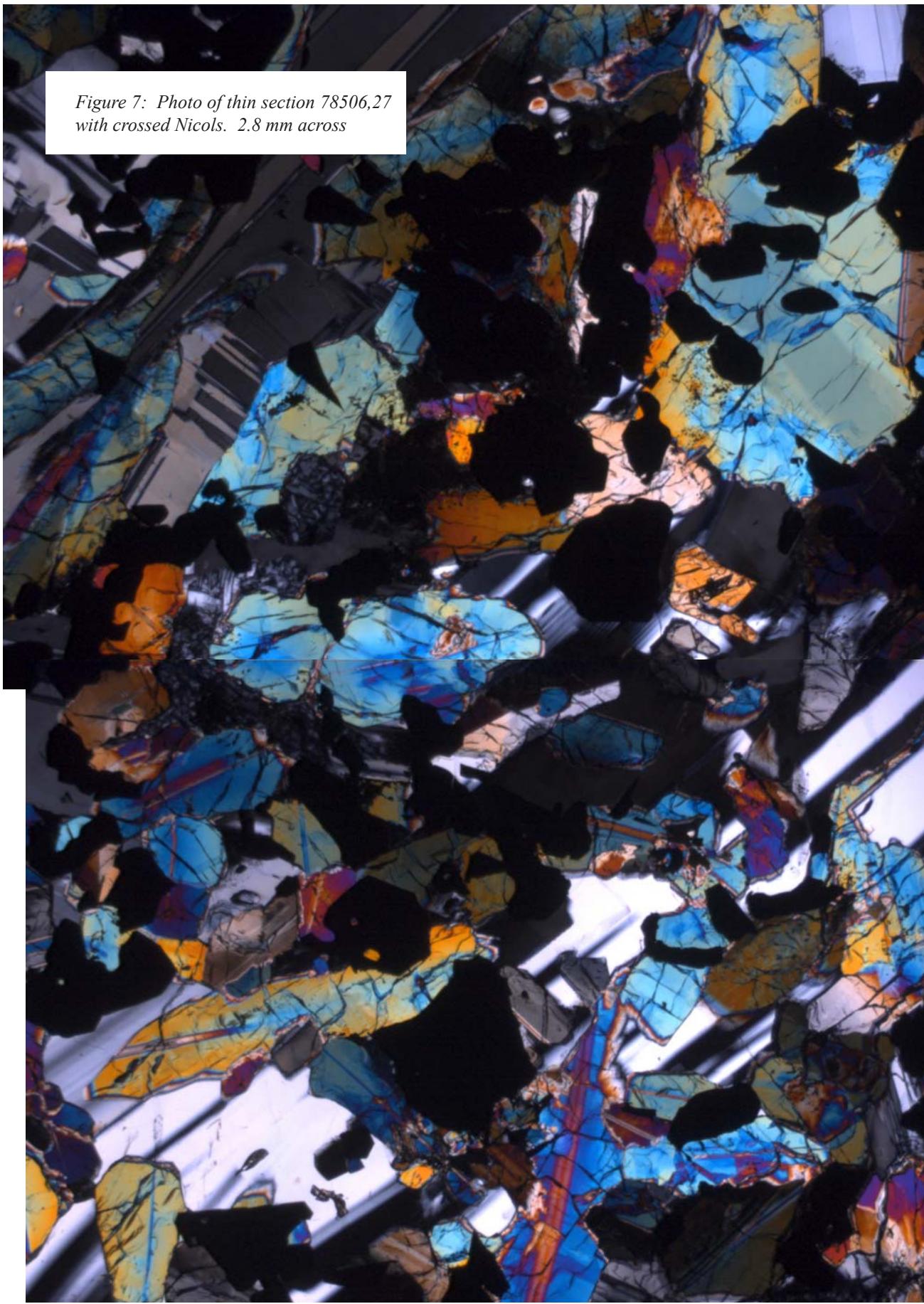


Lunar Sample Compendium  
C Meyer 2011



Lunar Sample Compendium  
C Meyer 2011

*Figure 7: Photo of thin section 78506,27  
with crossed Nicols. 2.8 mm across*



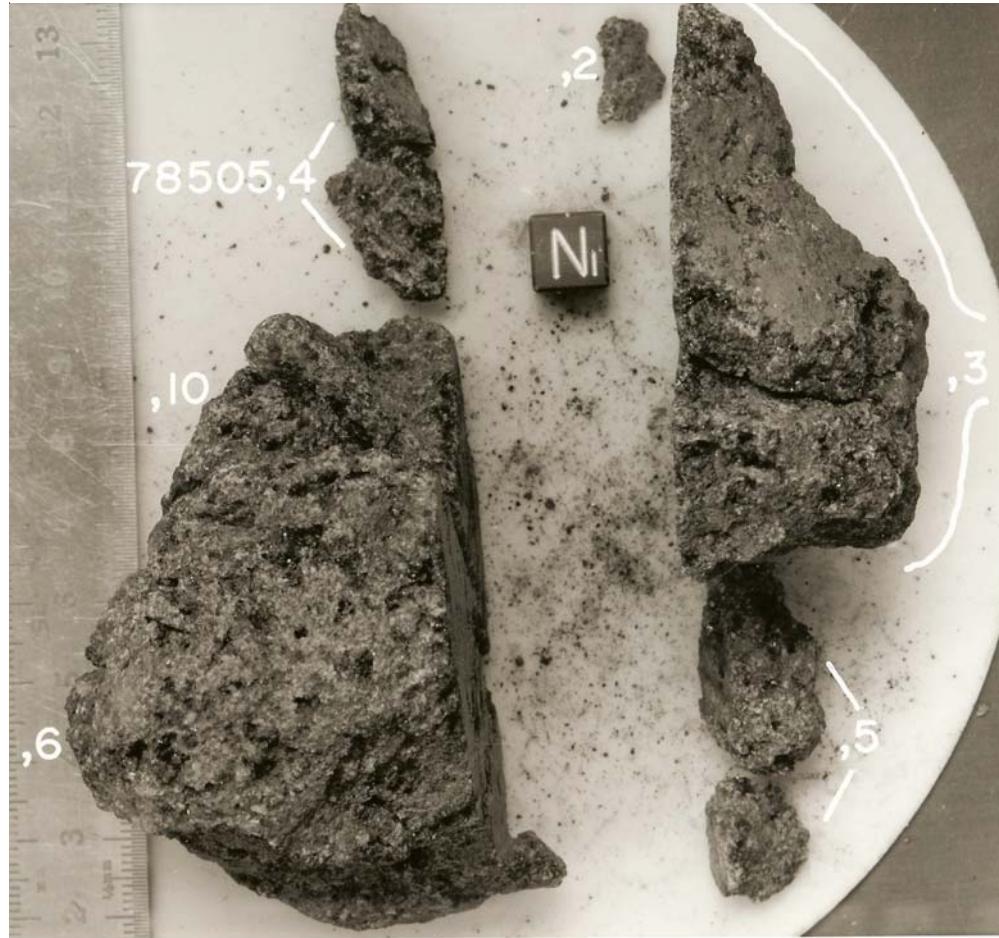
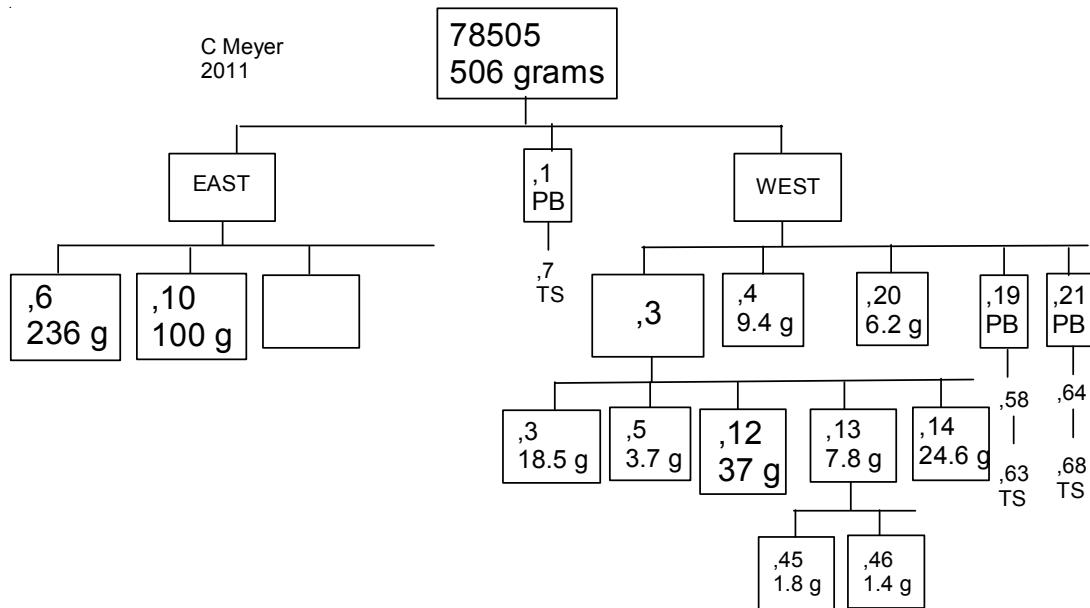


Figure 8: Processing photo for 78505. S74-19019 Cube is 1 cm<sup>3</sup>.



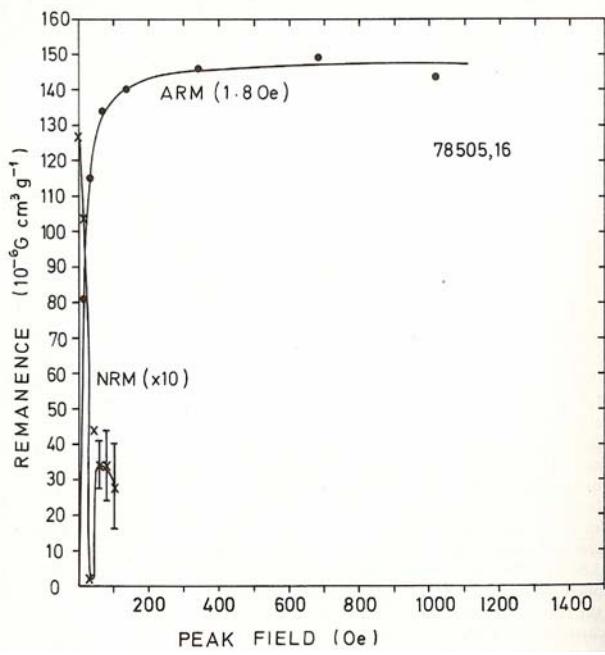


Figure 9: Magnetic properties of 78505 (Stephenson et al. 1975).

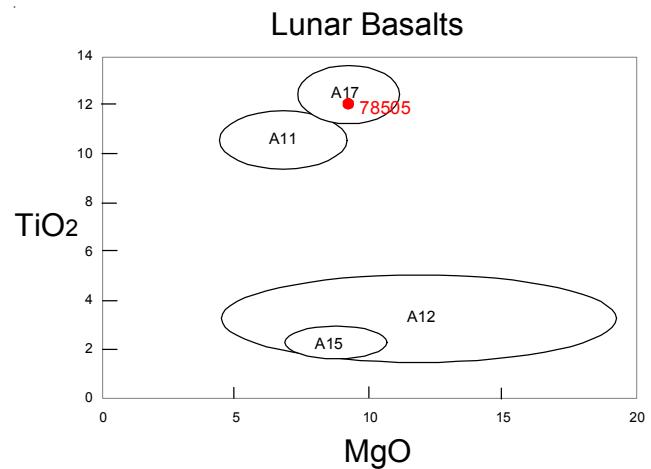


Figure 10: Composition of 78505 compared with other Apollo basalts.

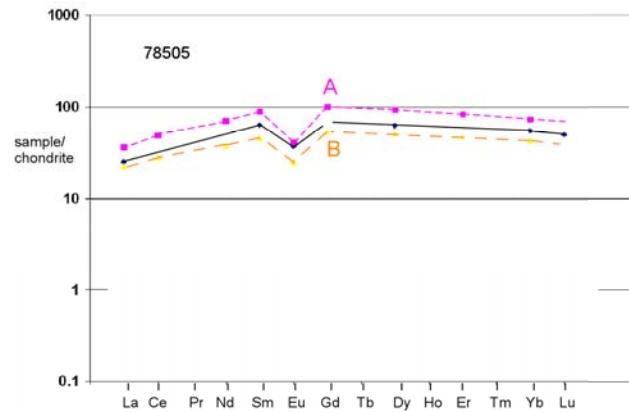


Figure 11: Normalized rare-earth-element diagram for 78505 compared with A and B types of Apollo 17 basalt.

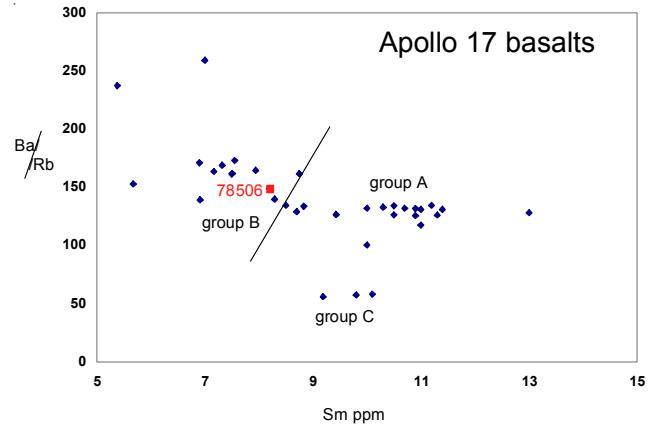


Figure 12: Trace element ratios for 78506 compared with that of other Apollo 17 basalts.

**Table 1. Chemical composition of 78505.**

	78505		78506		78507		78509
reference	Warner75	Keith 74	Rhodes76	Nyquist76	Warner79	Warner79	
SiO <sub>2</sub> %			38.55	(d)			
TiO <sub>2</sub>	12	(a)	12.93	(d)	11.9	12.3	(a)
Al <sub>2</sub> O <sub>3</sub>	10.6	(a)	8.99	(d)	8.8	9.2	(a)
FeO	18.6	(a)	19.36	(d)	18	19	(a)
MnO	0.227	(a)	0.27	(d)	0.222	0.252	(a)
MgO	9.5	(a)	9.59	(d)	10	8	(a)
CaO	9.9	(a)	9.94	(d)	9.7	10.9	(a)
Na <sub>2</sub> O	0.458	(a)	0.39	(d)	0.407	0.414	(a)
K <sub>2</sub> O	0.07	(a)	0.061	(b)	0.05	0.037	0.04
P <sub>2</sub> O <sub>5</sub>				0.02	(d)		
S %				0.16	(d)		
<i>sum</i>							
Sc ppm	74	(a)	73	(b)	79	89	(a)
V	108	(a)			130	101	(a)
Cr	2983	(a)	3489	(d)	3856	2657	(a)
Co	18.7	(a)	17.6	(b)	21	22	(a)
Ni							
Cu							
Zn							
Ga							
Ge ppb							
As							
Se							
Rb				0.44	(c)		
Sr				175	(c)		
Y							
Zr							
Nb							
Mo							
Ru							
Rh							
Pd ppb							
Ag ppb							
Cd ppb							
In ppb							
Sn ppb							
Sb ppb							
Te ppb							
Cs ppm							
Ba			66	(c)			
La	5.9	(a)	5.1	(c)	3.4	3.9	(a)
Ce			17.8	(c)	13	14	(a)
Pr							
Nd			19.6	(c)	16	16	(a)
Sm	9.4	(a)	8.19	(c)	6	5.8	(a)
Eu	2.1	(a)	1.85	(c)	1.59	1.22	(a)
Gd			12.9	(c)			
Tb					1.4	1.3	(a)
Dy	15	(a)	14.9	(c)	10	9	(a)
Ho							
Er							
Tm							
Yb	8.9	(a)	7.99	(c)	5.8	5.5	(a)
Lu	1.2	(a)	1.11	(b)	0.86	0.8	(a)
Hf			8.2	(b)	5.5	5.1	(a)
Ta					1.3	1.3	(a)
W ppb							
Re ppb							
Os ppb							
Ir ppb							
Pt ppb							
Au ppb							
Th ppm			0.39	(b)			
U ppm			0.135	(b)			

technique: (a) INAA, (b) radiation count. (c) IDMS, (d) XRF

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